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CENTER FOR

ASTROPHYSICS

HARVARD & SMITHSONIAN



**LIVERPOOL
JOHN MOORES
UNIVERSITY**

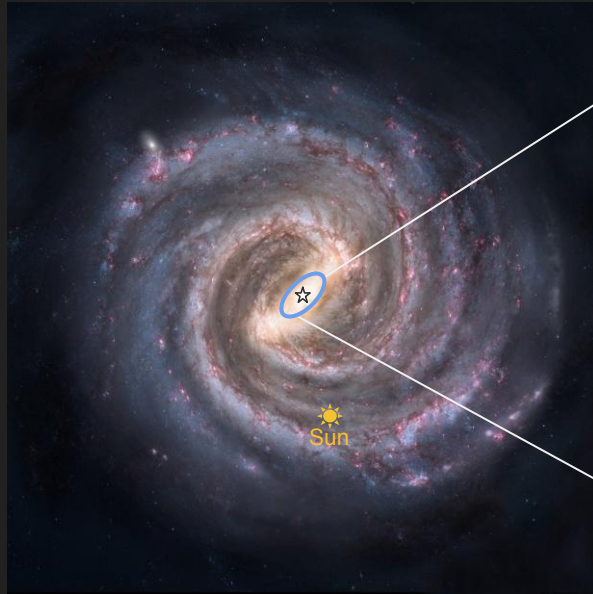
ACES: Resolving the super to subsonic gas transition in the Galactic center for the first time



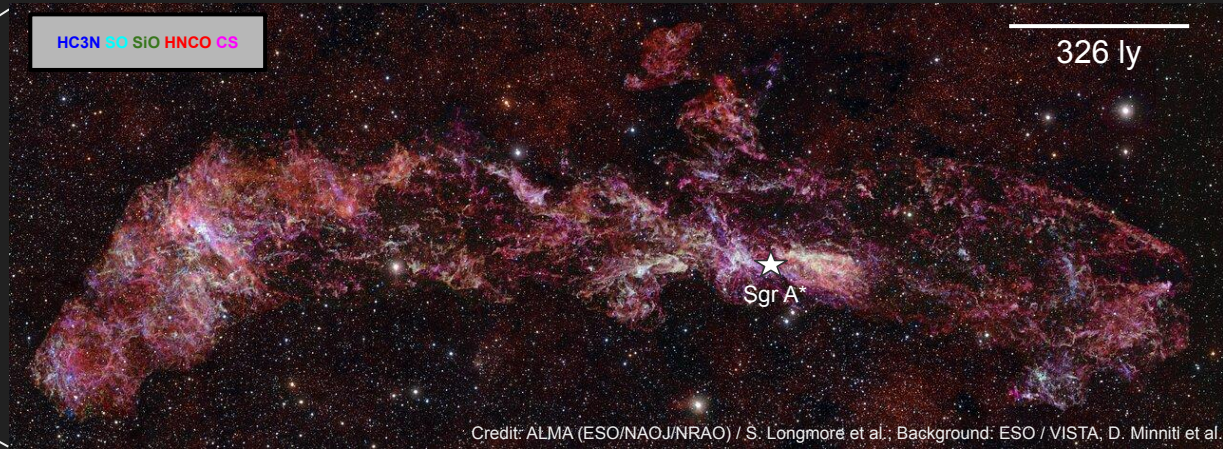
Rojita P. Buddhacharya, Jonathan D. Henshaw, Steven N. Longmore, Daniel L. Walker,
Rebecca J. Houghton, Qizhou Zhang, Ashley T. Barnes, Adam Ginsburg and
the ACES Team

Galactic Center: A surprisingly inefficient star factory!

A window into how our own Solar system formed



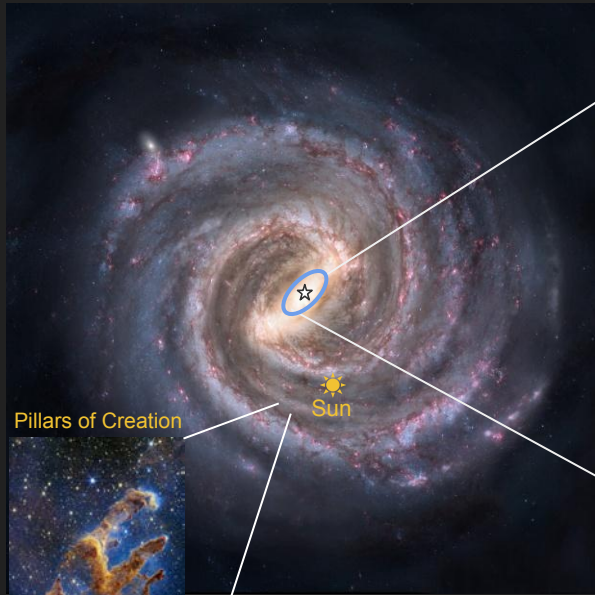
ACES: The Milky Way's Central Molecular Zone (Gas + Stars)



Credit: ALMA (ESO/NAOJ/NRAO) / S. Longmore et al.; Background: ESO / VISTA, D. Minniti et al.

Galactic Center: A surprisingly inefficient star factory!

A window into how our own Solar system formed



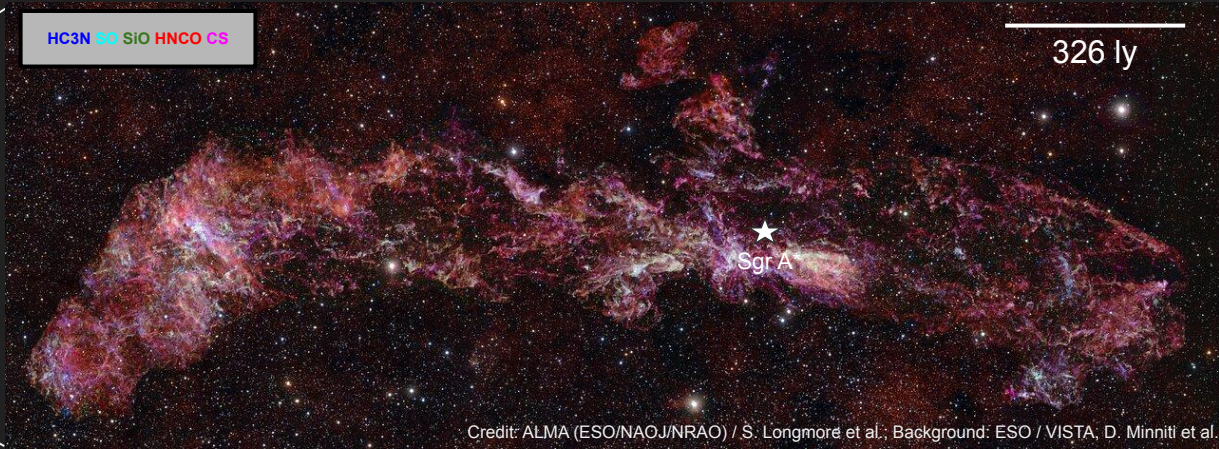
Pillars of Creation

Sun

Star formation rate \propto gas density

Credit: JWST

ACES: The Milky Way's Central Molecular Zone (Gas + Stars)

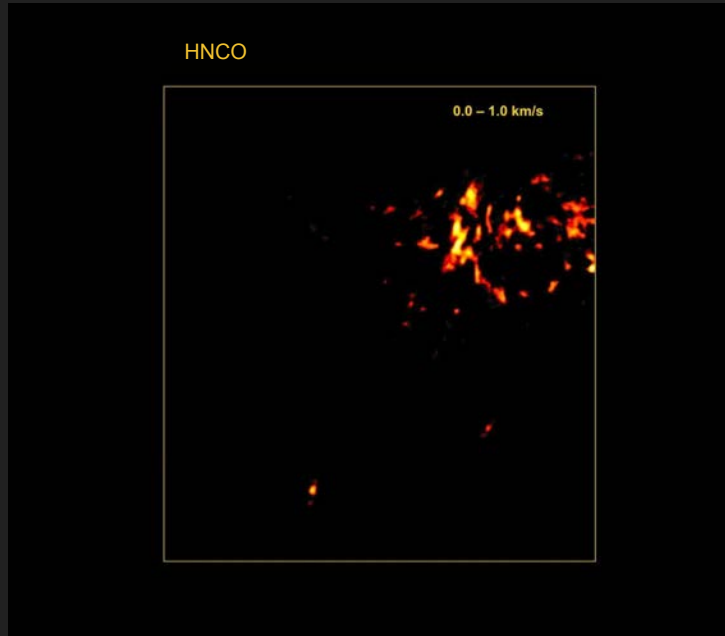


- ★ Why is star formation so inefficient in the Galactic Center?
- ★ CMZ: the closest lab for how stars like our Sun were born!

The suspect: Turbulence?

Turbulence tears clouds apart ...

Swirling ball of gas in a CMZ cloud



Islands of calm in a turbulent sea on Earth

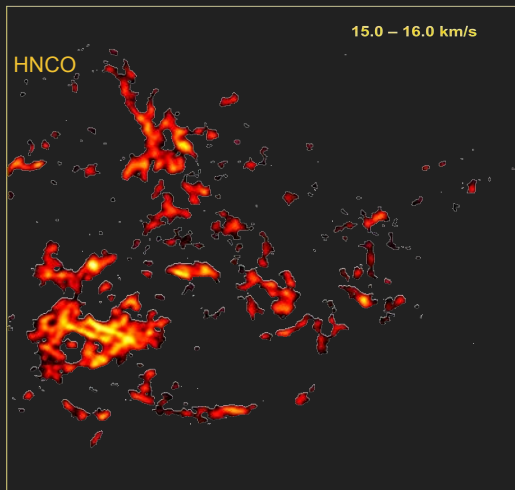


Supersonic gas → subsonic gas → Early stage of star formation

The suspect: Turbulence?

Turbulence brings gas together ...

Swirling ball of gas in a CMZ cloud



Islands of calm in a turbulent sea on Earth



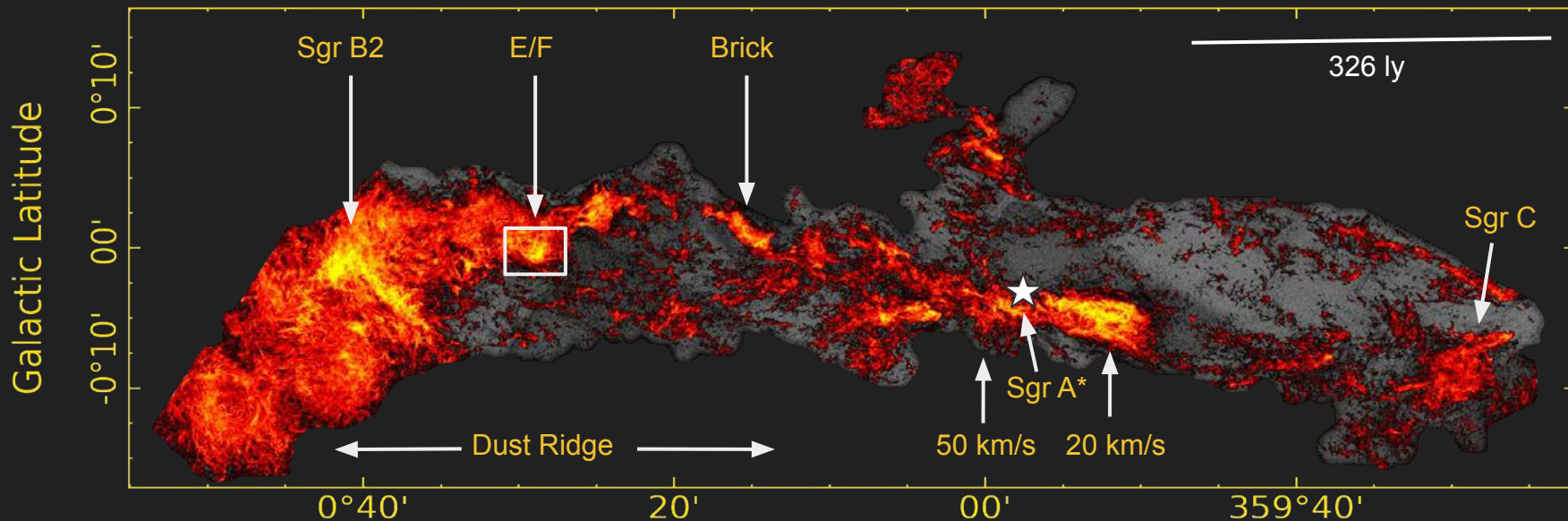
- ★ Islands of calm in a turbulent sea = Zen state of gas in ISM
- ★ First direct observation in Perseus molecular cloud (Pineda et al. 2010)
- ★ Widespread subsonic gas motions in high mass star forming clouds (Henshaw et al. 2014, Sokolov et al. 2018, Li et al. 2020)



Galactic Center: Can we find Islands of calm?

ALMA Exploration CMZ Survey (ACES) does the magic!

HNCO (4-3) Peak Intensity Map of the CMZ



- ★ Spectral resolution = 0.25 km/s
- ★ Spatial resolution = 0.3 ly

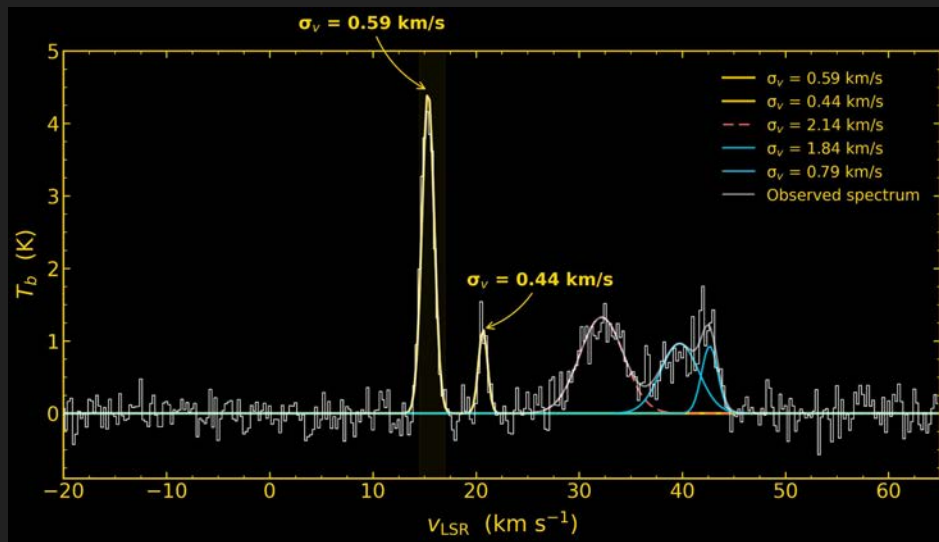
Galactic Longitude

Longmore et al. 2026, Walker et al. 2026

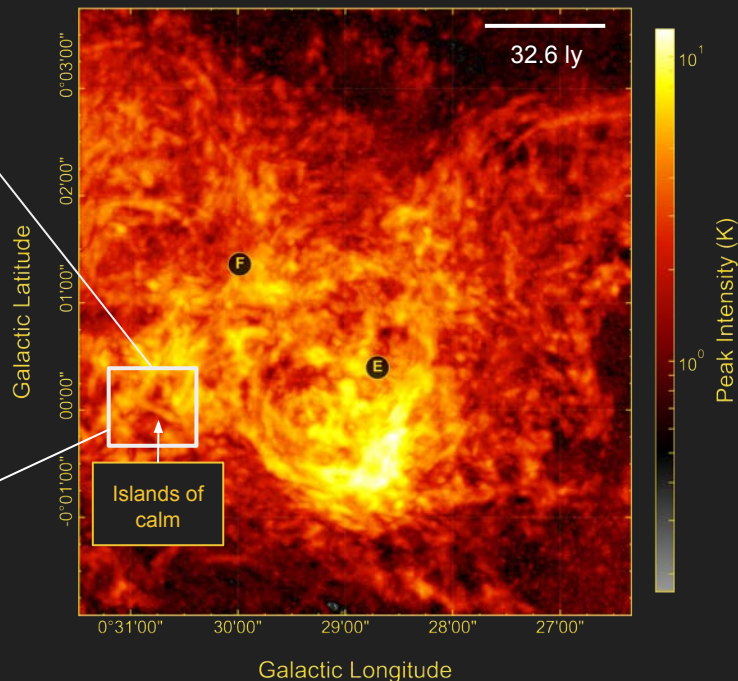
Galactic Center: Where is Islands of calm?

..... in the Dust Ridge

An Example Spectrum



HNCO (4-3) Peak Intensity map of Cloud E/F



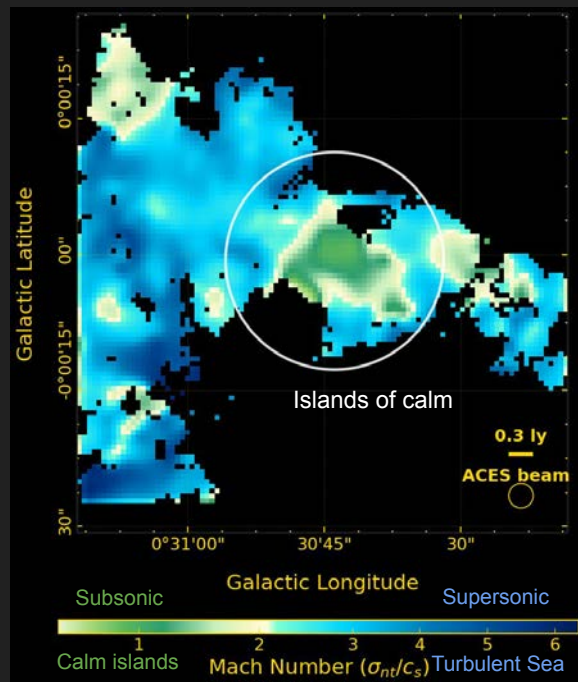
- ★ $\sim 0.3 \text{ ly wide, } \sim 2 \text{ ly long}$
- ★ Large $\sigma_v =$ stormy, turbulent gas
- ★ Small $\sigma_v =$ calm, coherent gas

Galactic Center: Islands of calm in a turbulent sea

Resolving the super to subsonic gas transition !

- ★ Calm (**subsonic**) gas surrounded by turbulent (**supersonic**) gas.

Islands of calm in the CMZ

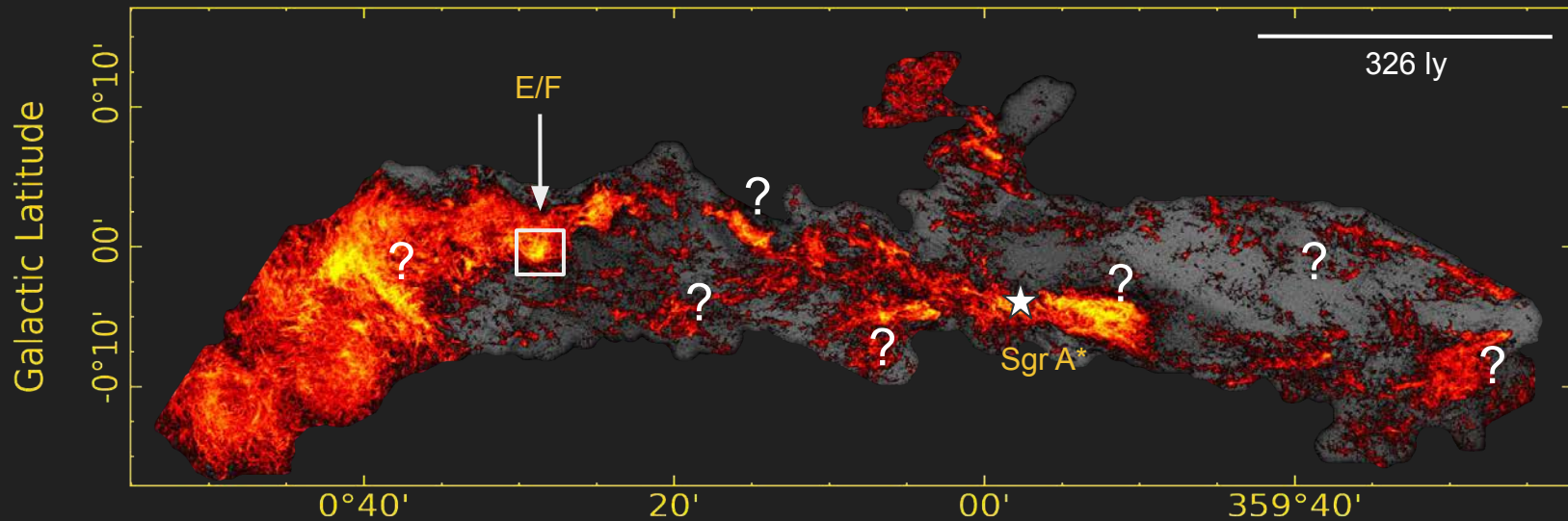


Islands of calm on Earth



Galactic Center: Where will next generations of stars will form?

HNCO (4-3) Peak Intensity Map of the CMZ



- ★ Each pixel is a clue. Every Island of calm is a potential stellar birthsites.
- ★ ML tools to search Islands of calm

Galactic Longitude

Longmore et al. 2026, Walker et al. 2026,
Buddhacharya et al. in prep



Dr. Qizhou Zhang

Center for Astrophysics | Harvard & Smithsonian



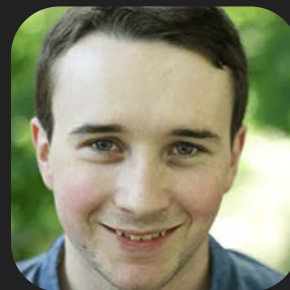
Prof. Steve Longmore

Astrophysics Research Institute,
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Max Planck Institute for Astronomy



Dr. Daniel Walker

UK ALMA Regional Centre Node,
Jodrell Bank Centre for Astrophysics



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Max Planck Institute for Astronomy



Dr. Ashley Barnes

European Southern Observatory



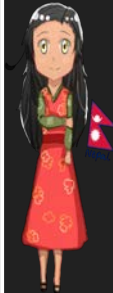
Prof. Adam Ginsburg

Department of Astronomy, University
of Florida



The ACES Collaboration

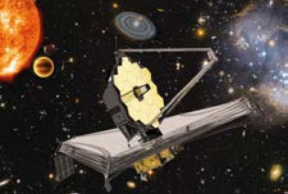
What we found and why it matters?



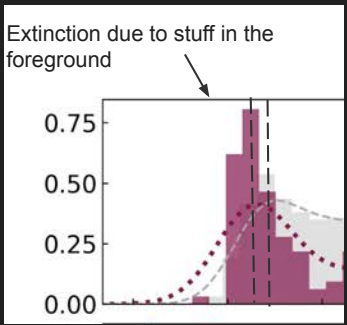
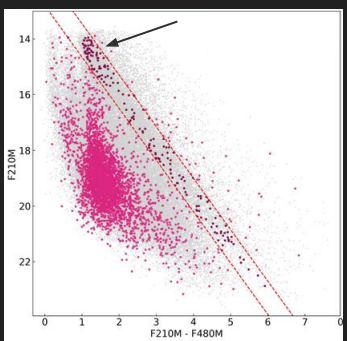
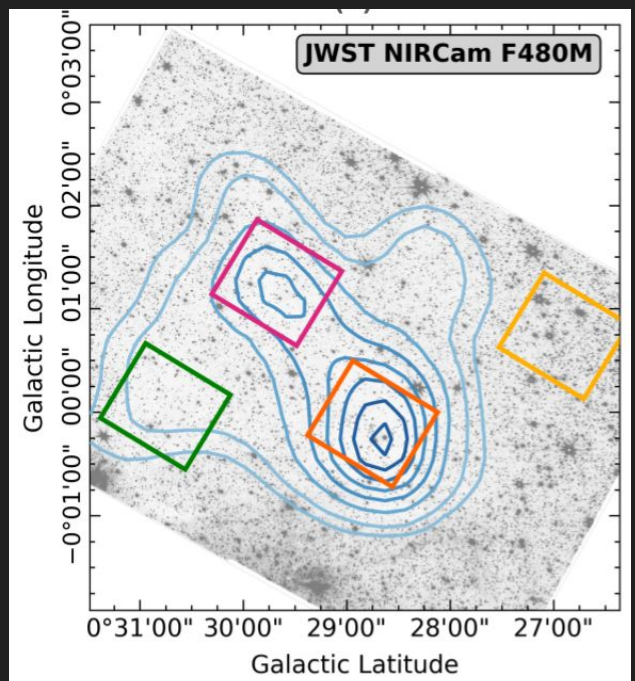
First detection of **Islands of Calm** in stormy **Galactic Center** where gravity may begin to overcome turbulence and trigger gas collapse.

Initial conditions for star formation may be similar across diverse environments including where our Sun formed.

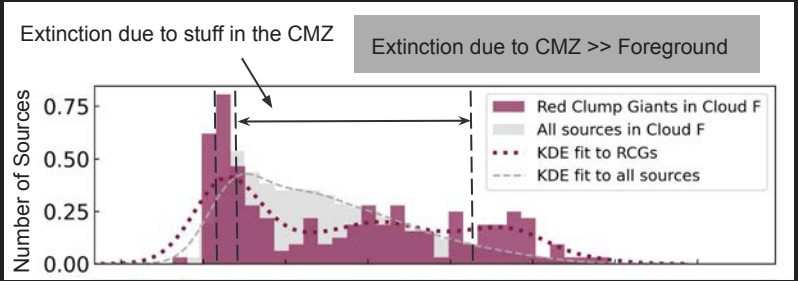
Additional Slides



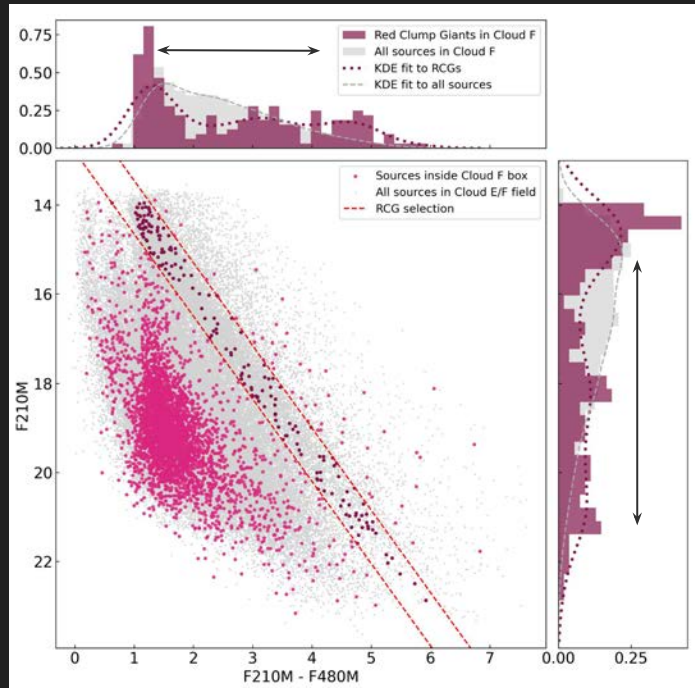
Red Clump Giants in Cloud E/F



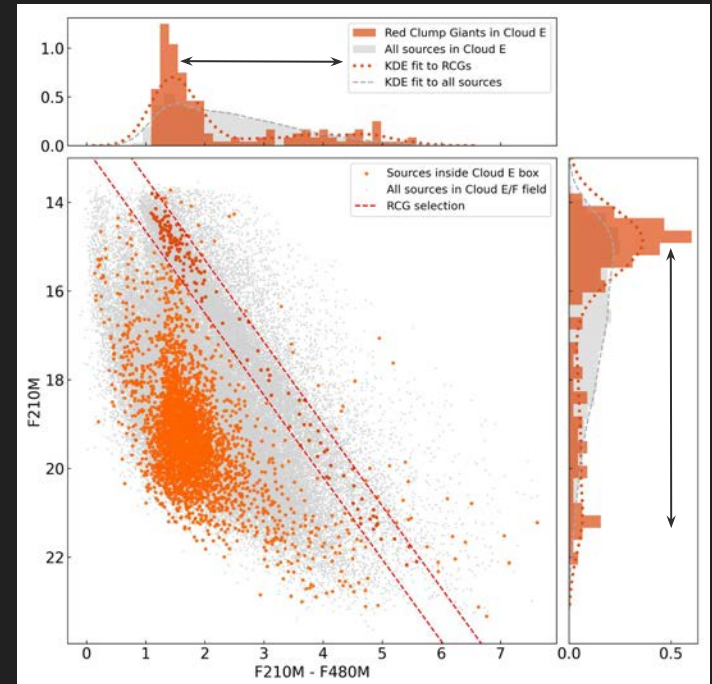
- ★ Red Clump Giants are evolved, Helium-burning stars.
- ★ Standard light bulbs in space



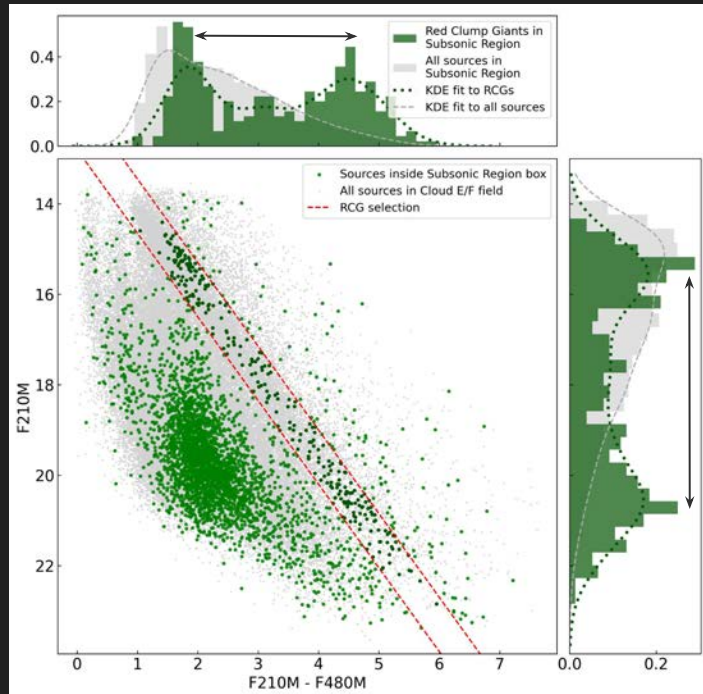
Red Clump Giants in Cloud E/F



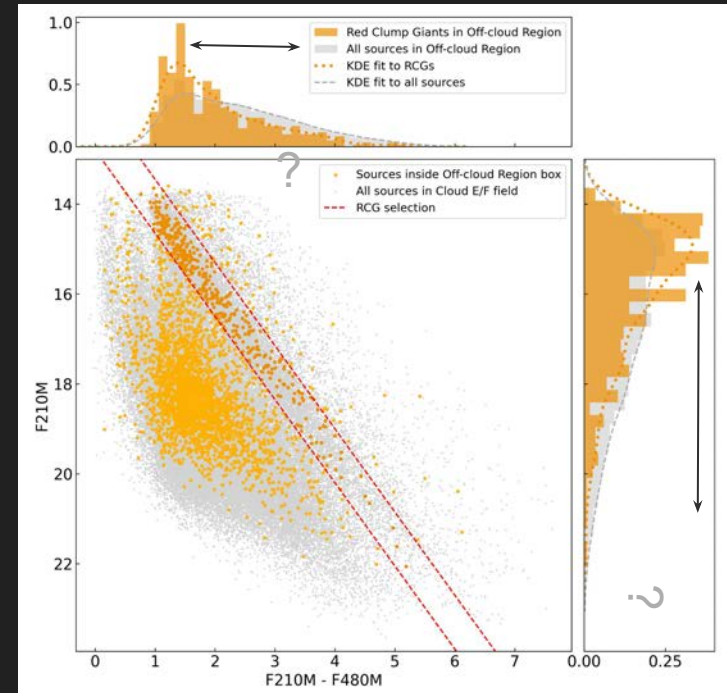
Two peaks in histograms =
Extinction due to CMZ



Red Clump Giants in Cloud E/F

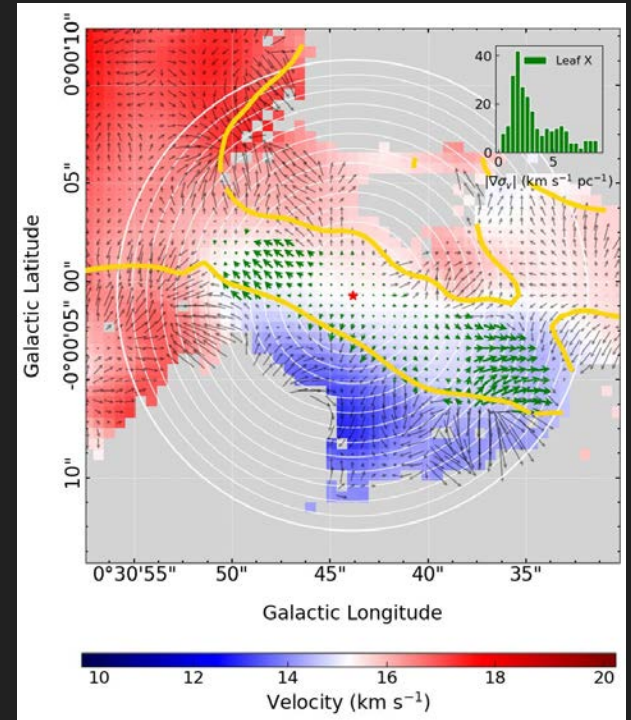


Two peaks in histograms in the SCR but not in the off-Cloud region

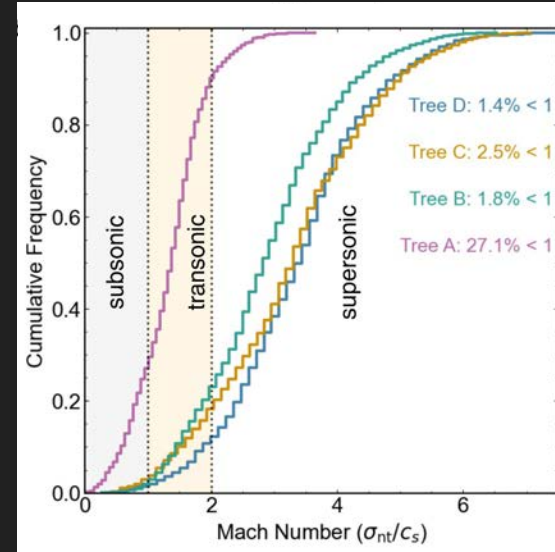
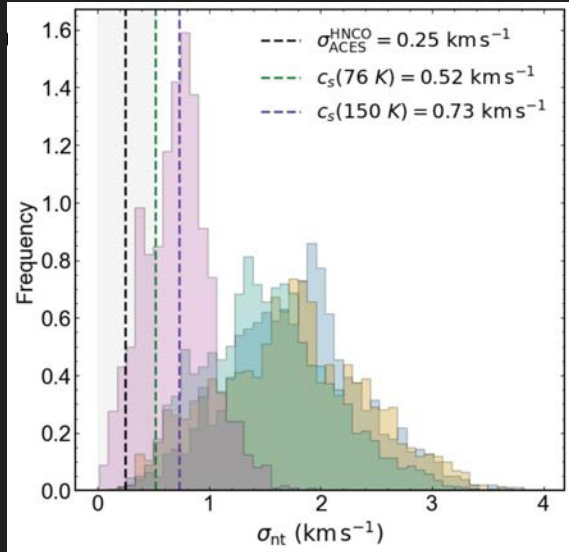


Velocity dispersion gradient: Transition to coherence

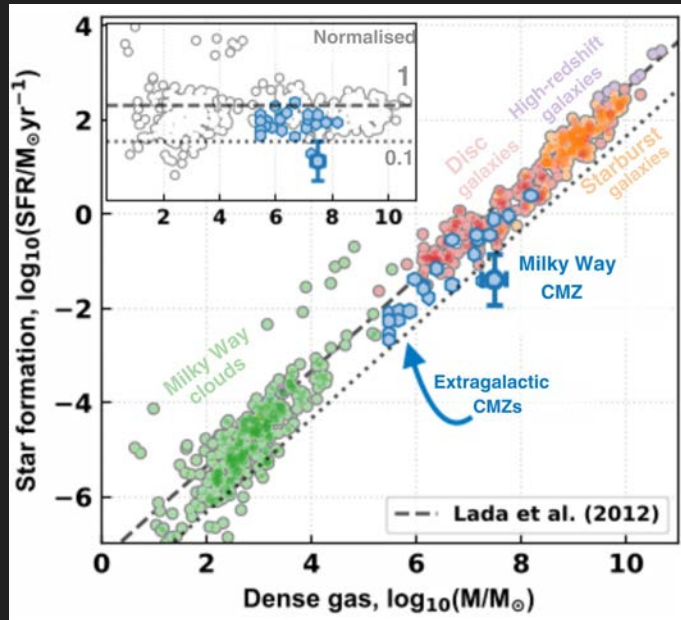
- ★ Length of vector = gradient magnitude
- ★ Anti-correlation between peak intensity and velocity dispersion



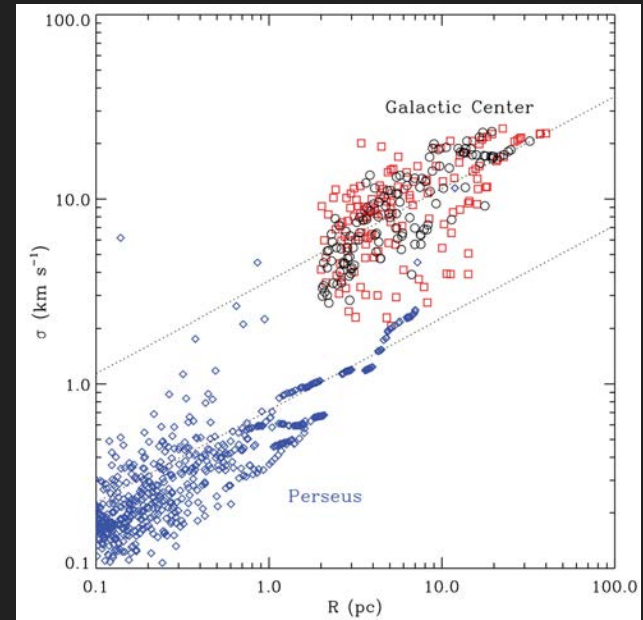
Mach number: Supersonic to subsonic gas transition



Star formation scenario

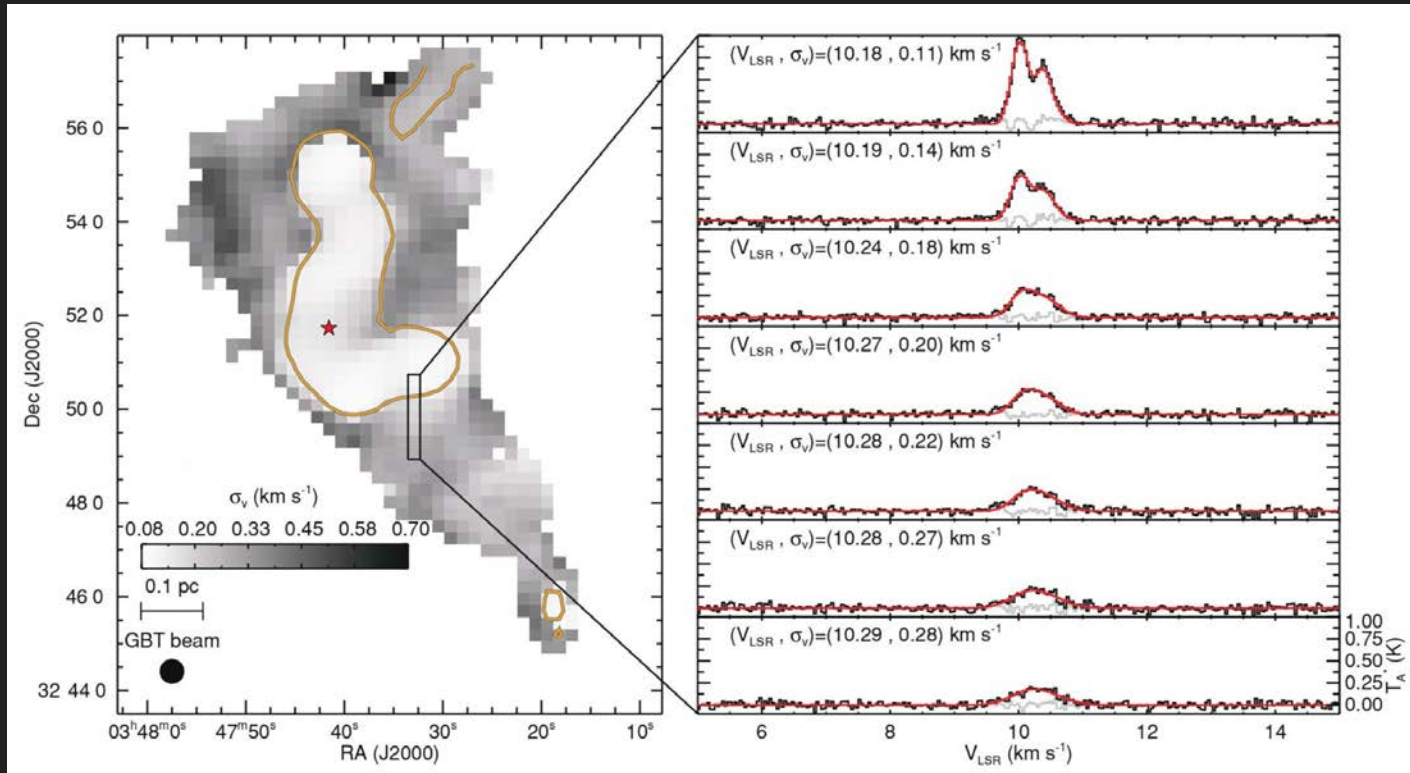


Lada et al. 2010, 2012, Longmore et al. 2013, Barnes et al. 2017, Henshaw et al. 2023



Shetty et al. 2012a

First observation in Perseus molecular cloud



Pineda et al. 2010