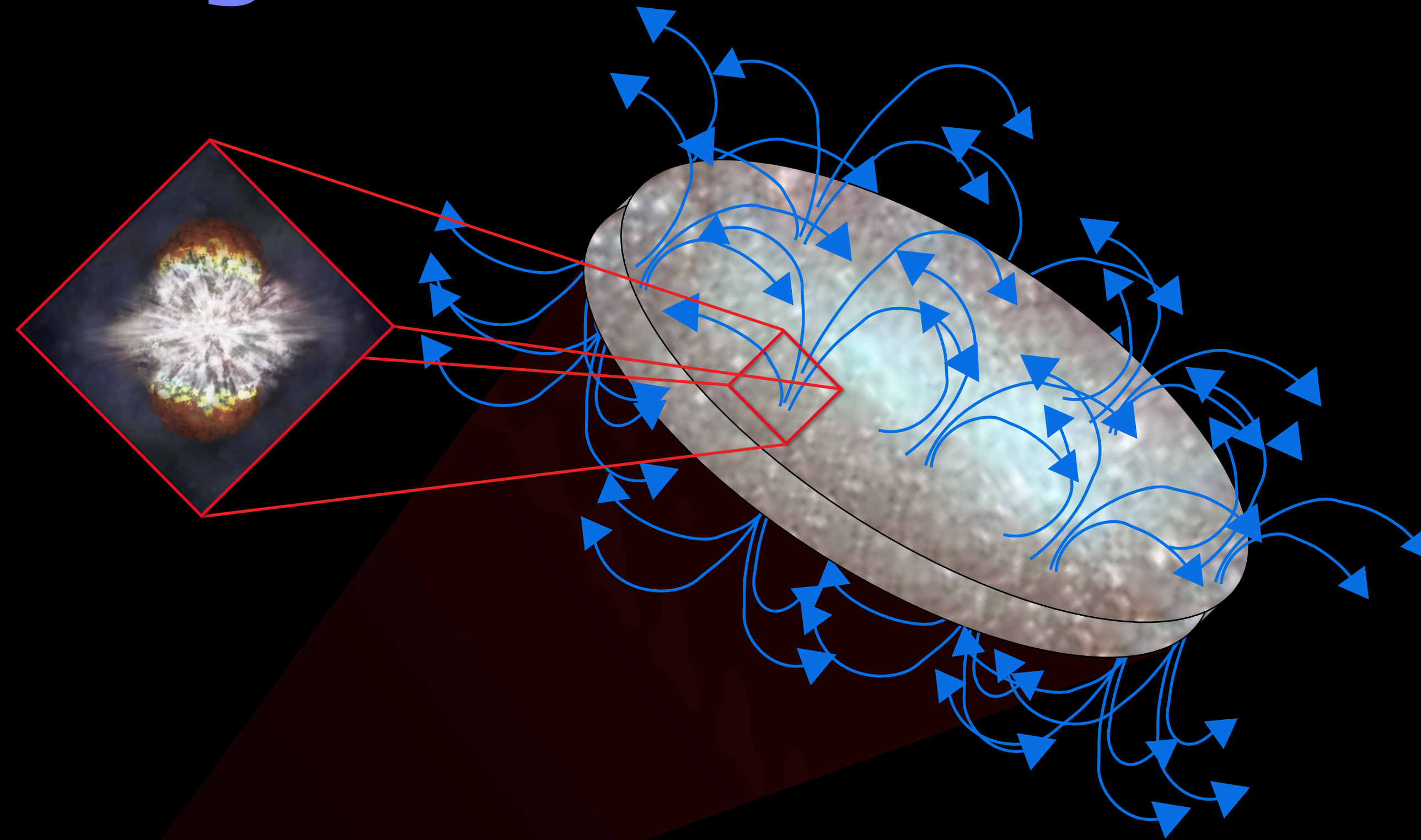


The Supernovae-Driven Winds of the Large Magellanic Cloud Galaxy



AAS Talk: 427.01
Thurs at 10 am



Wisconsin H-alpha Mapper

Dr. Kat Barger
Associate Professor
Texas Christian University

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Image credit (modified):
LMC: Eckhard Slawik
Supernove: NASA/CXC/M. Weiss

Supernovae are driving a wide spread wind

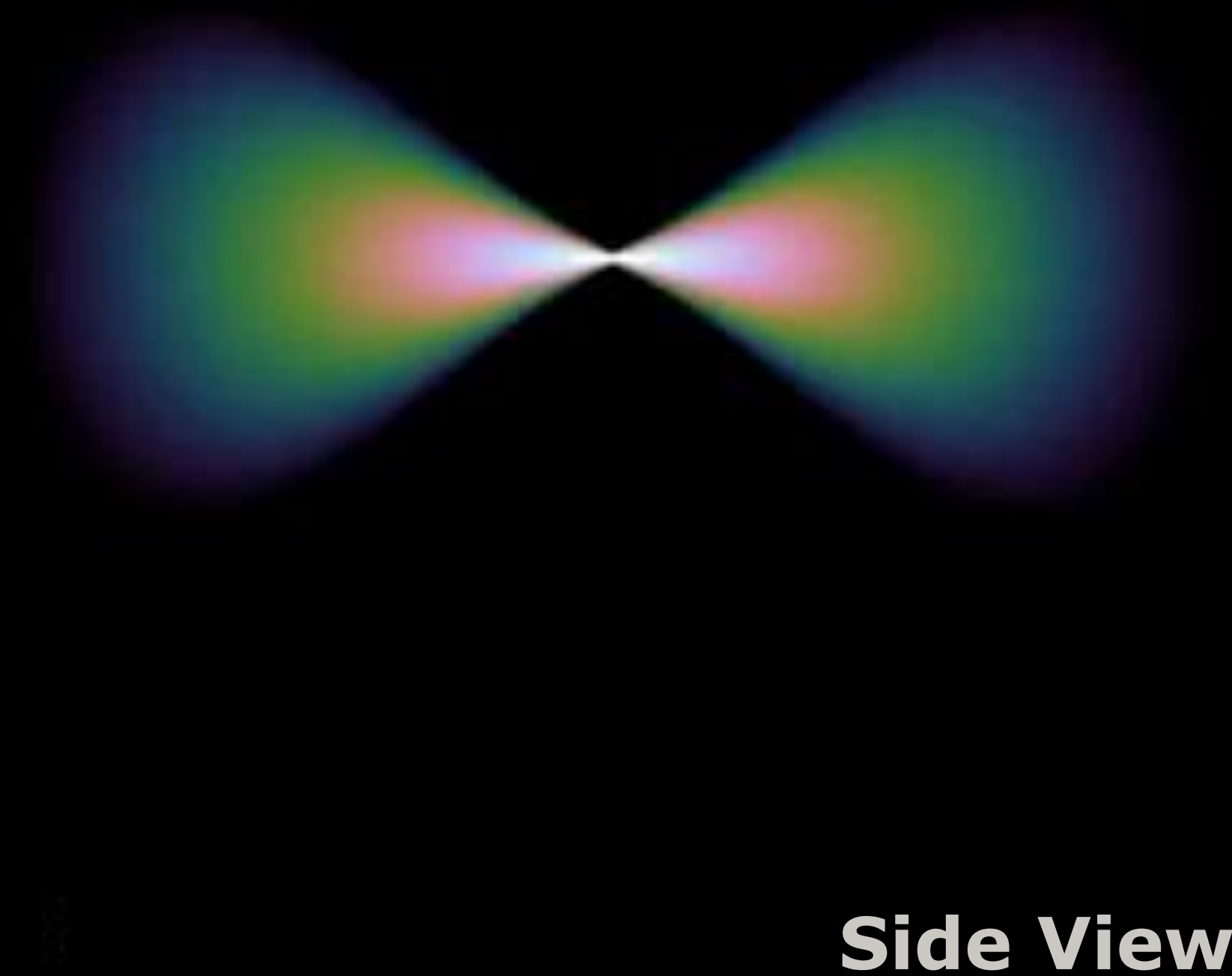
Diffuse disk emission

MCELS: Smith+2005
Winkler+2015



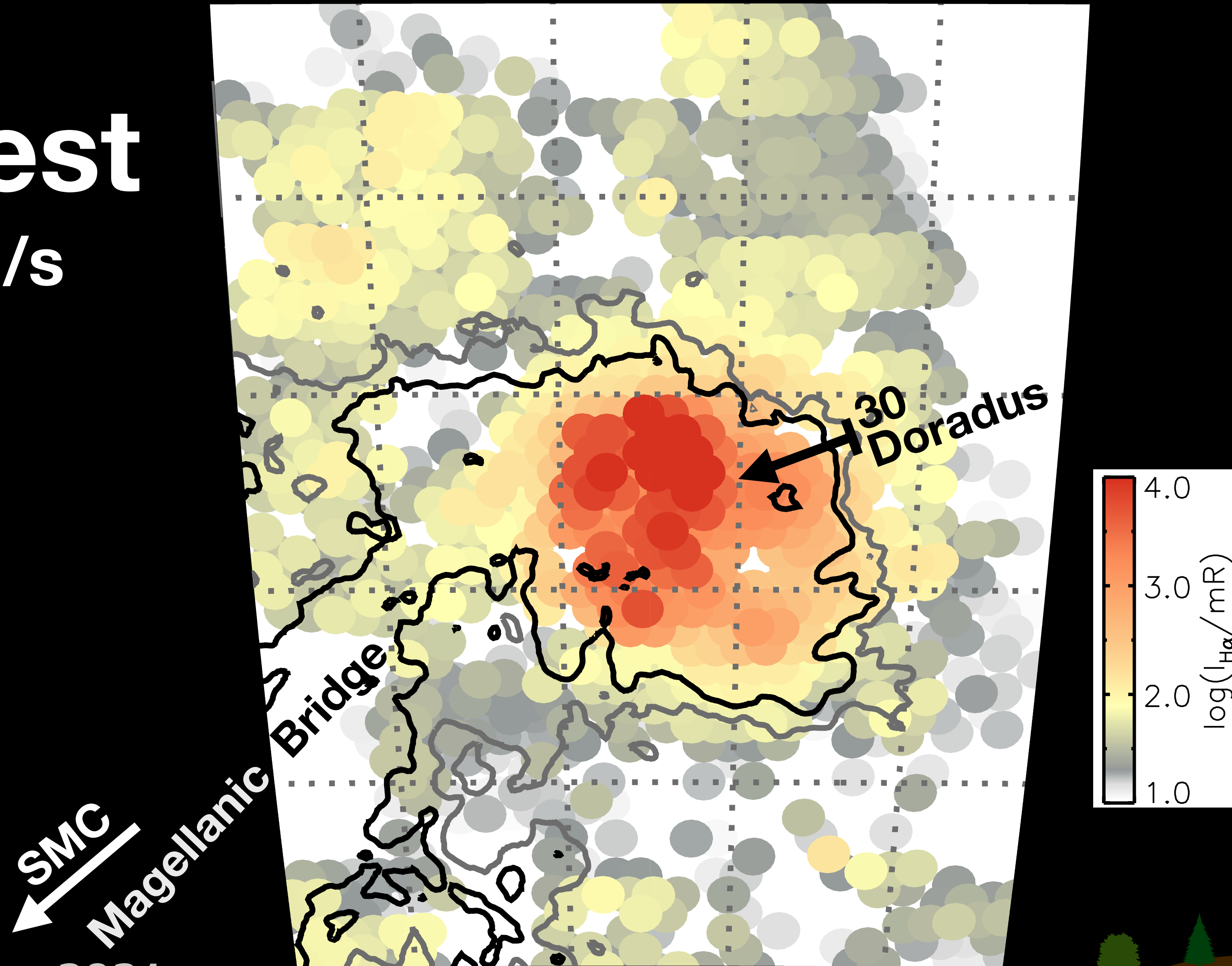
Simulations with Supernovae Explosions

Fielding+2017



Wind on near side of LMC in H α emission

Slowest
-75 km/s

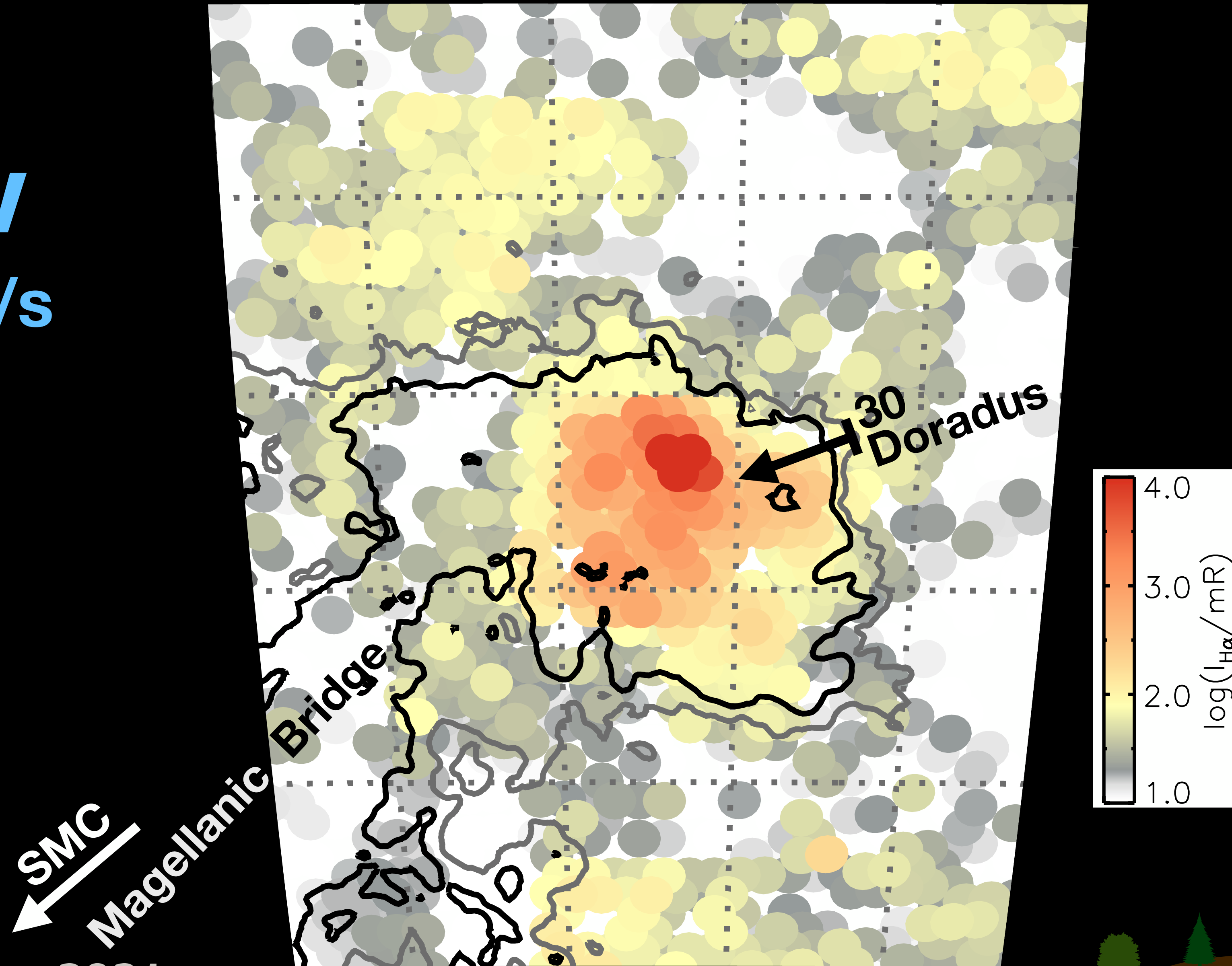


WHAM



Wind on near side of LMC in H α emission

Slow
-100 km/s

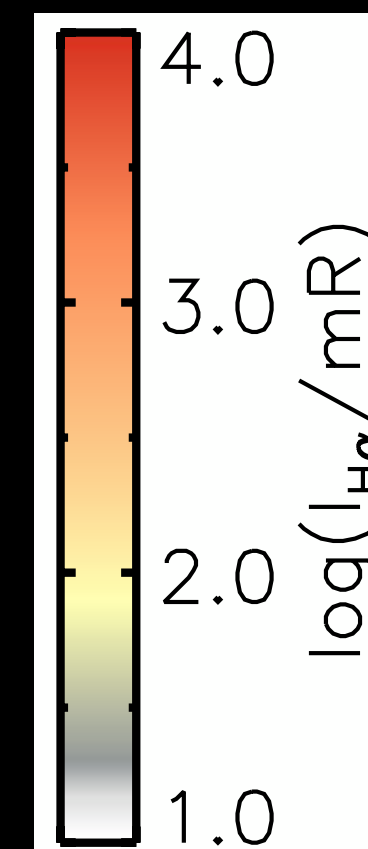
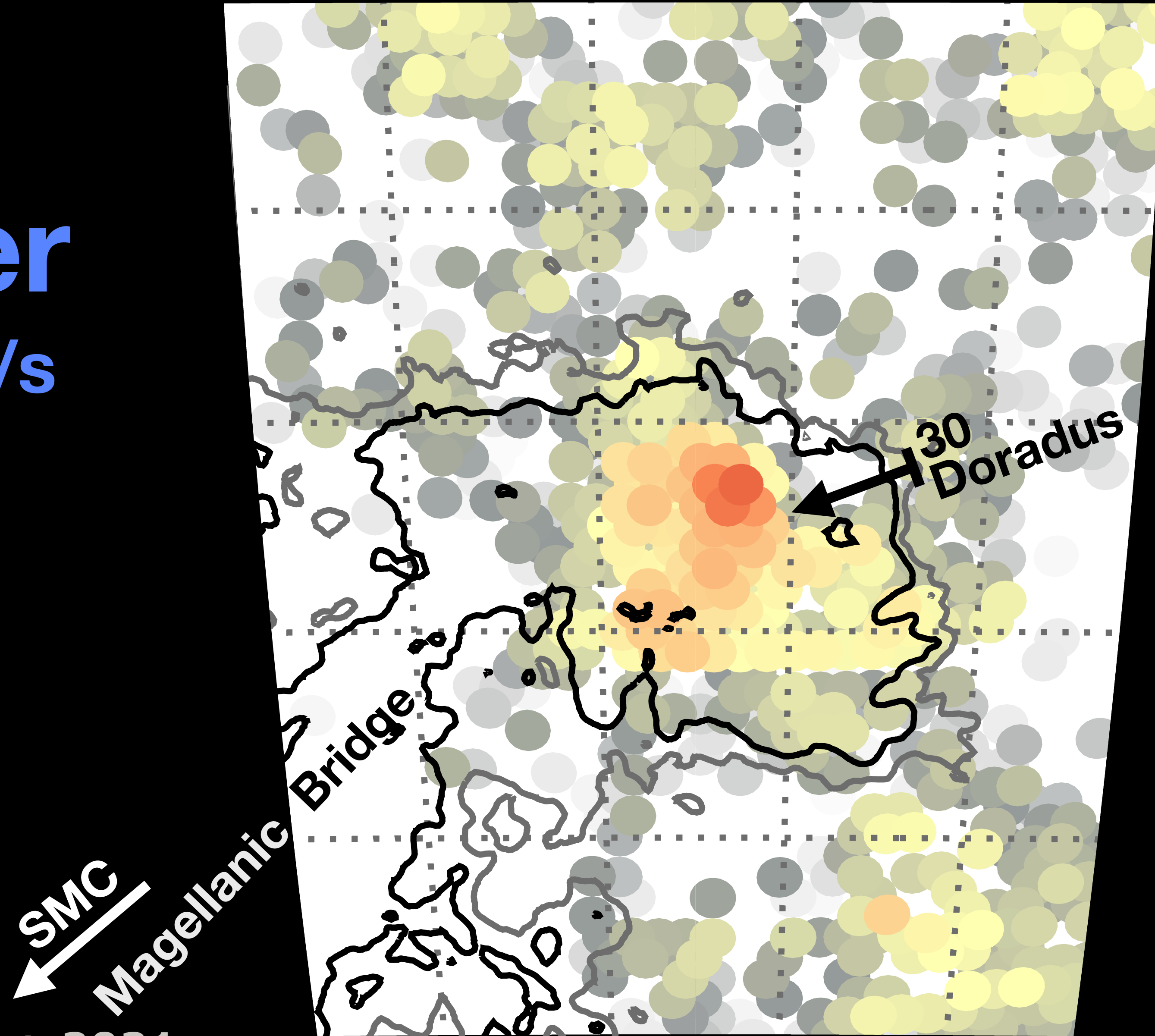


WHAM



Wind on near side of LMC in H α emission

Faster
-130 km/s

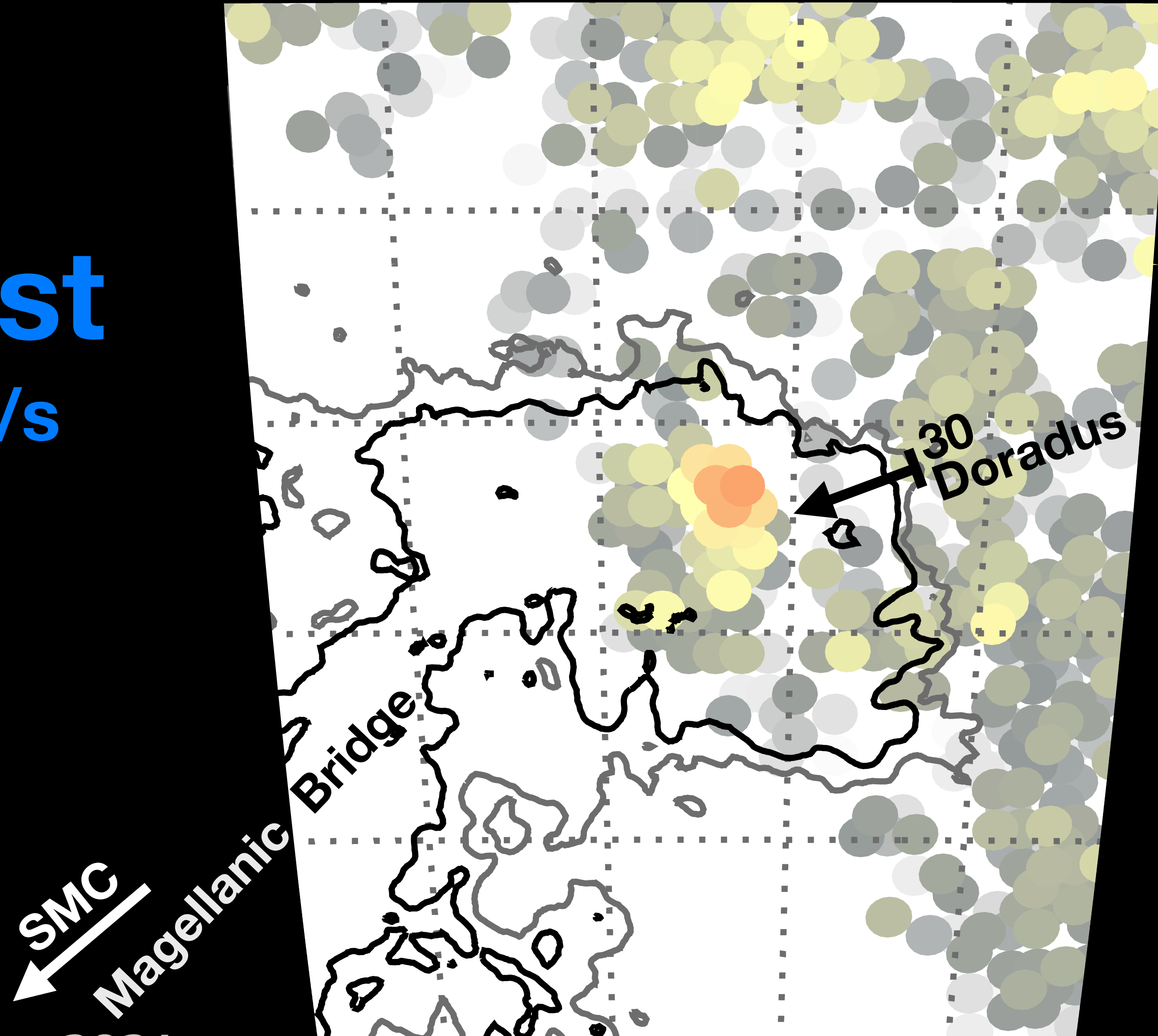


WHAM

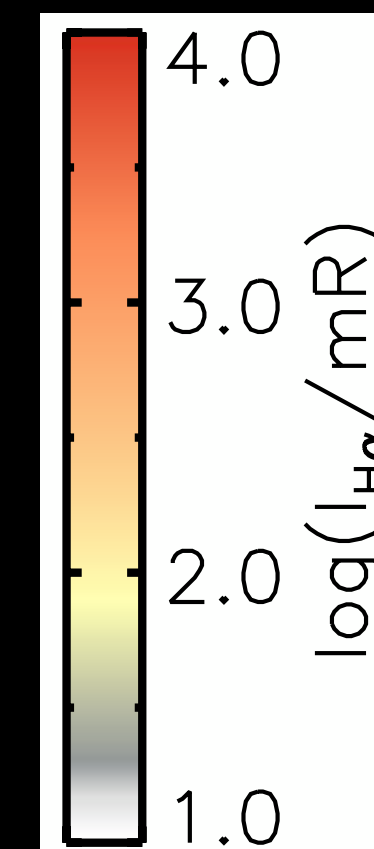


Wind on near side of LMC in H α emission

Fastest
-160 km/s



**85x10⁶ Msun of
gas is being
ejected at a rate
of 1.4 Msun/yr**



SMC

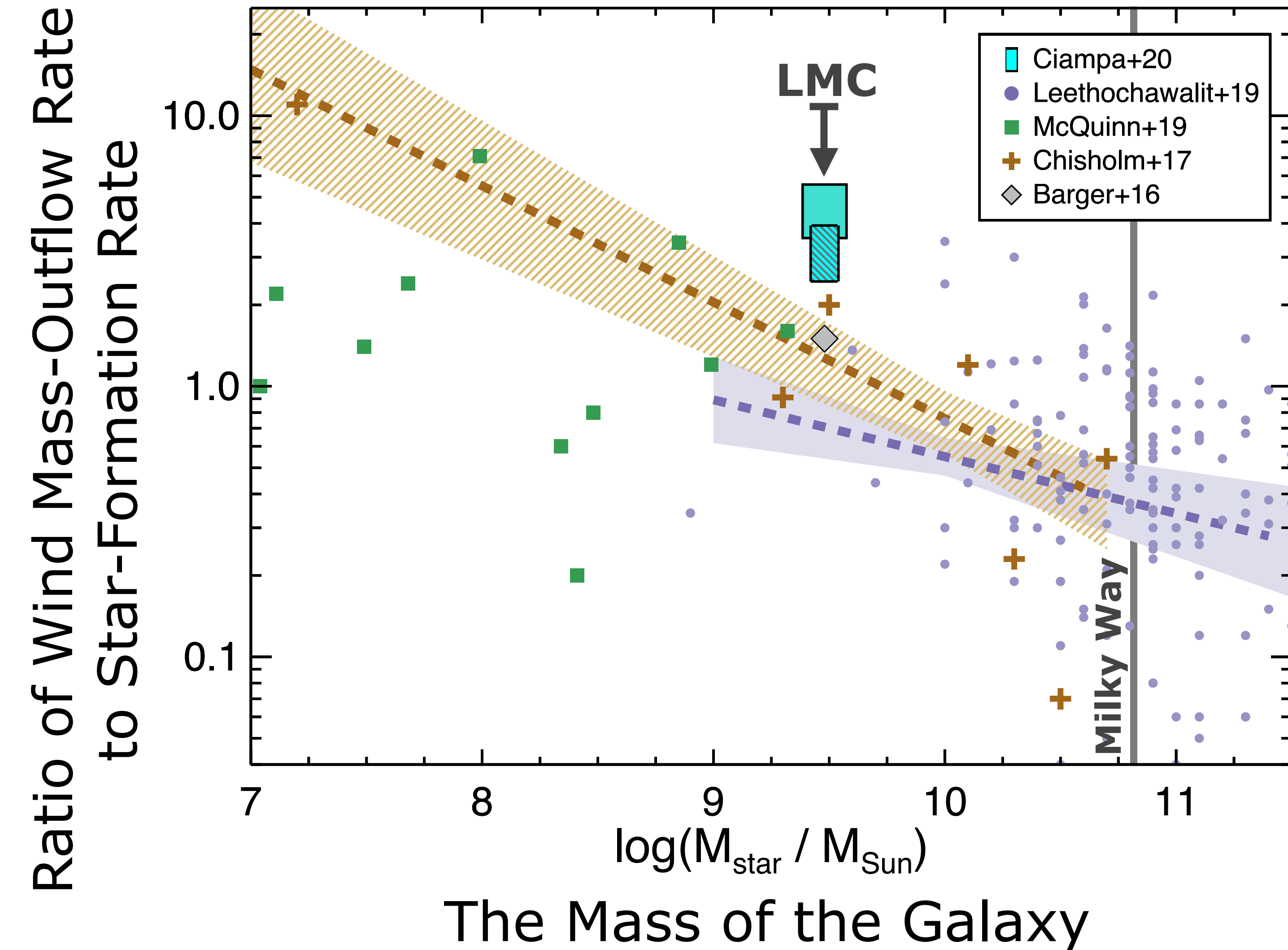
Magellanic Bridge

30 Doradus

WHAM



Mass-Loading Factor



This galaxy is losing 4.5x more gas through the wind than through forming new stars and planets!

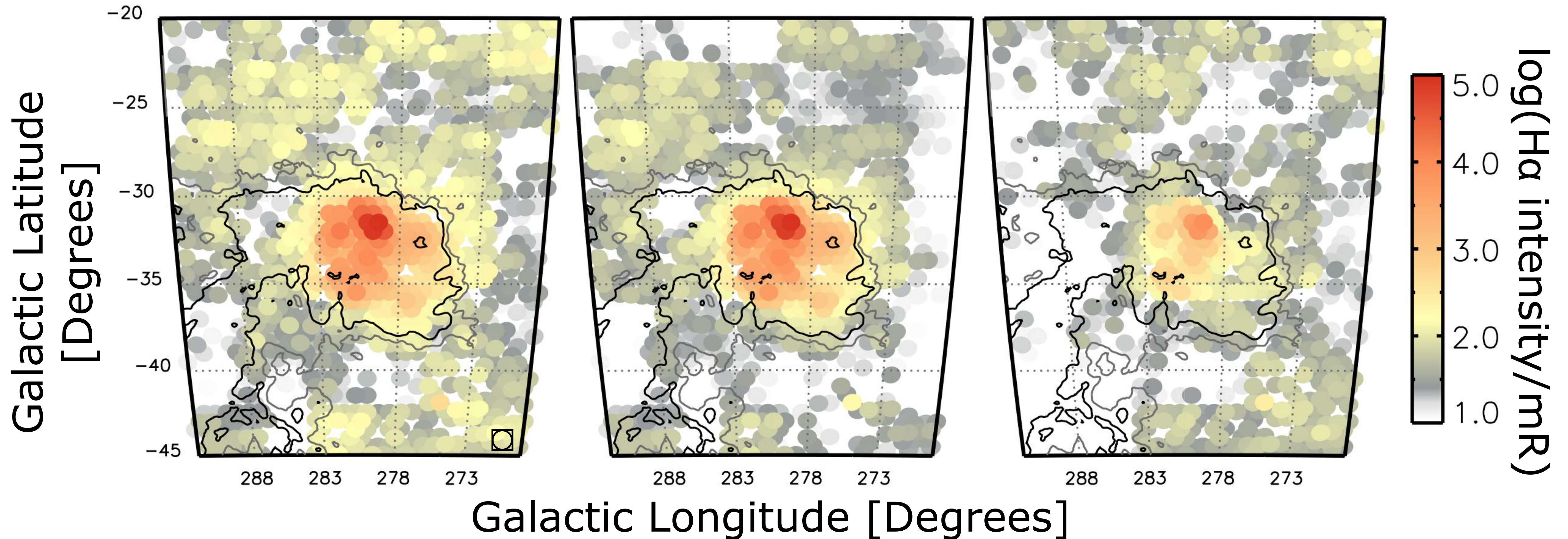
This ratio is much larger than for other galaxies of the same mass.

But our observations are 10x more sensitive & the wind is fully mapped.

All nearside
of wind

Slowest
part of wind

Fastest
part of wind

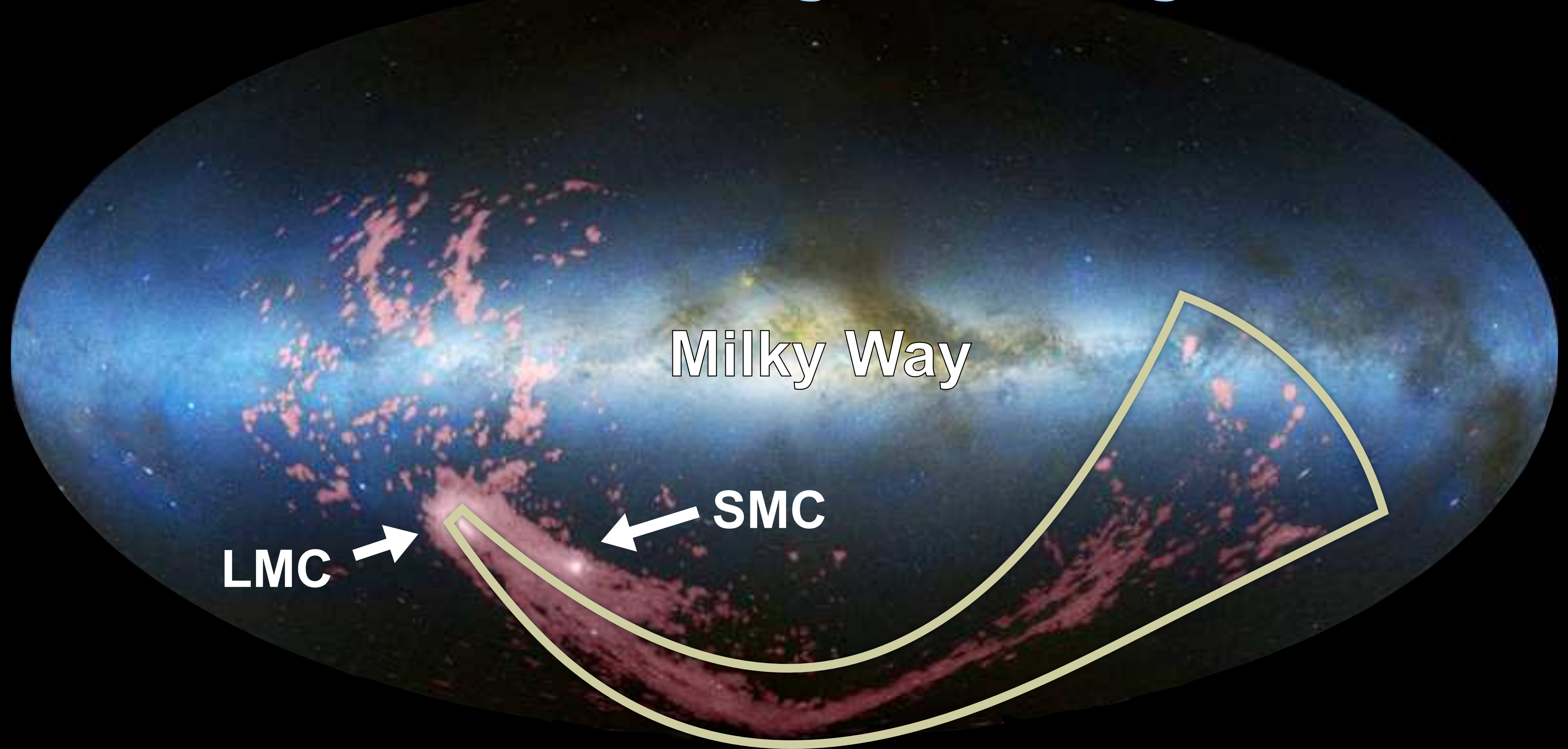


Ciampa + Barger + 2021

Will likely fall
back to the LMC

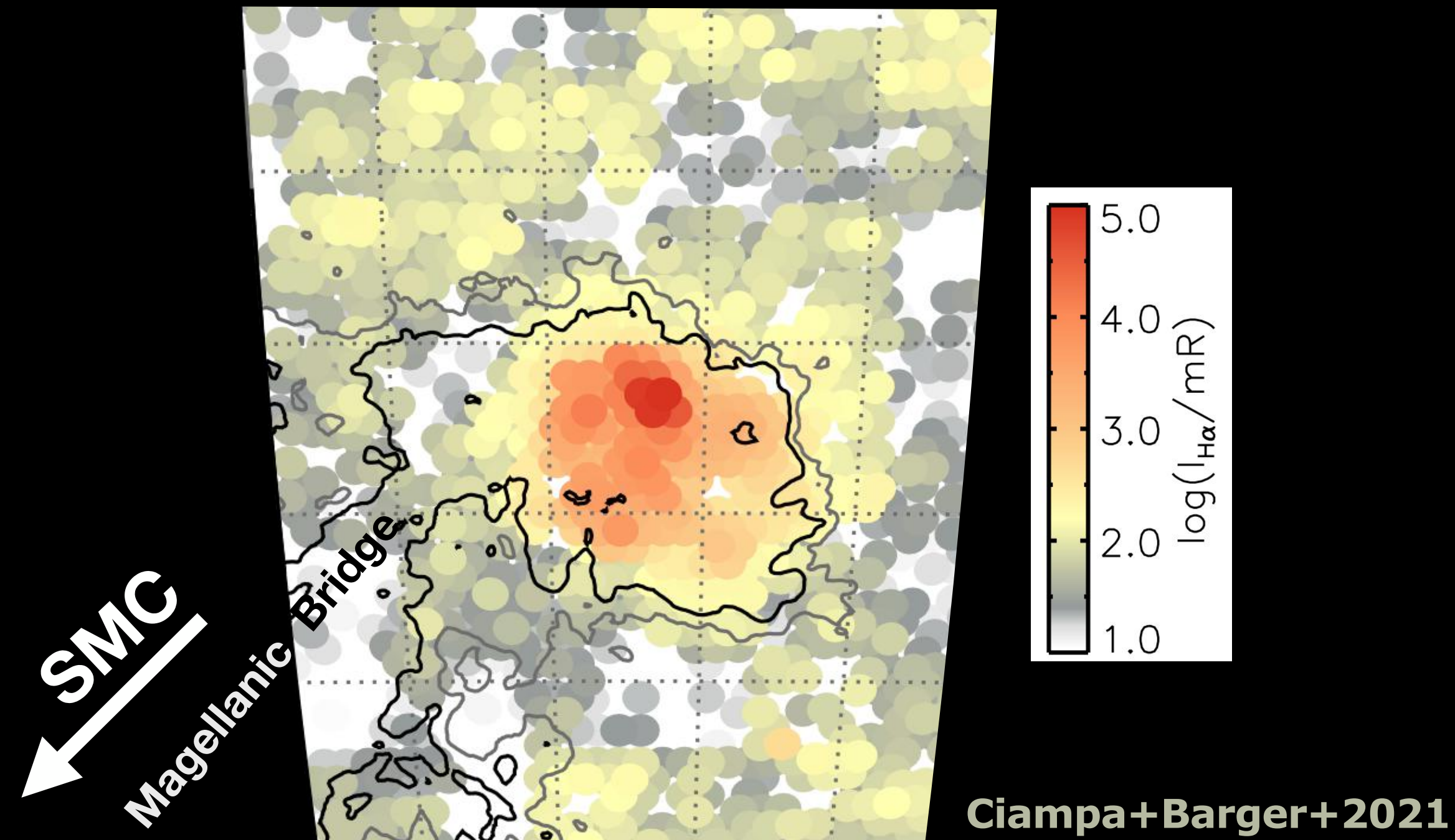
Will likely escape
and could be
captured by the
Milky way

Is the LMC's outflow feeding the trailing tidal stream?

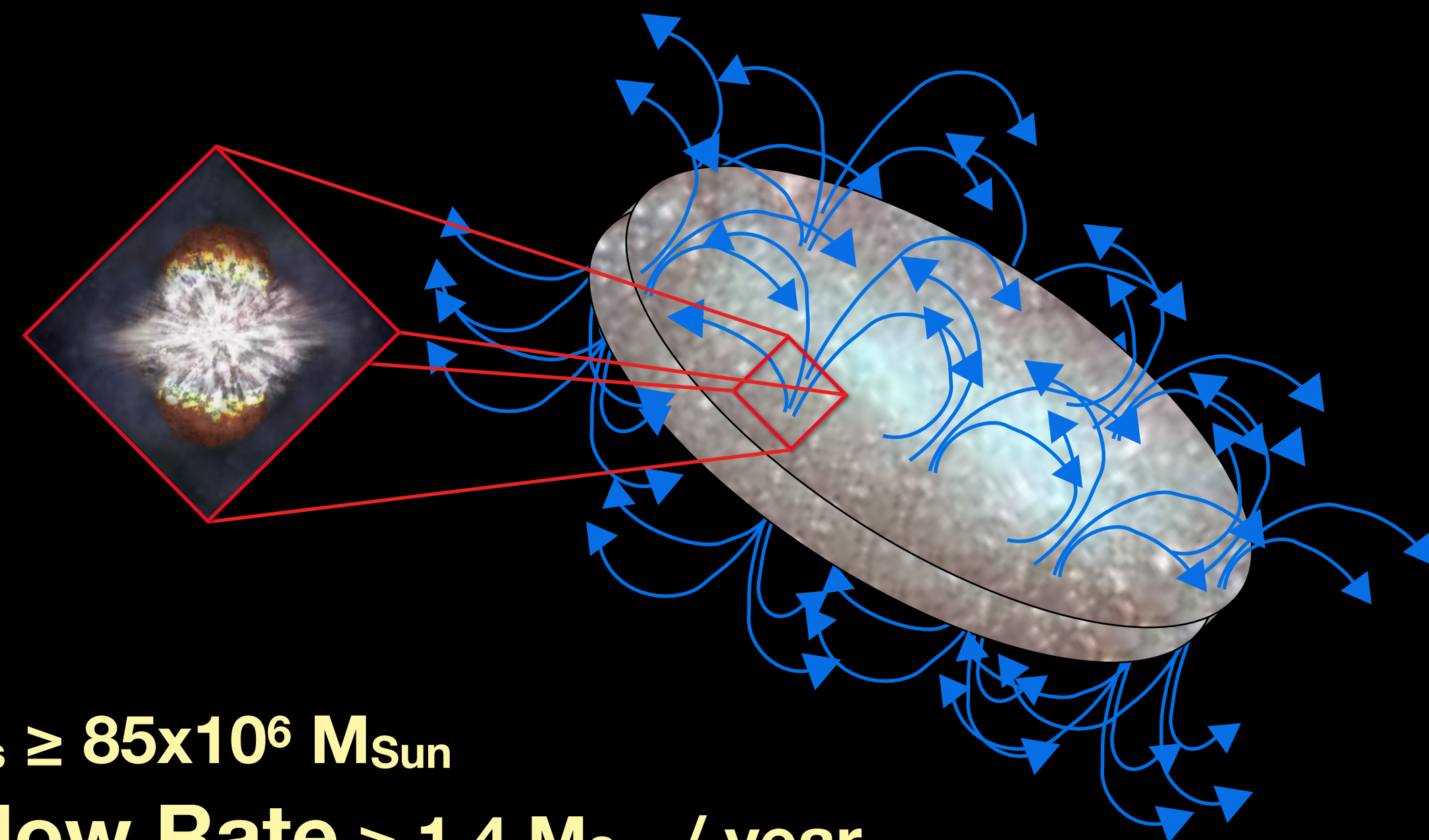


**How much of this gas will be captured by our galaxy?
And, how many new stars and planets will it form as a result?**

Near side of wind in H α emission



Stellar activity is driving a scale galactic wind out of the Large Magellanic Cloud galaxy



$M_{winds} \geq 85 \times 10^6 M_{Sun}$

Outflow Rate $\geq 1.4 M_{Sun} / \text{year}$

Barger + 2016

Ciampa + Barger+ 2021

AAS Talk: 427.01

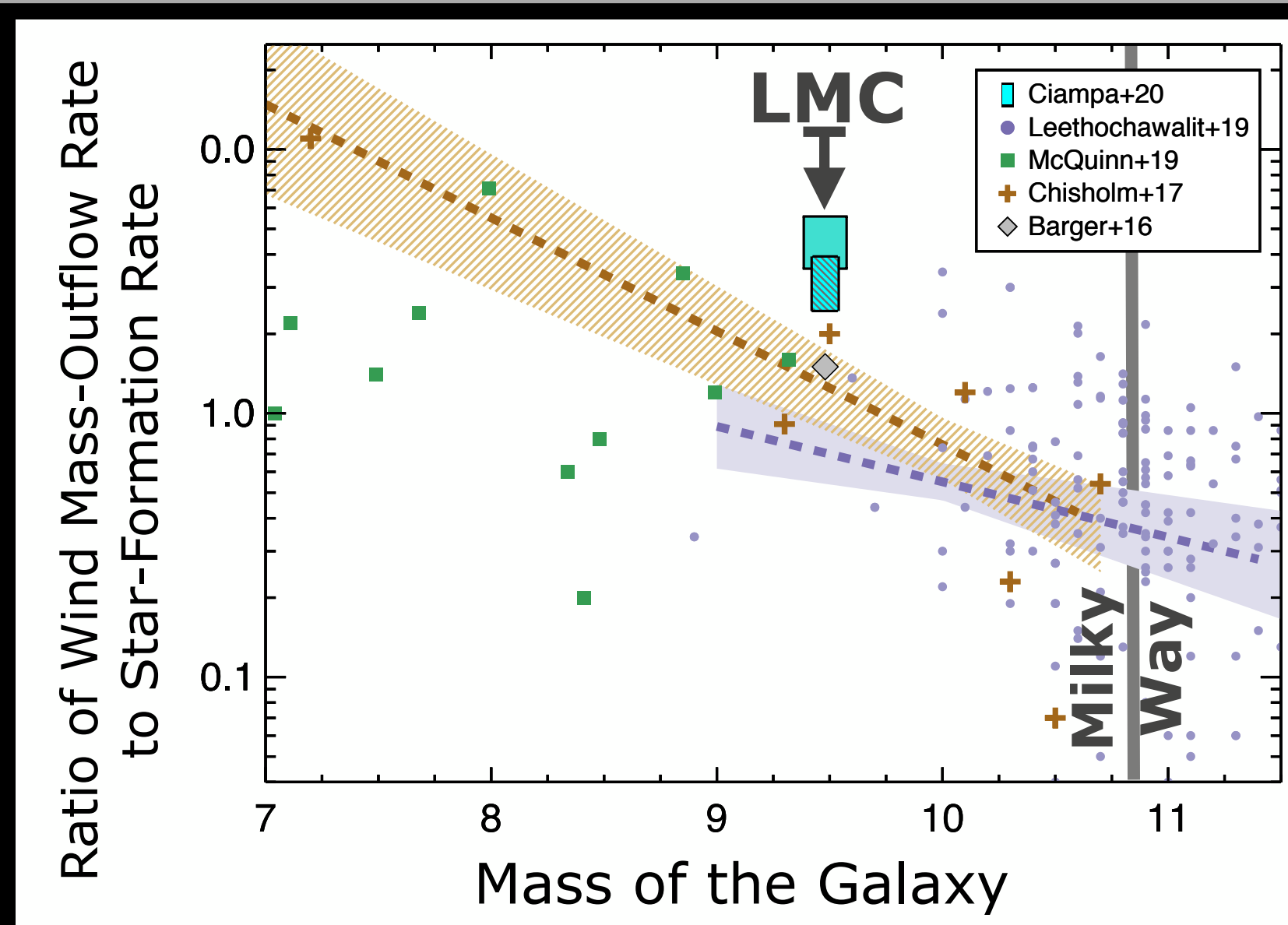
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Loosing
4.5x more
gas in wind
than gas
that it is
converting
into stars
and
planets!