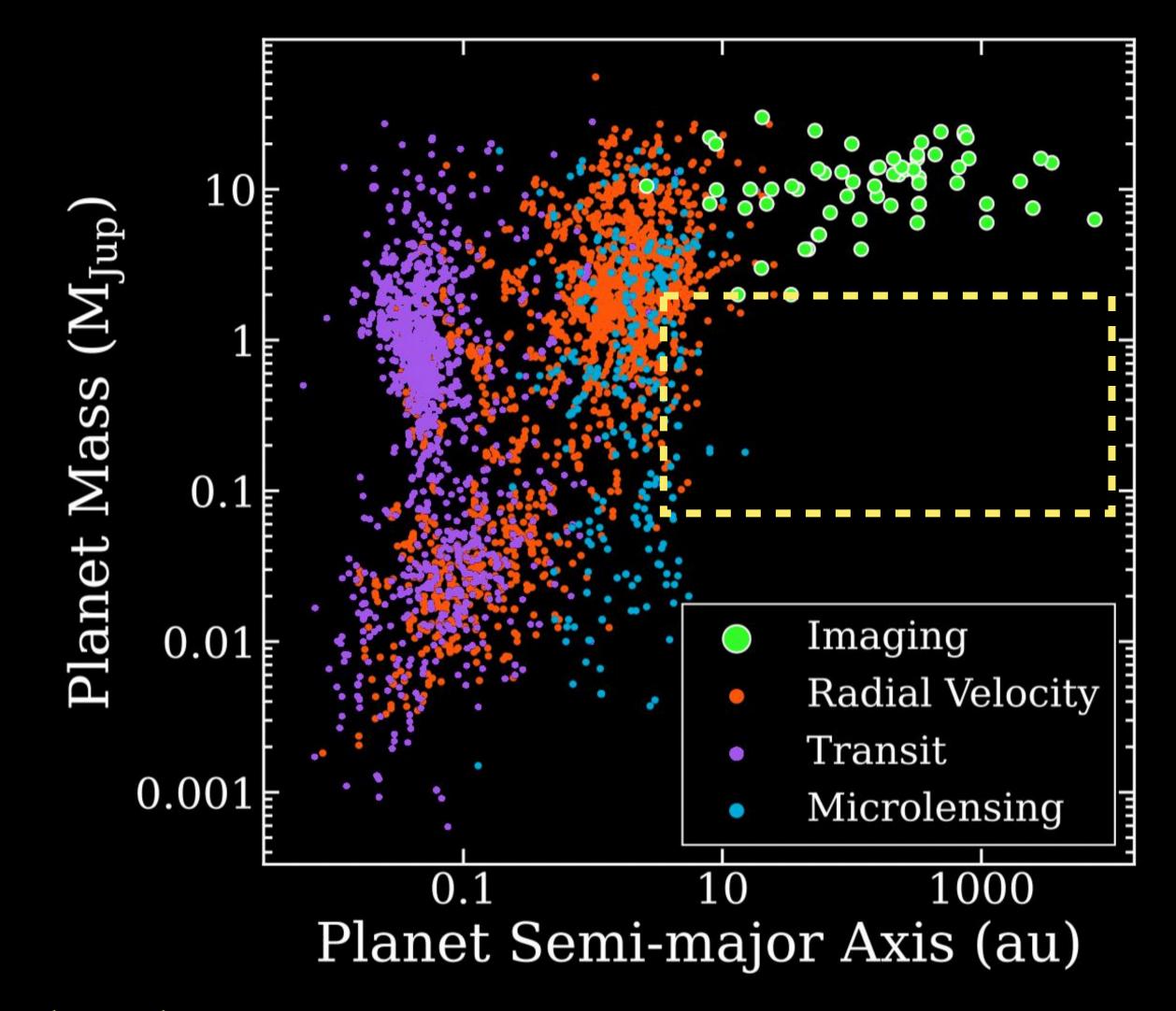


# Direct Imaging of Exoplanet Systems with JWST

- Imaging exoplanets is difficult because of much brighter diffracted starlight
- JWST helps by providing:
  - a stable space-based platform
  - observations at infrared wavelengths

### JWST GTO Program 1184

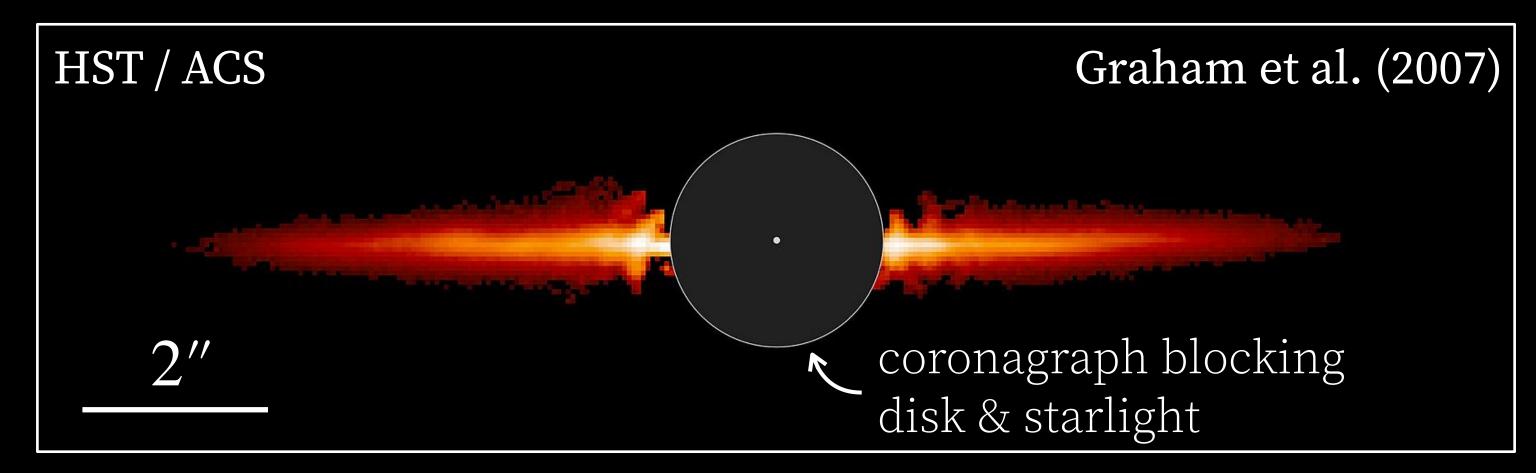
- Uses JWST's NIRCam to search for planets
- Targets 9 red dwarf stars
  - → fainter stars → planets easier to detect

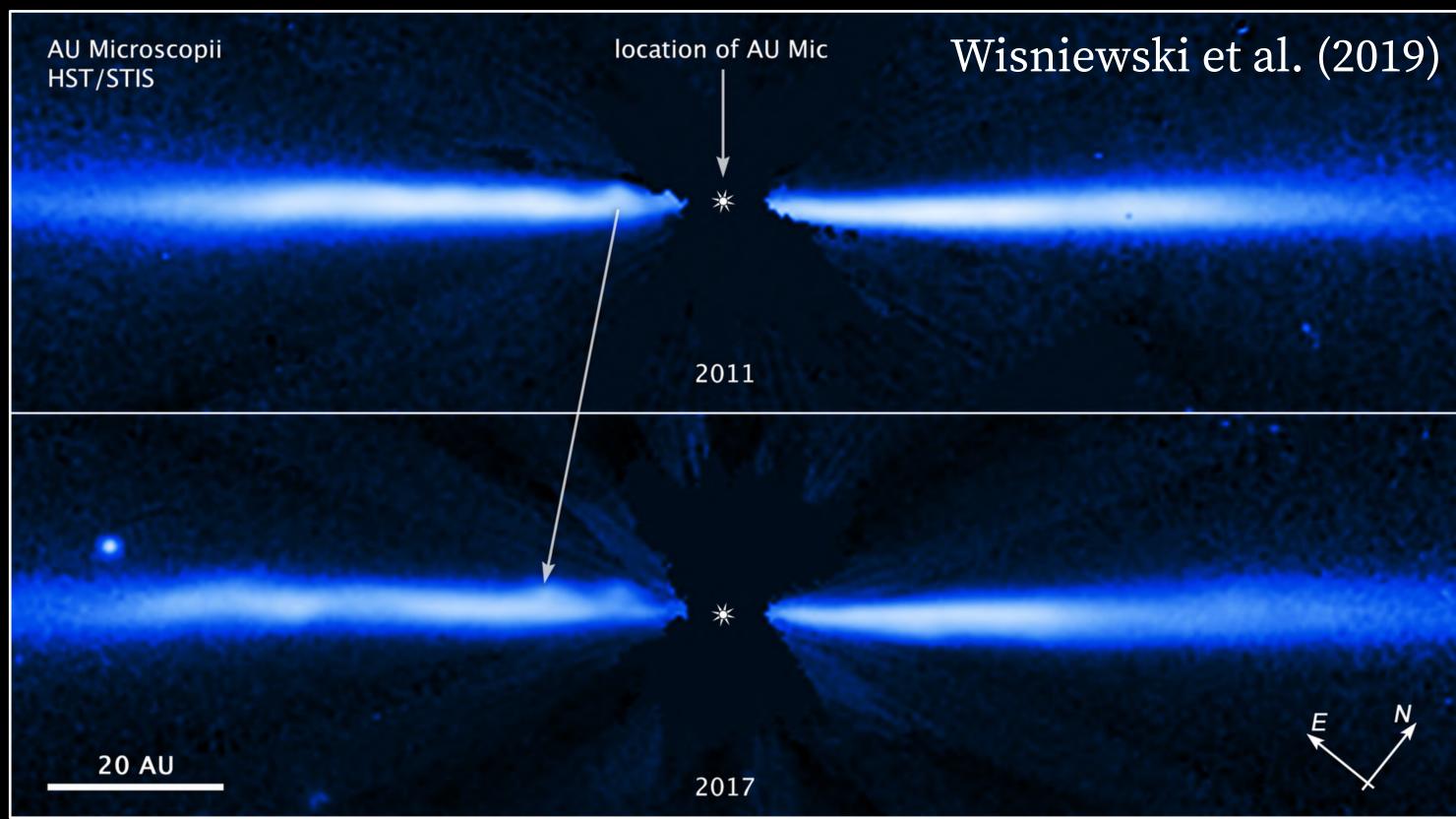


- Sensitive to sub-Saturn mass exoplanets on wide orbits

## The AU Mic System

- 23 million year old red dwarf star
- 32 light years away
- Nearly "edge-on" debris disk
- Numerous fast-moving "clumps" in the disk
- Two small sep. planets detected via transit

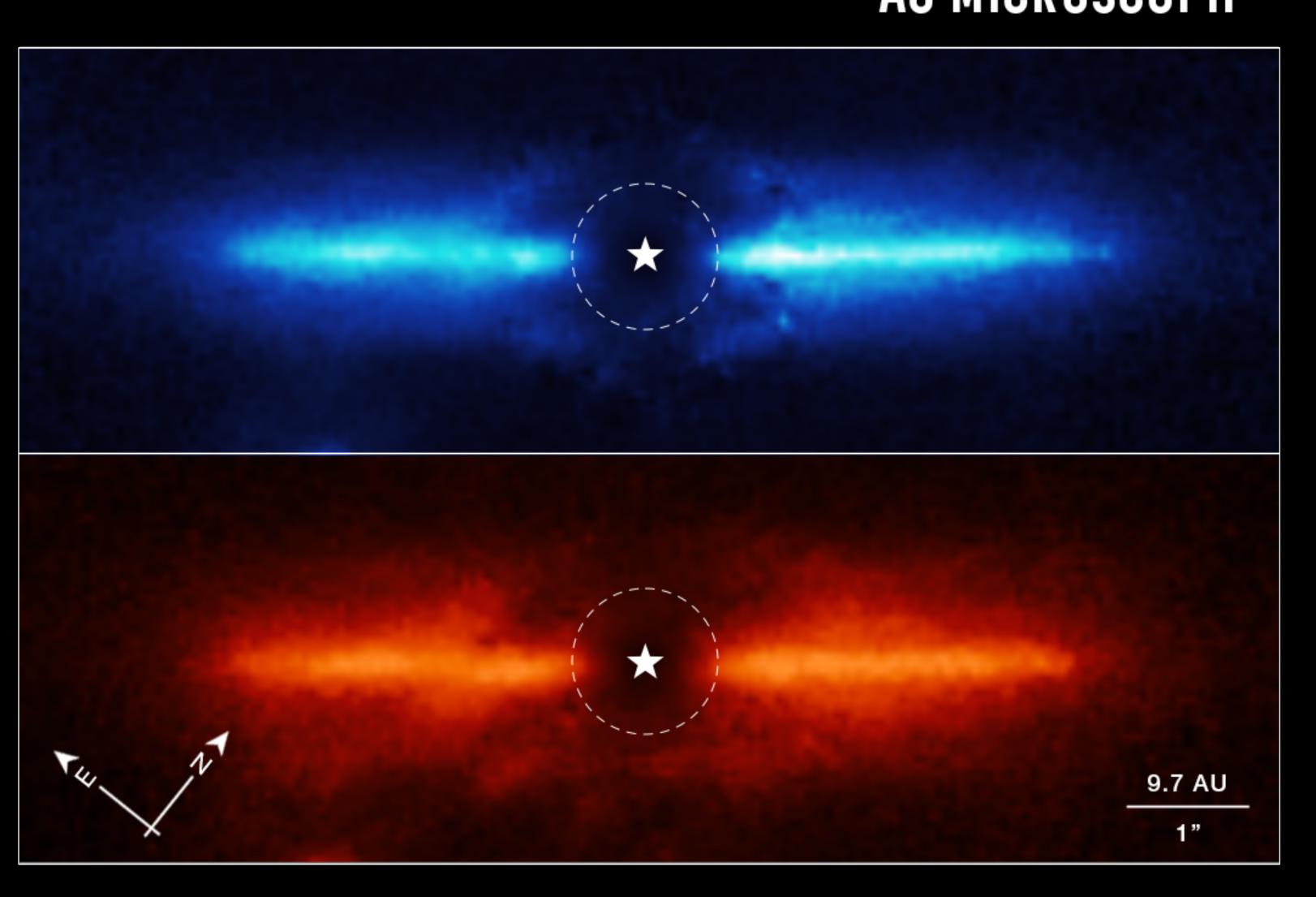




## AU Mic: As Seen by JWST

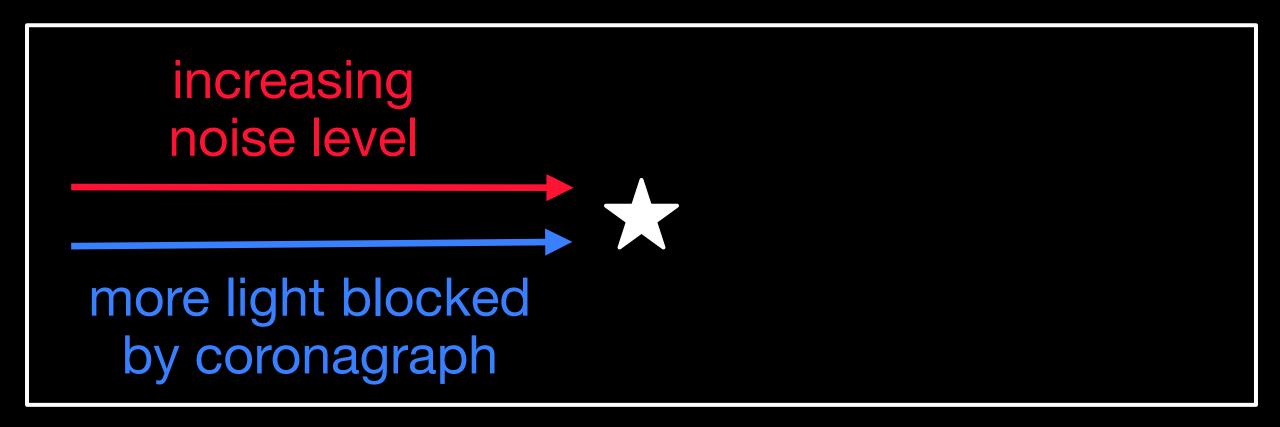
# **AU MICROSCOPII**

- Targeted AU Mic as part of GTO 1184
- Observed at two nearinfrared wavelengths (3.5 & 4.4 microns)
- Detect the disk with high significance
- First detection of the disk at these wavelengths
- Reveals significant "blue" disk color

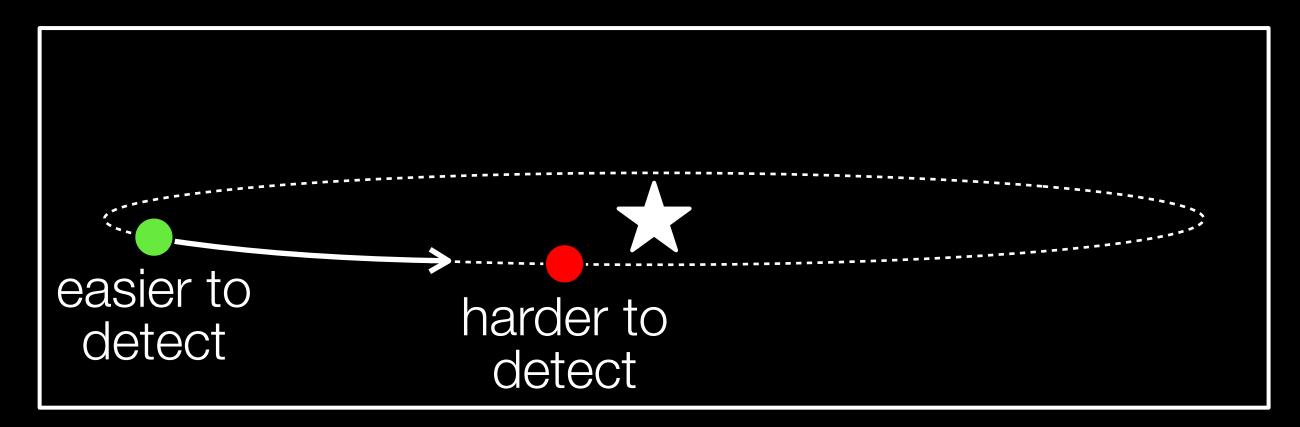


### **AU Mic: The Search for Planets**

- No planets detected
- Measure sensitivity to determine planets we would have seen
- Planets' orbits should be aligned with the disk
- In infrared: more mass = more heat = brighter
- Compute chance of detection vs orbital radius & planet mass

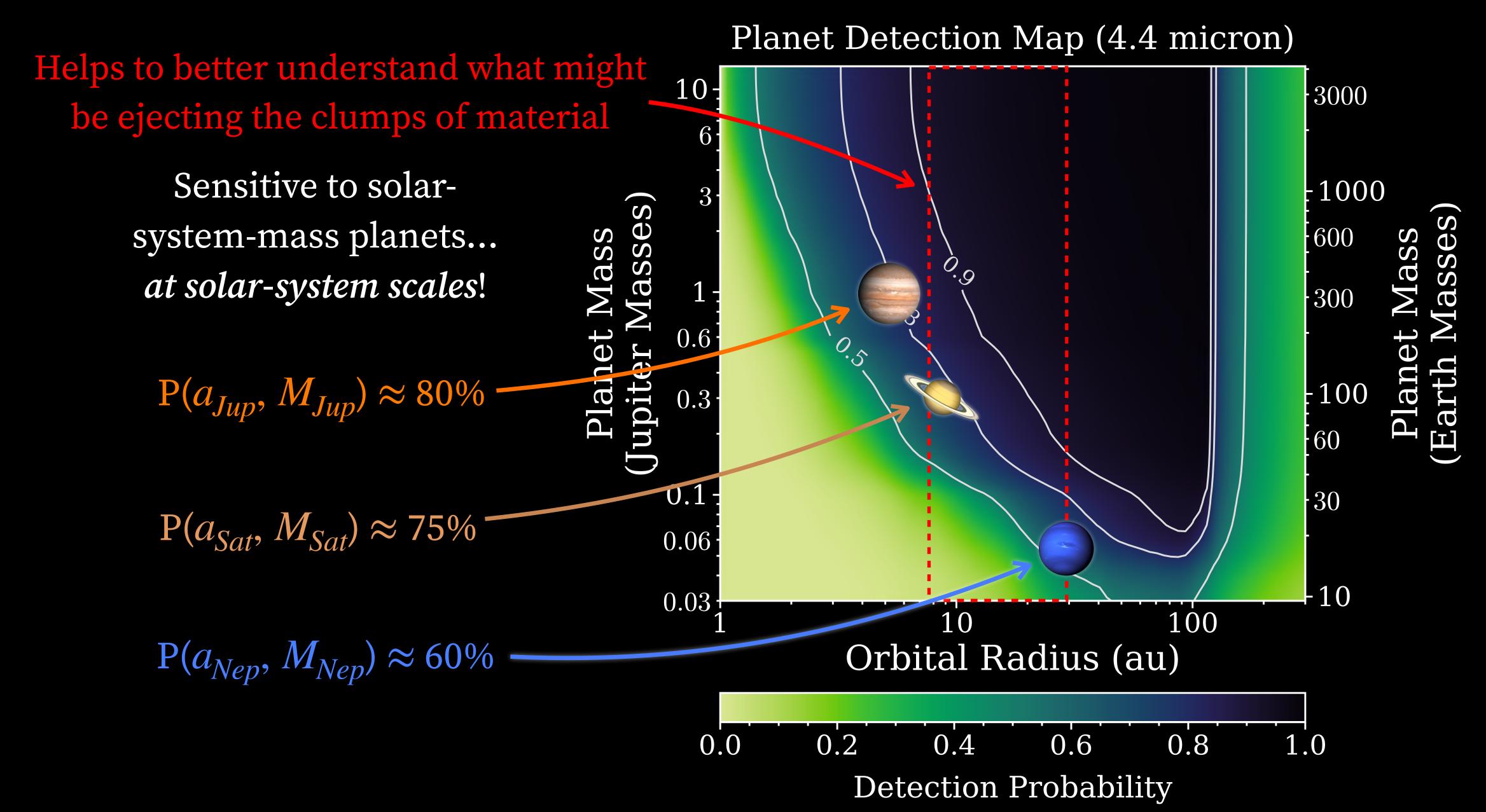


A planet is harder to detect closer to the star



A planet on an inclined orbit is harder to detect at some parts of its orbit

### AU Mic: The Search for Planets



### Summary & Conclusions

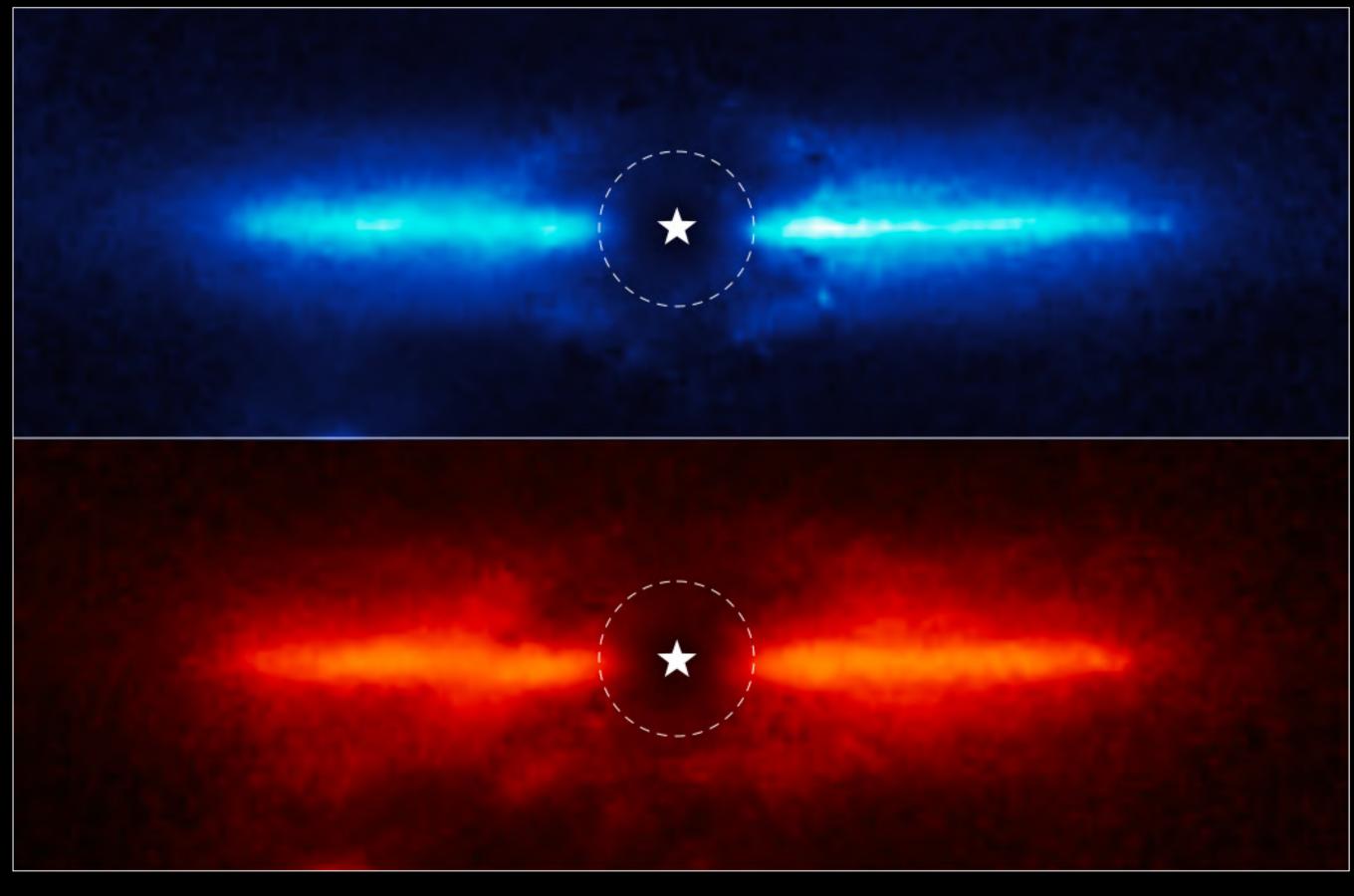
- 1) JWST's NIRCam easily recovers the first 3–4  $\mu m$  images of AU Mic's disk
- 2) Blue disk color at 3–4  $\mu m$
- 3) Data are sensitive to solarsystem-analog planets

JWST's sensitivity & wavelength coverage allow us to explore previously uncharted territory for direct imaging of wide separation planets and faint circumstellar disks.

Thank you!

#### **GTO 1184 Team**

J. Schlieder, M. Meyer, C. Beichman, J. Leisenring, M. Rieke, M. Ygouf, T. Barclay, E. Bogat, G. Bryden, P. Calissendorff, M. De Furio, T. Groff, M. McElwain from NASA-GSFC, U of Michigan, U of Arizona, Caltech-IPAC, and JPL



Contact: Kellen Lawson — kellen.d.lawson@nasa.gov