ALMA reveals extended cool gas and hot ionized outflows around a distant star-forming galaxy

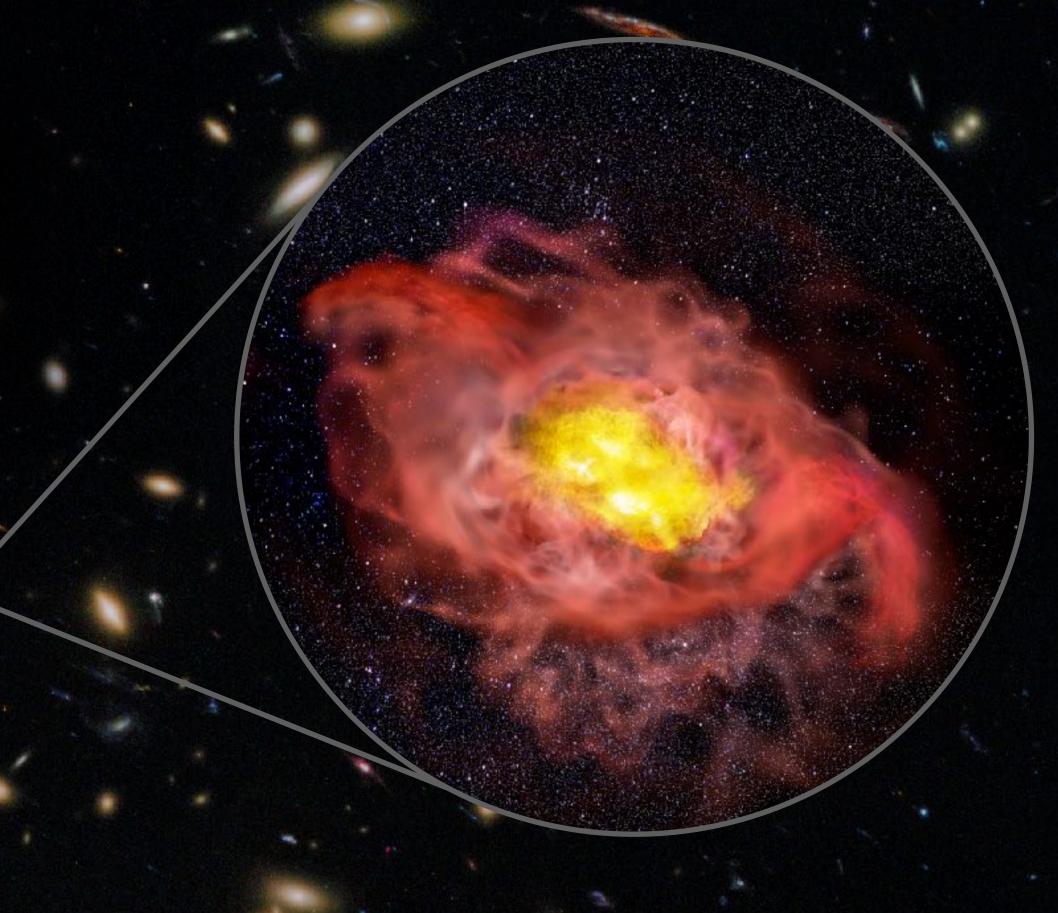
Hollis Akins (Grinnell College → UT Austin)

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Akins et al. (2022), ApJ in press iPoster 241.23: today, 5:30-6:30 pm PDT, hall A/B hollis.akins@gmail.com | (336) 681-8118







Background: Hubble view of Abell 1689 (ESA/Hubble)

Inset: Artist's illustration of A1689-zD1 (B. Saxon/NRAO/AUI/NSF)



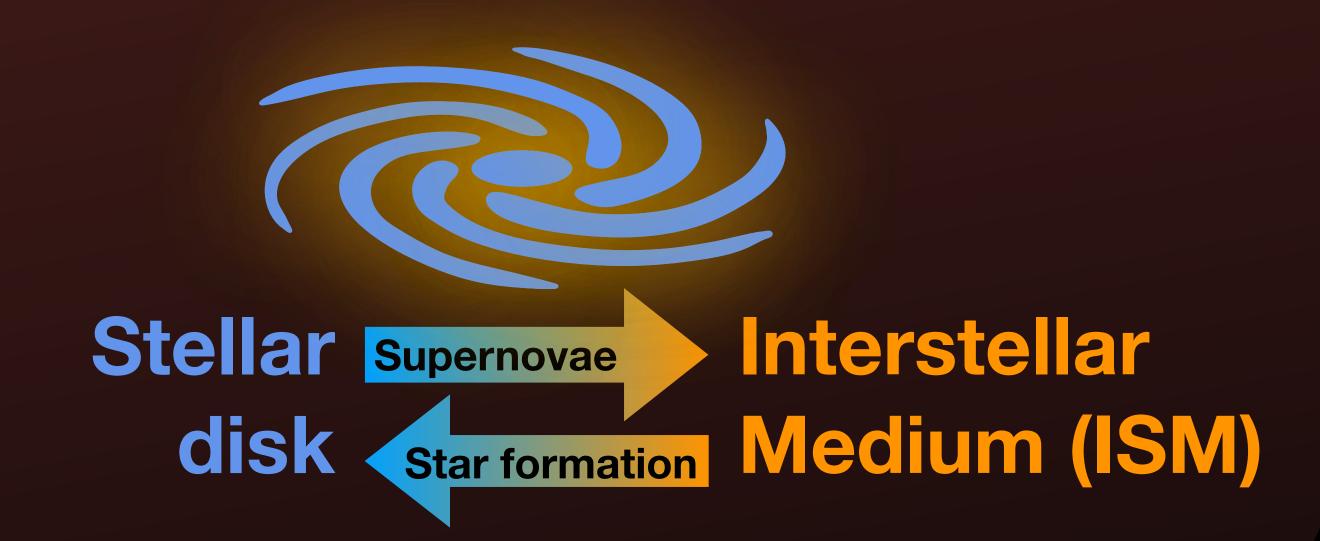
 The ISM is mostly hydrogen gas, with some heavier elements like carbon or oxygen



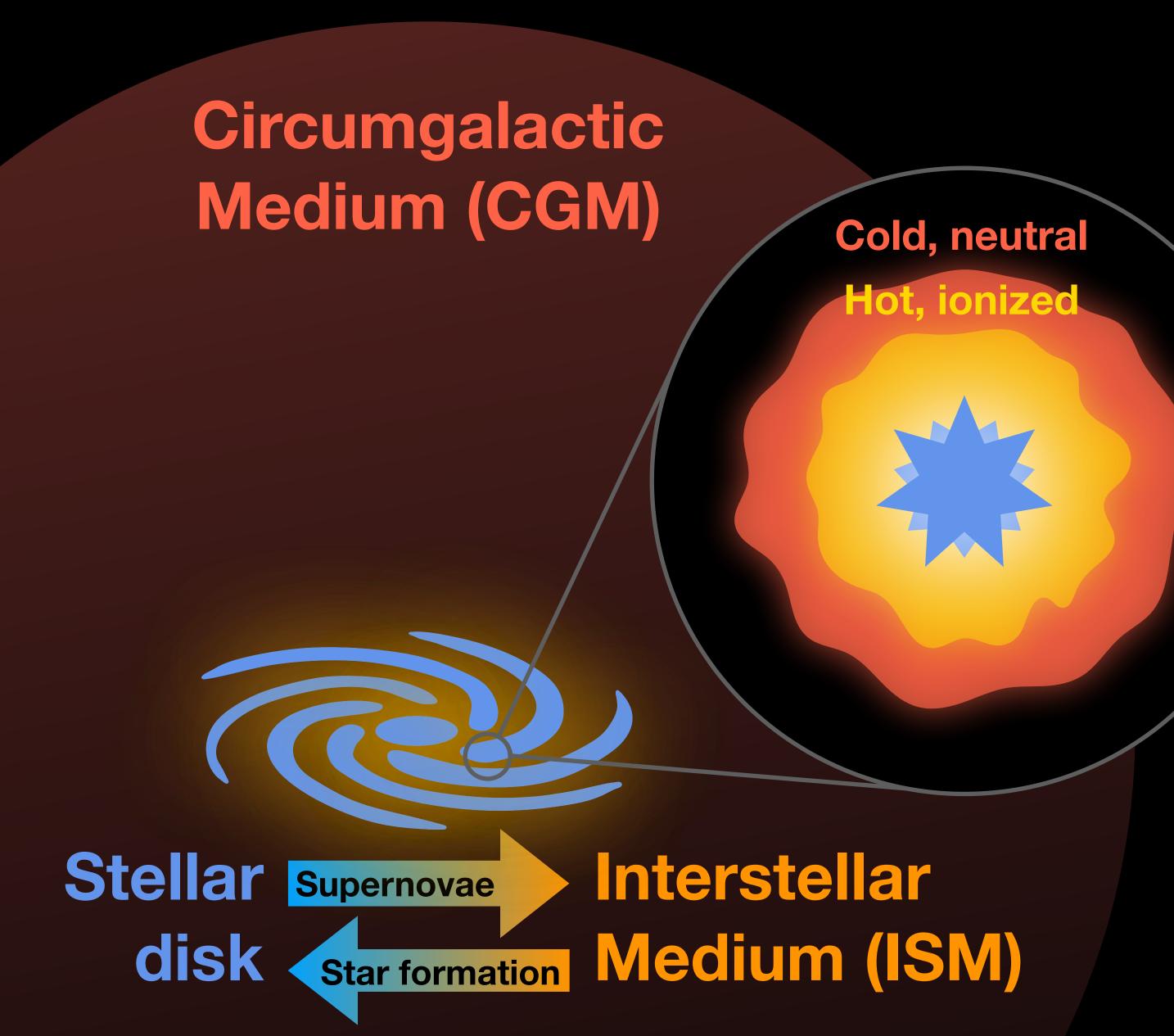
- The ISM is mostly hydrogen gas, with some heavier elements like carbon or oxygen
- ISM gas is constantly being formed into stars/ replenished by supernovae



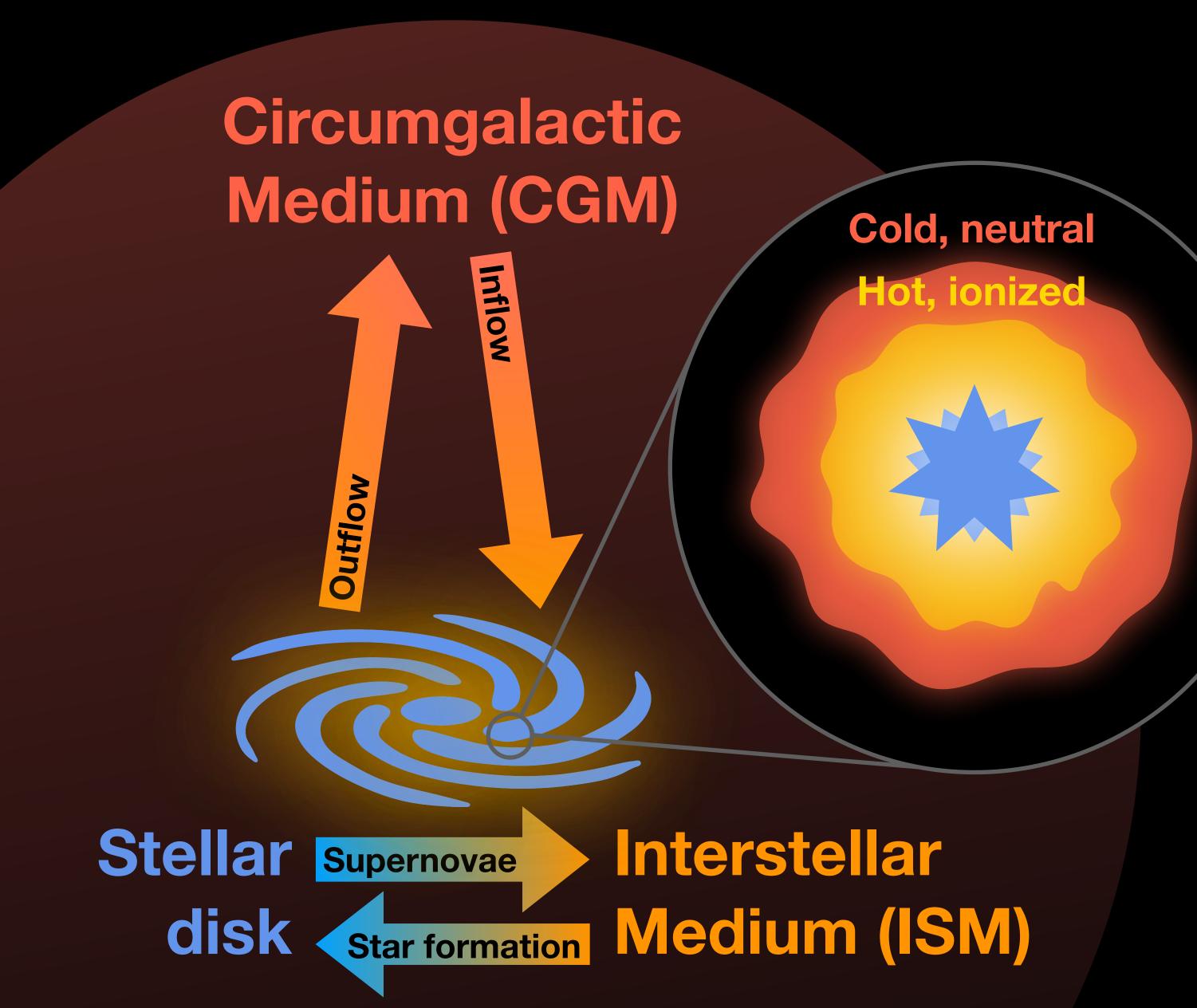
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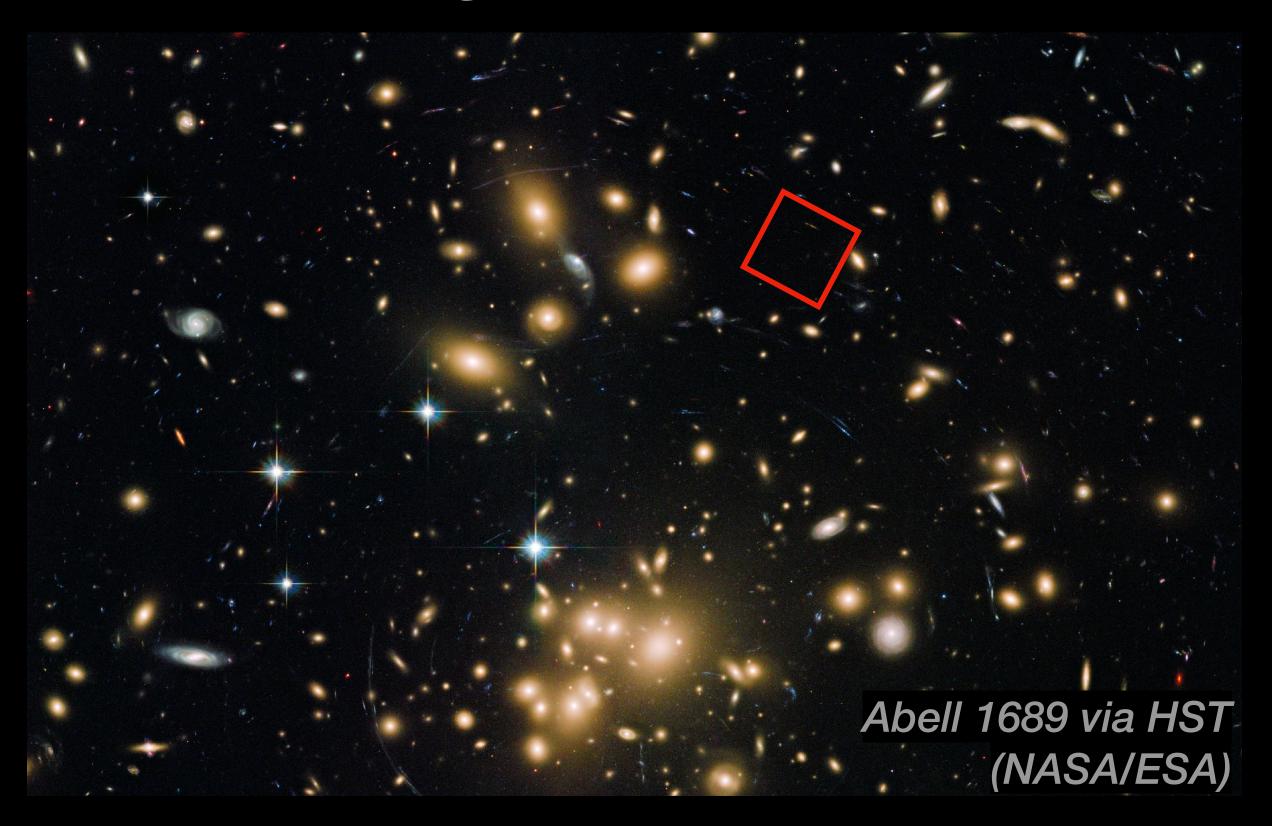


- The ISM is mostly hydrogen gas, with some heavier elements like carbon or oxygen
- ISM gas is constantly being formed into stars/
 replenished by supernovae
- The ISM can exist in many phases
- Gas can also be exchanged between the ISM and CGM

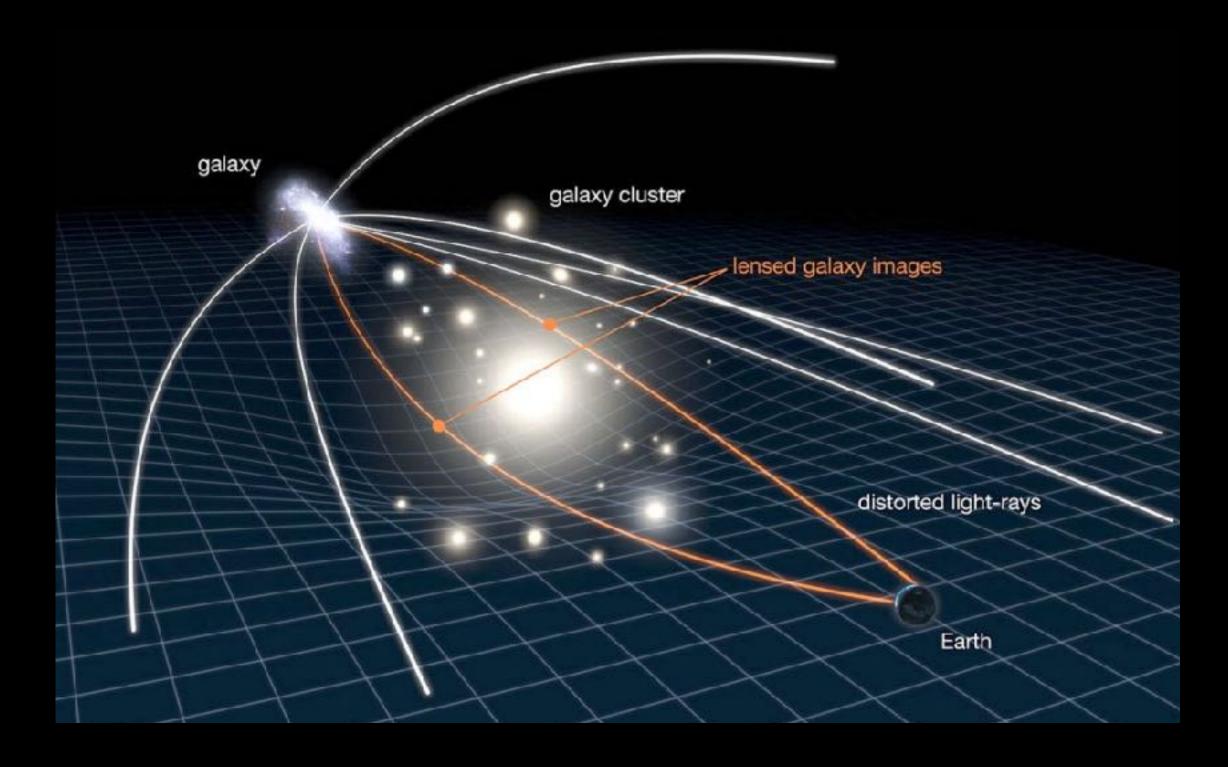




- A1689-zD1 is roughly 13 billion light years away

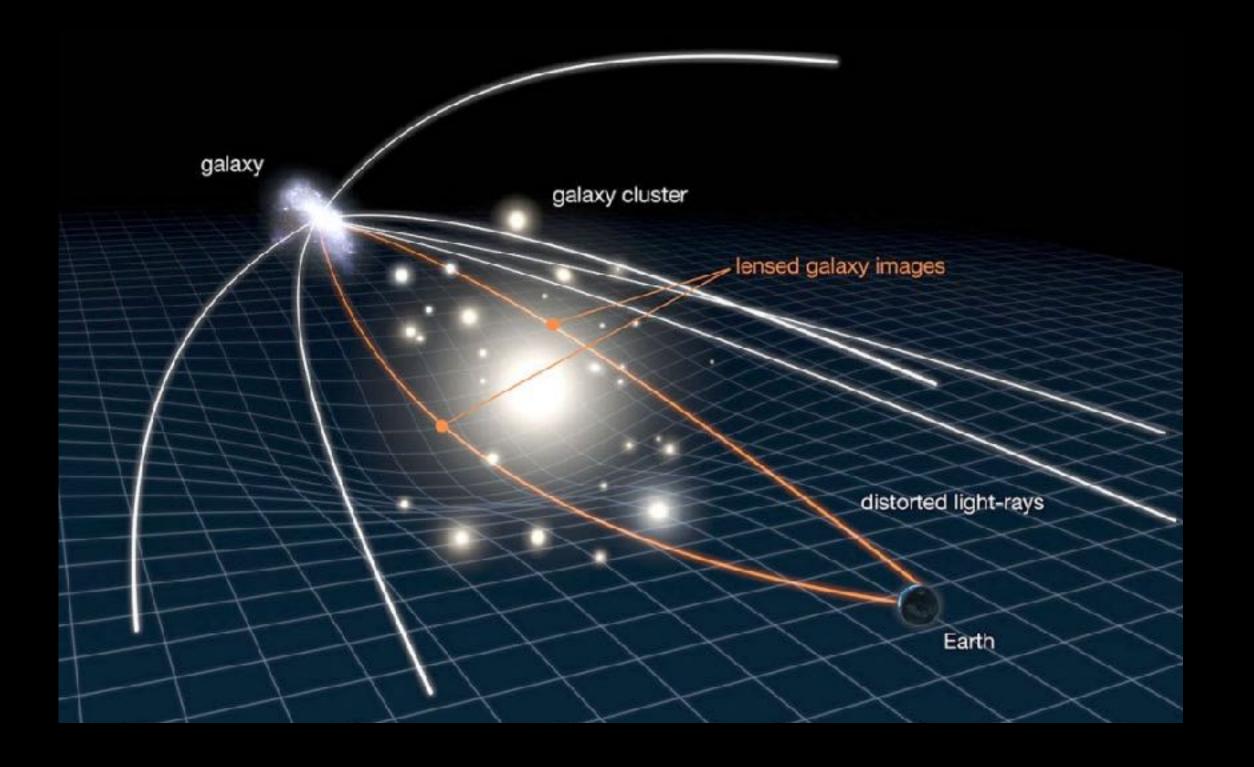


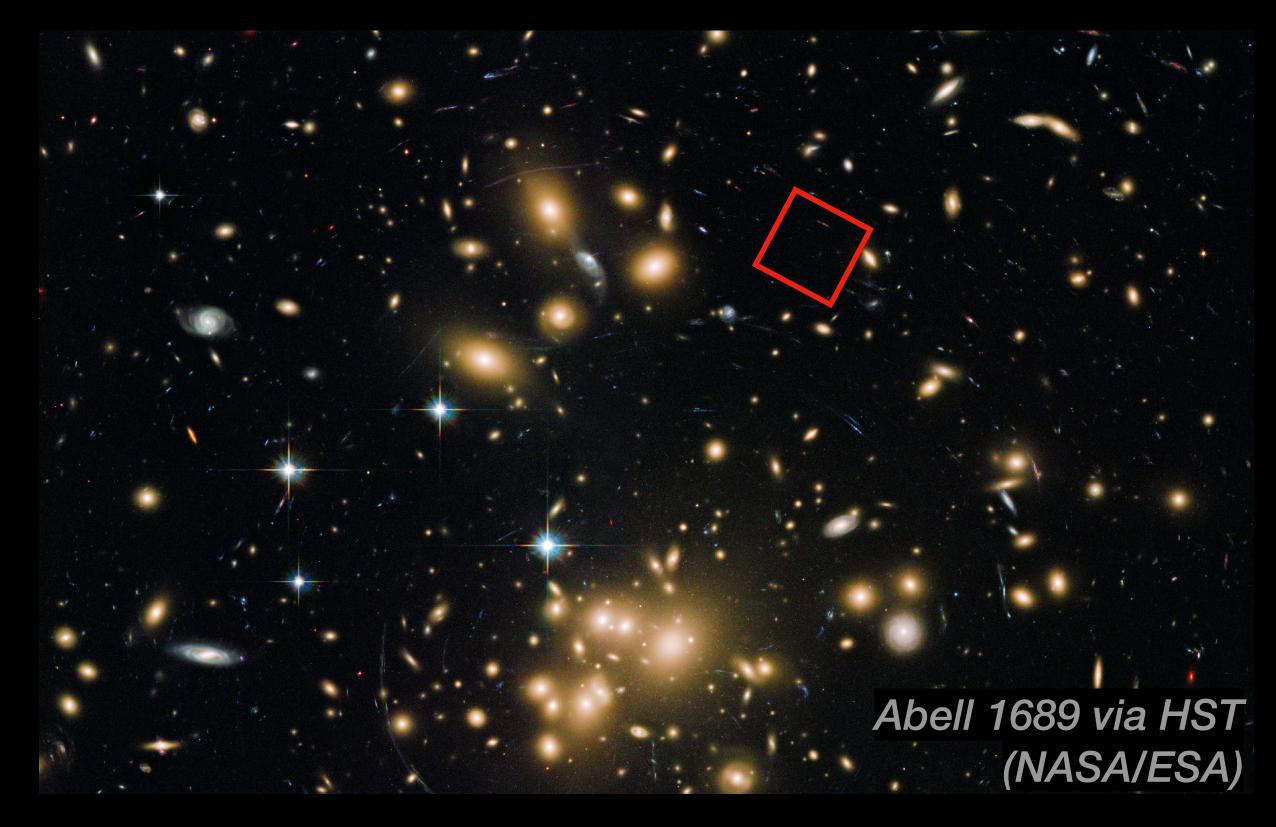
- A1689-zD1 is roughly 13 billion light years away
- The cluster Abell 1689 magnifies the galaxy by 10x





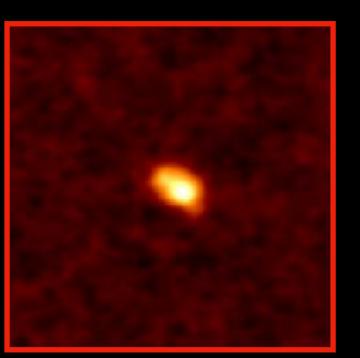
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- Relatively "normal" galaxy



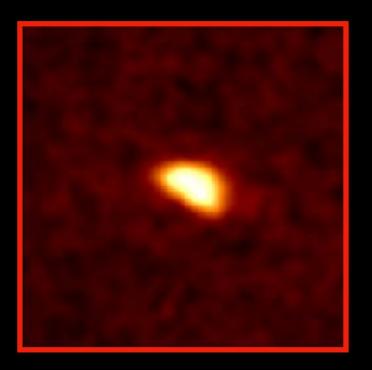


- A1689-zD1 is roughly 13 billion light years away
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- Relatively "normal" galaxy
- We use new ALMA observations of [OIII] and [CII] to observe the ISM and archival HST data to observe the ultraviolet emission from the stars

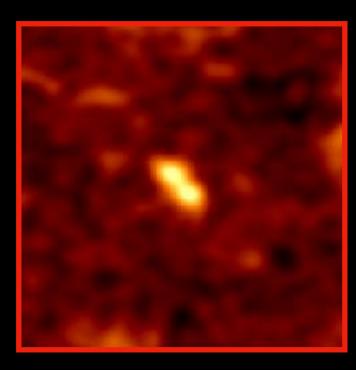




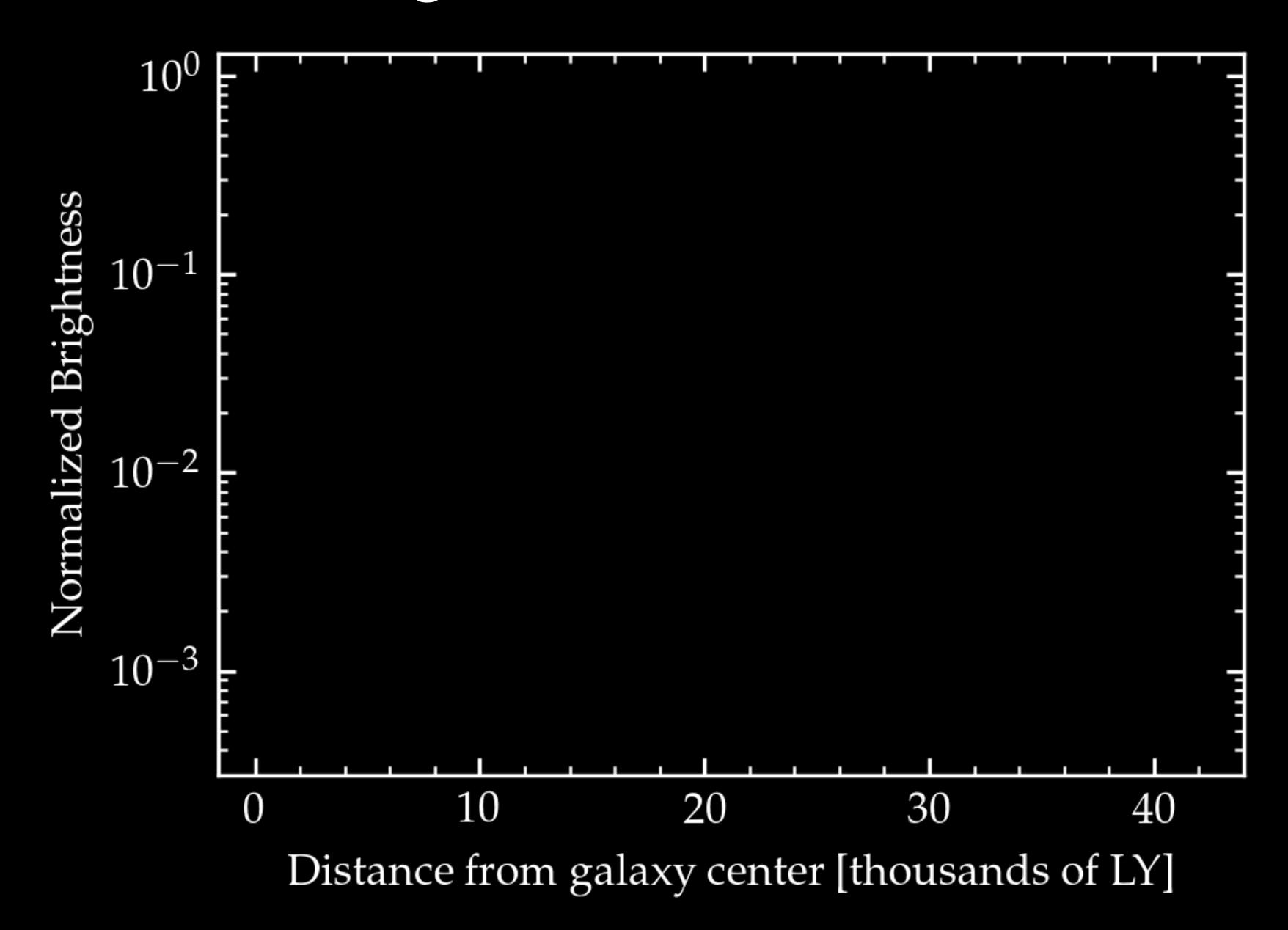


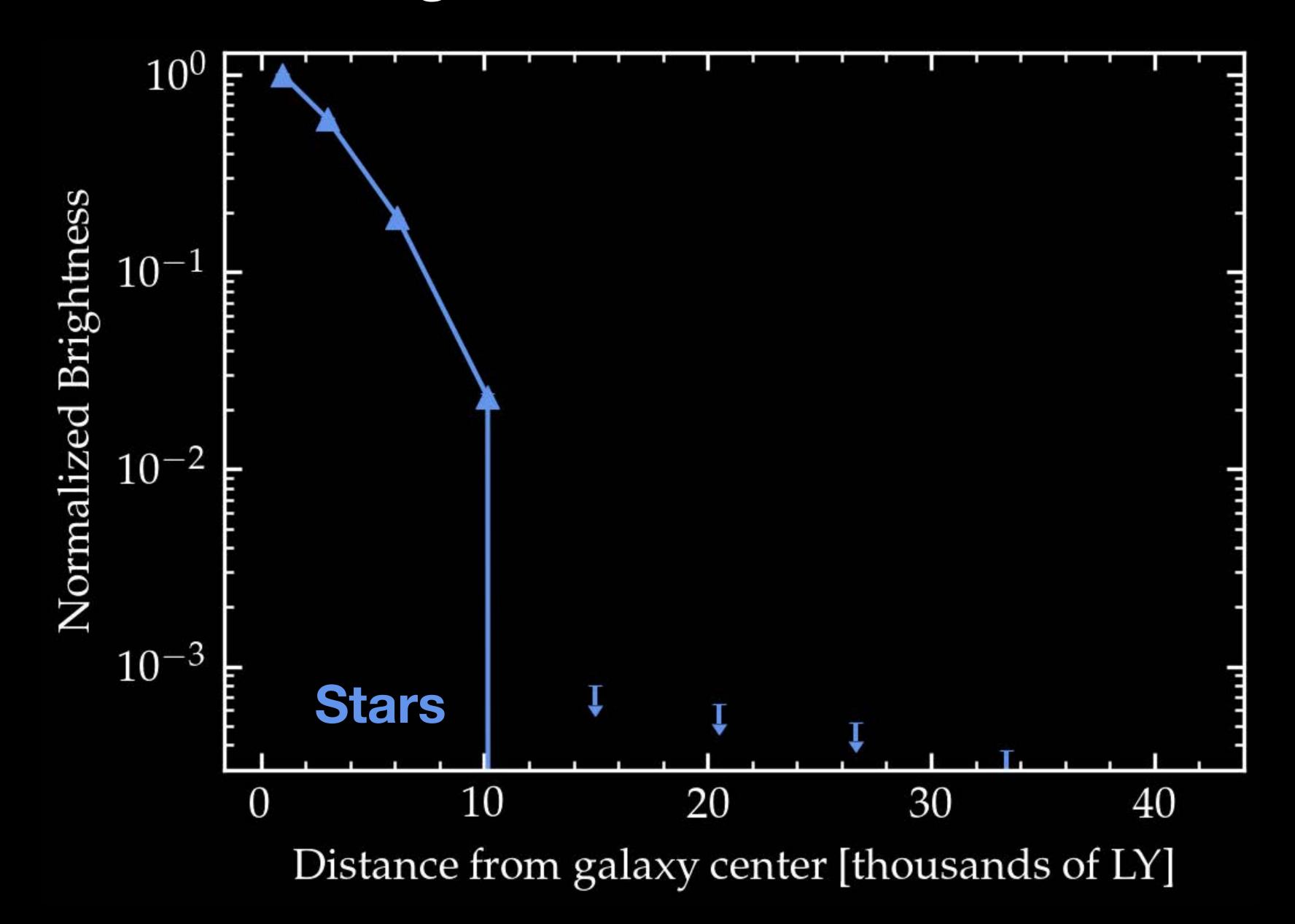


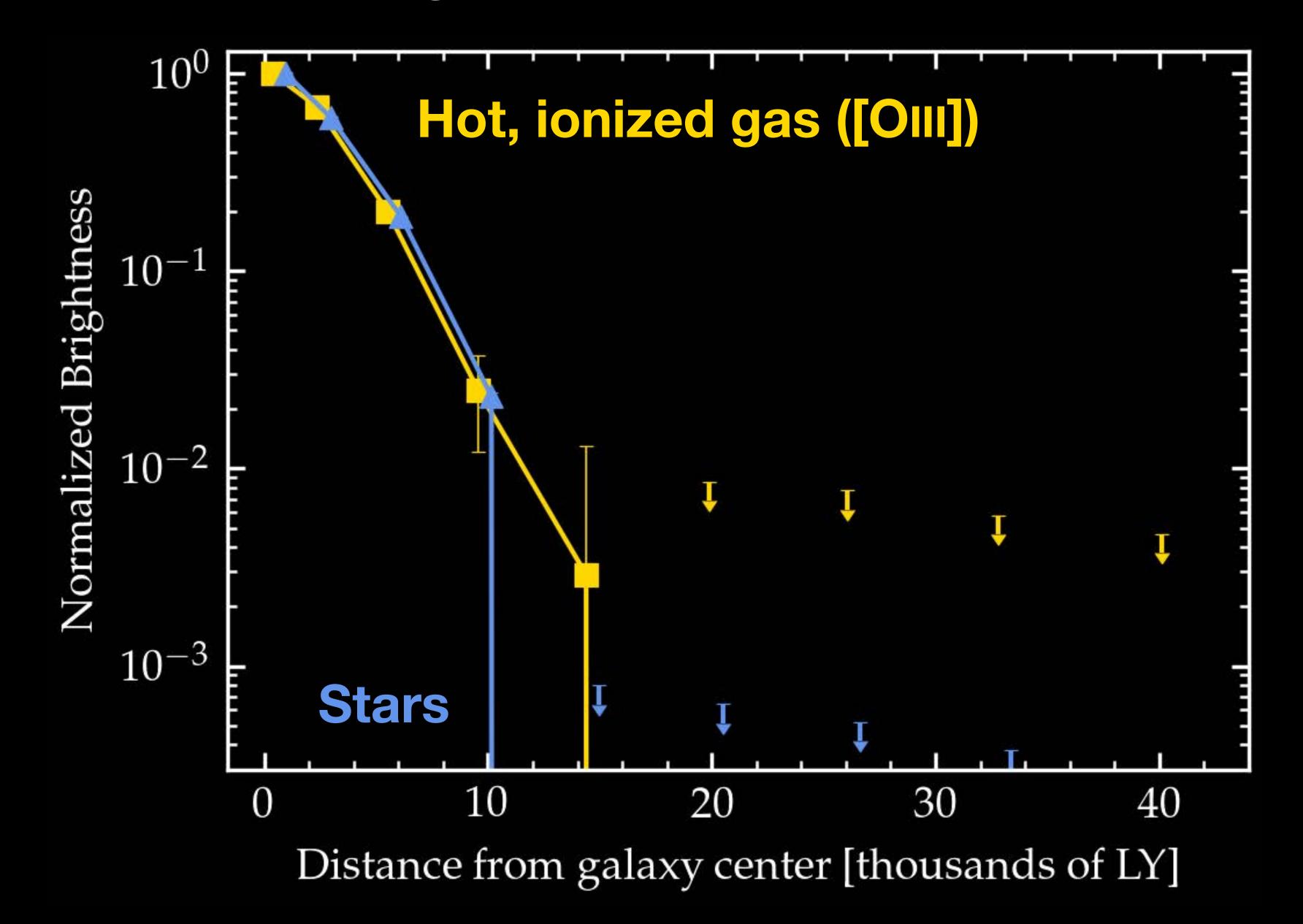
[CII] emission (cold, neutral

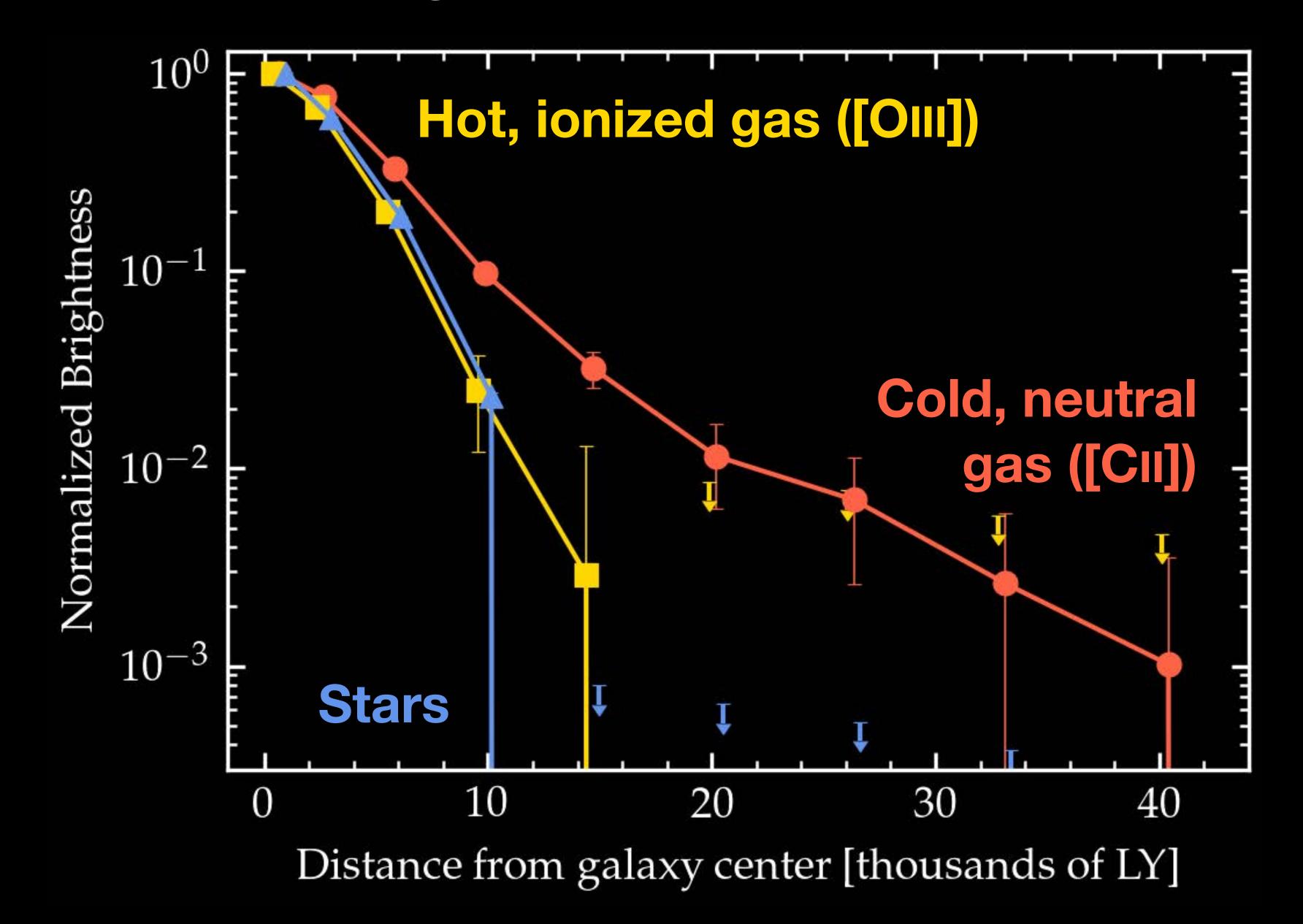


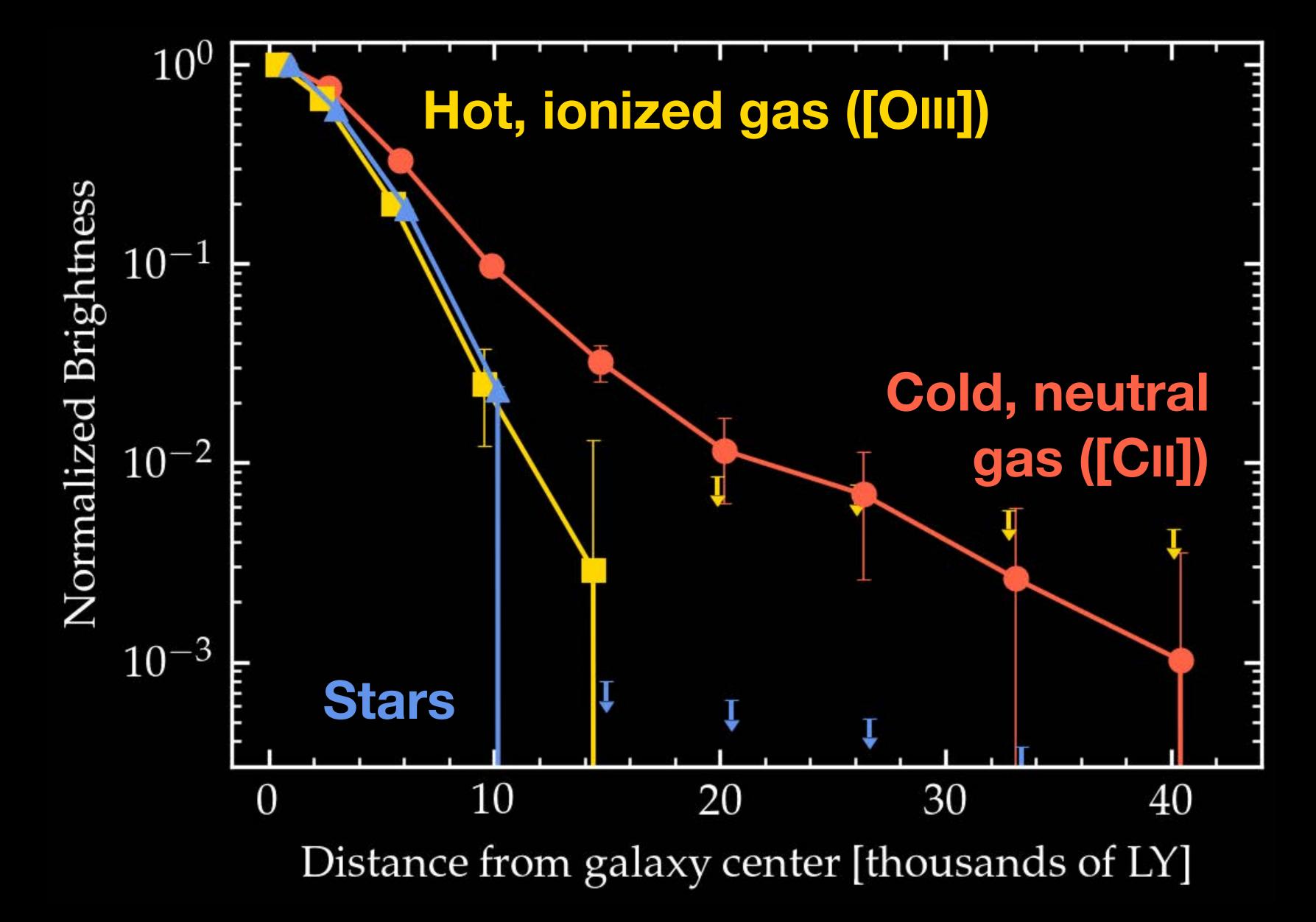
Ultraviolet emission (stars) via HST





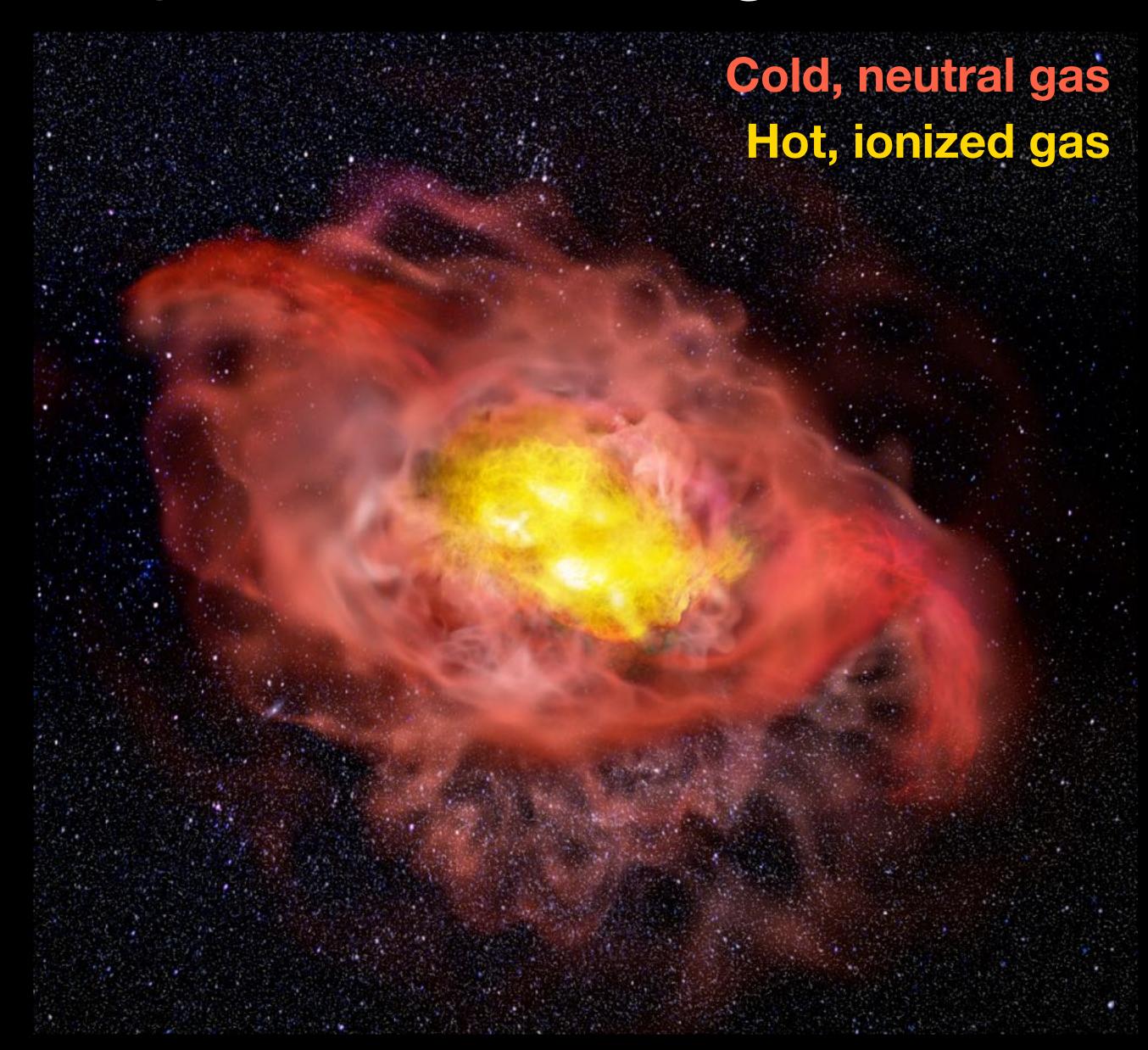






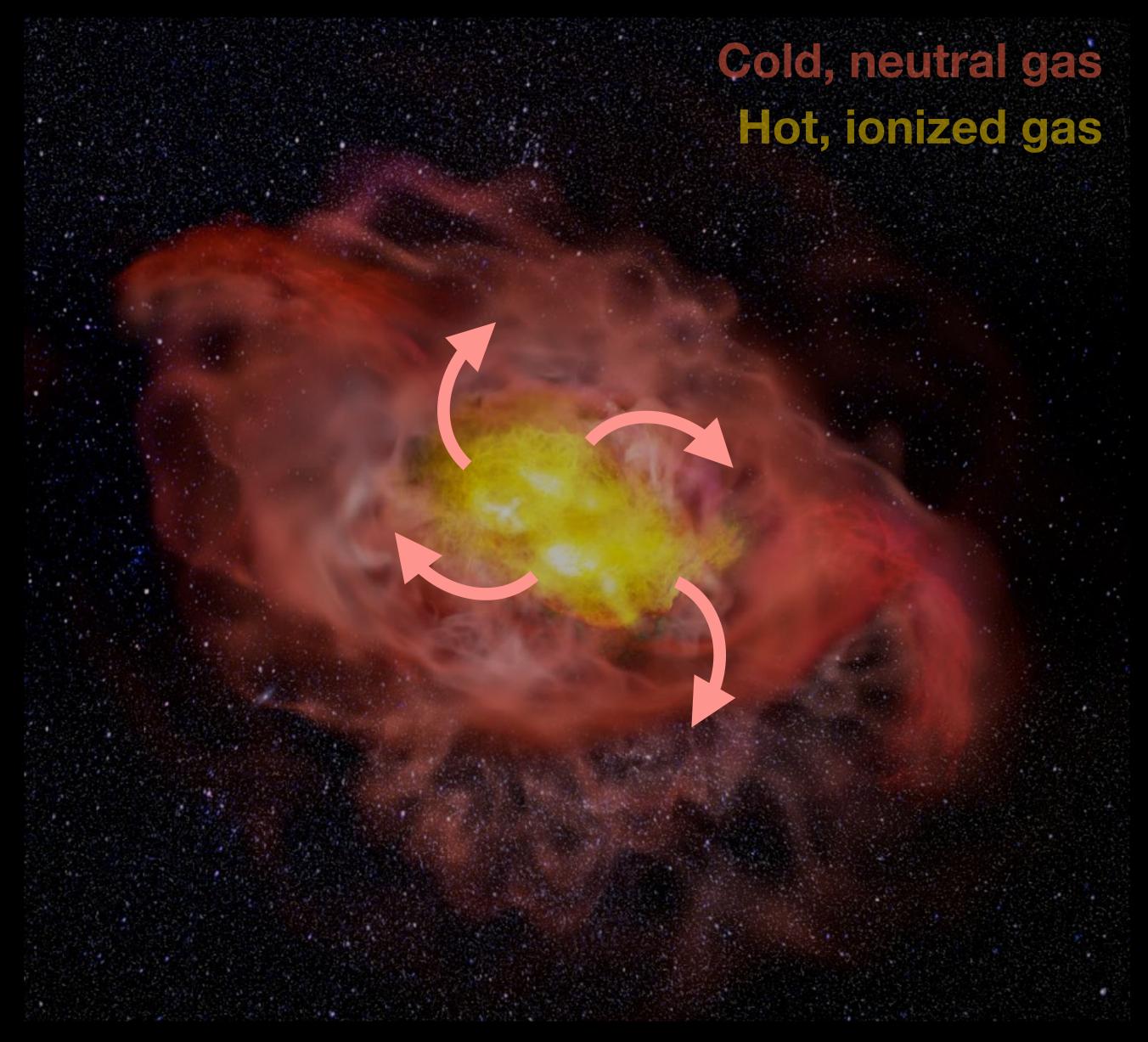
The cold, neutral gas extends significantly further than the stars or the hot, ionized gas

Why is the neutral gas extended?



Artist's illustration of A1689-zD1 (B. Saxon/NRAO/AUI/NSF)

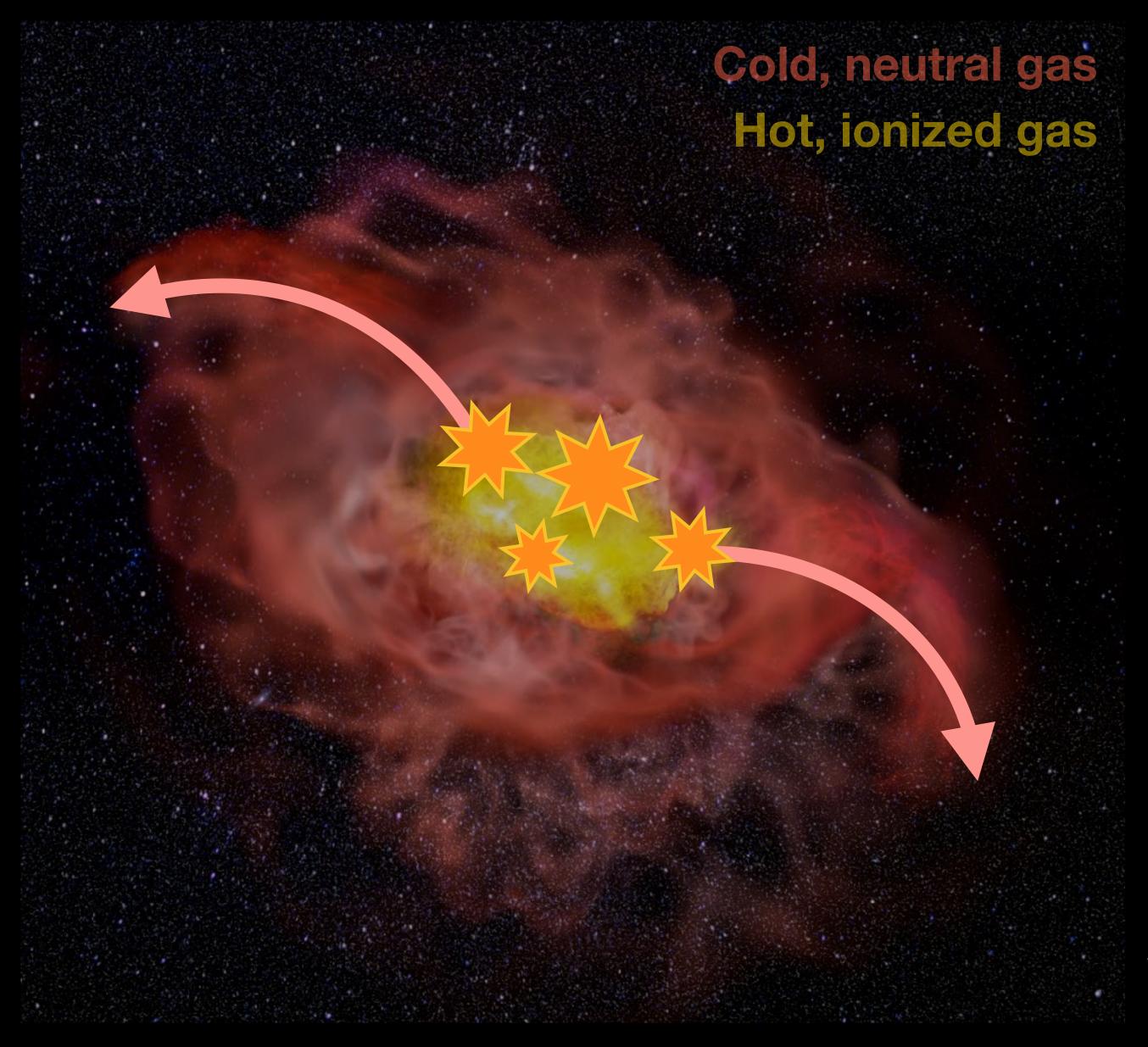
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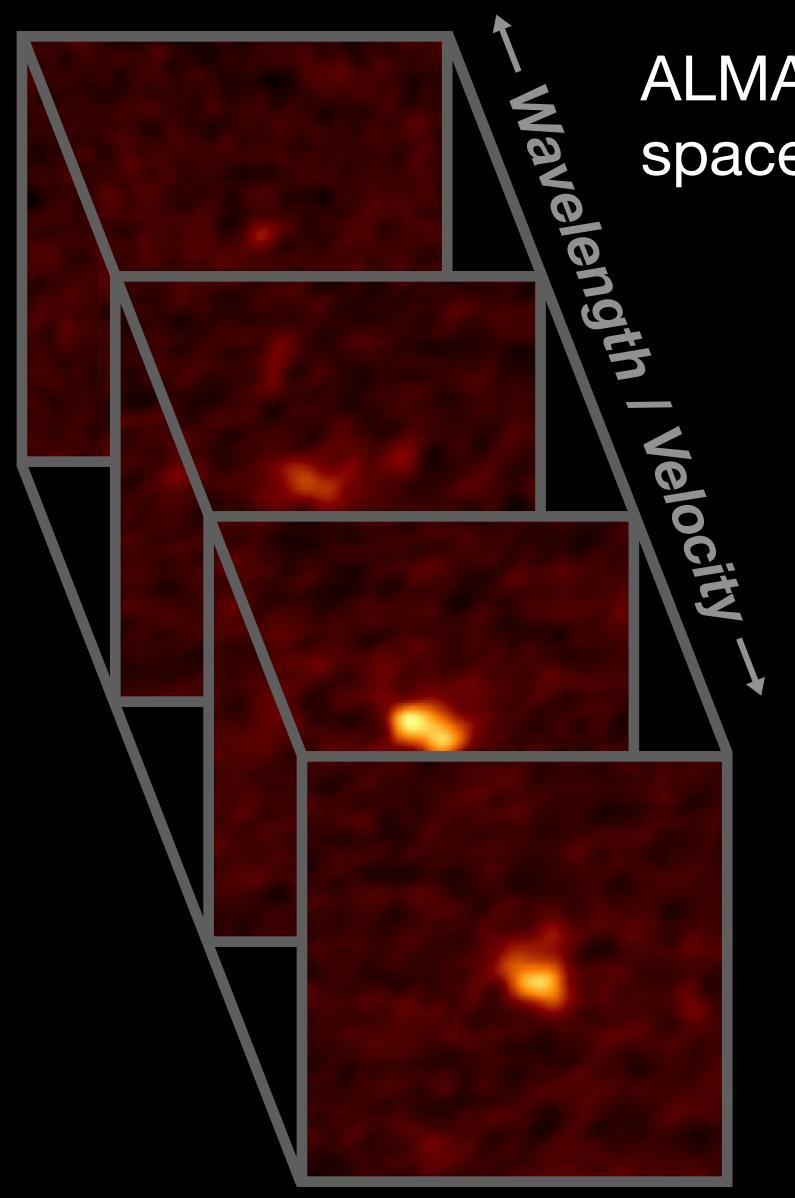
Why is the neutral gas extended?



- Past mergers/ gravitational interactions with other galaxies
- 2. Outflow of gas driven by galactic activity (supernovae, AGN)

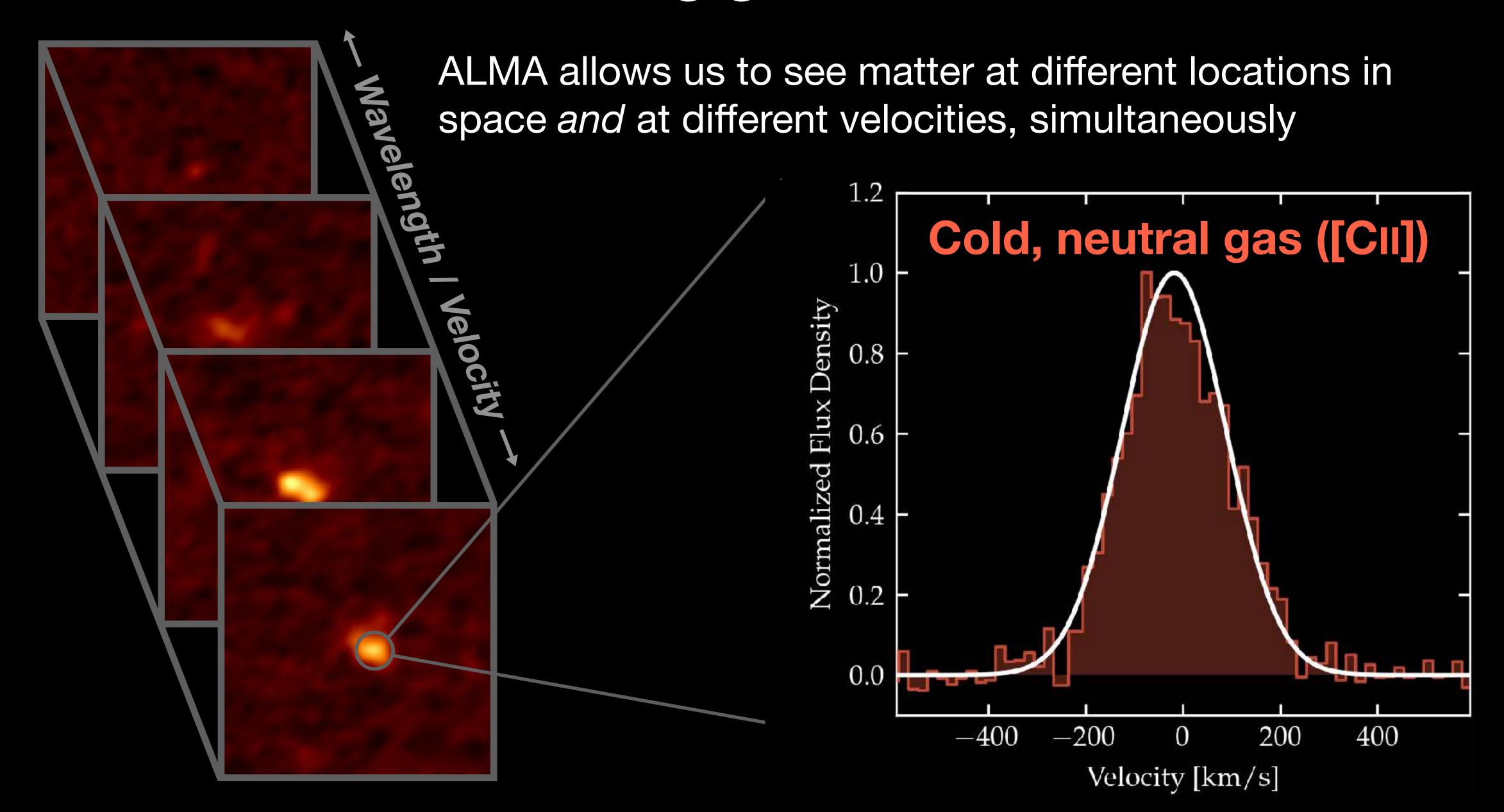
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Evidence for outflowing gas

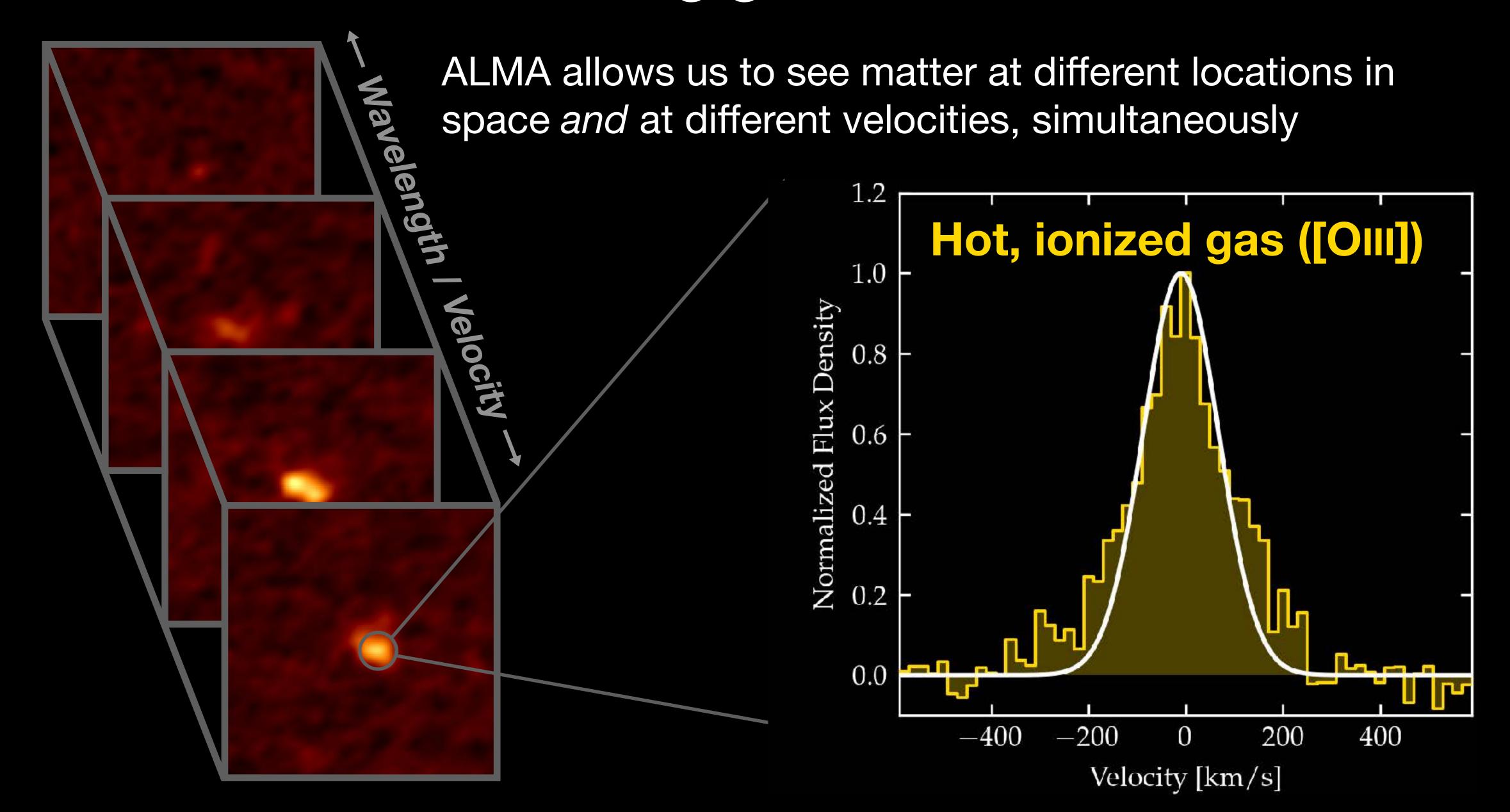


ALMA allows us to see matter at different locations in space and at different velocities, simultaneously

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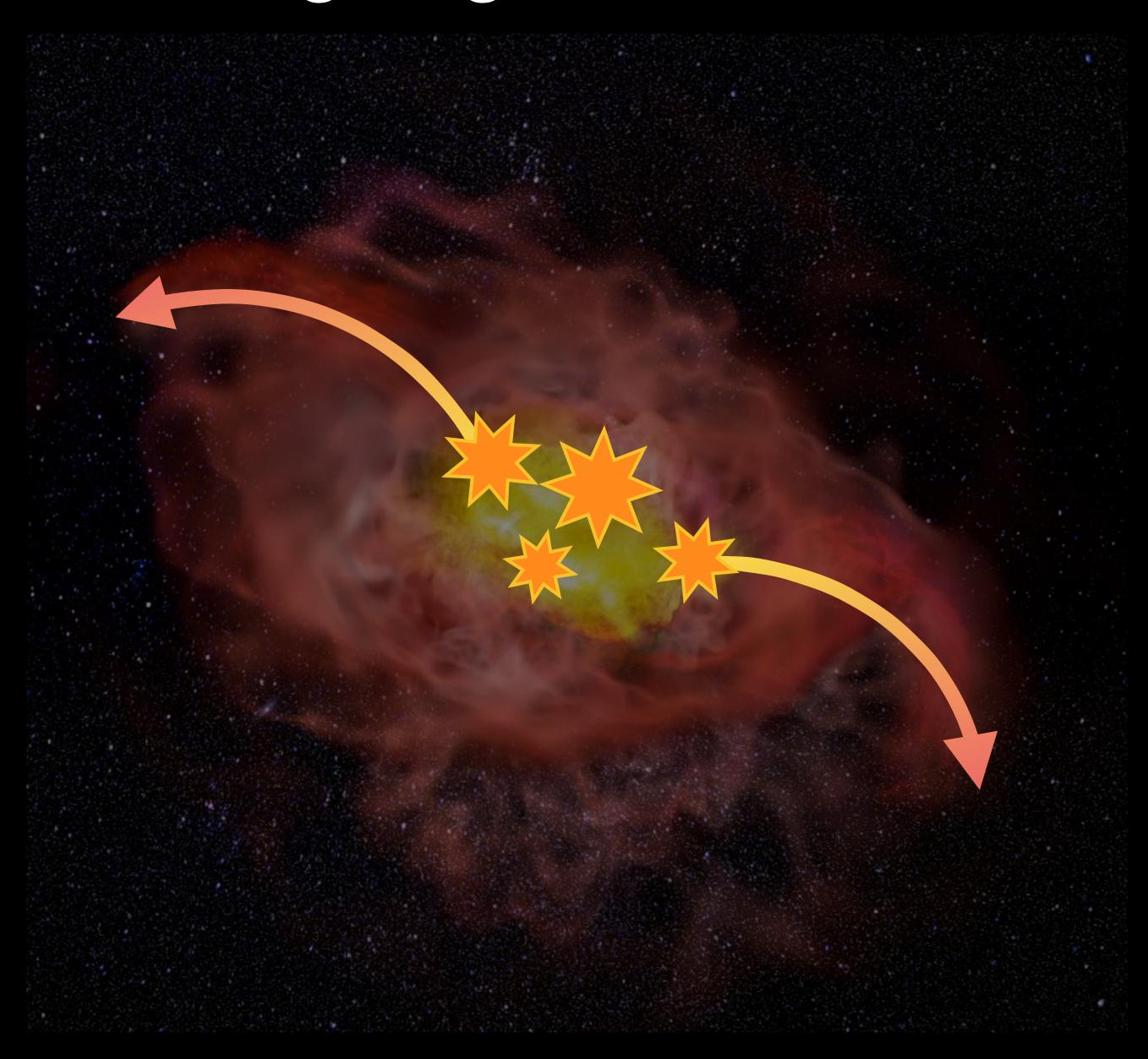


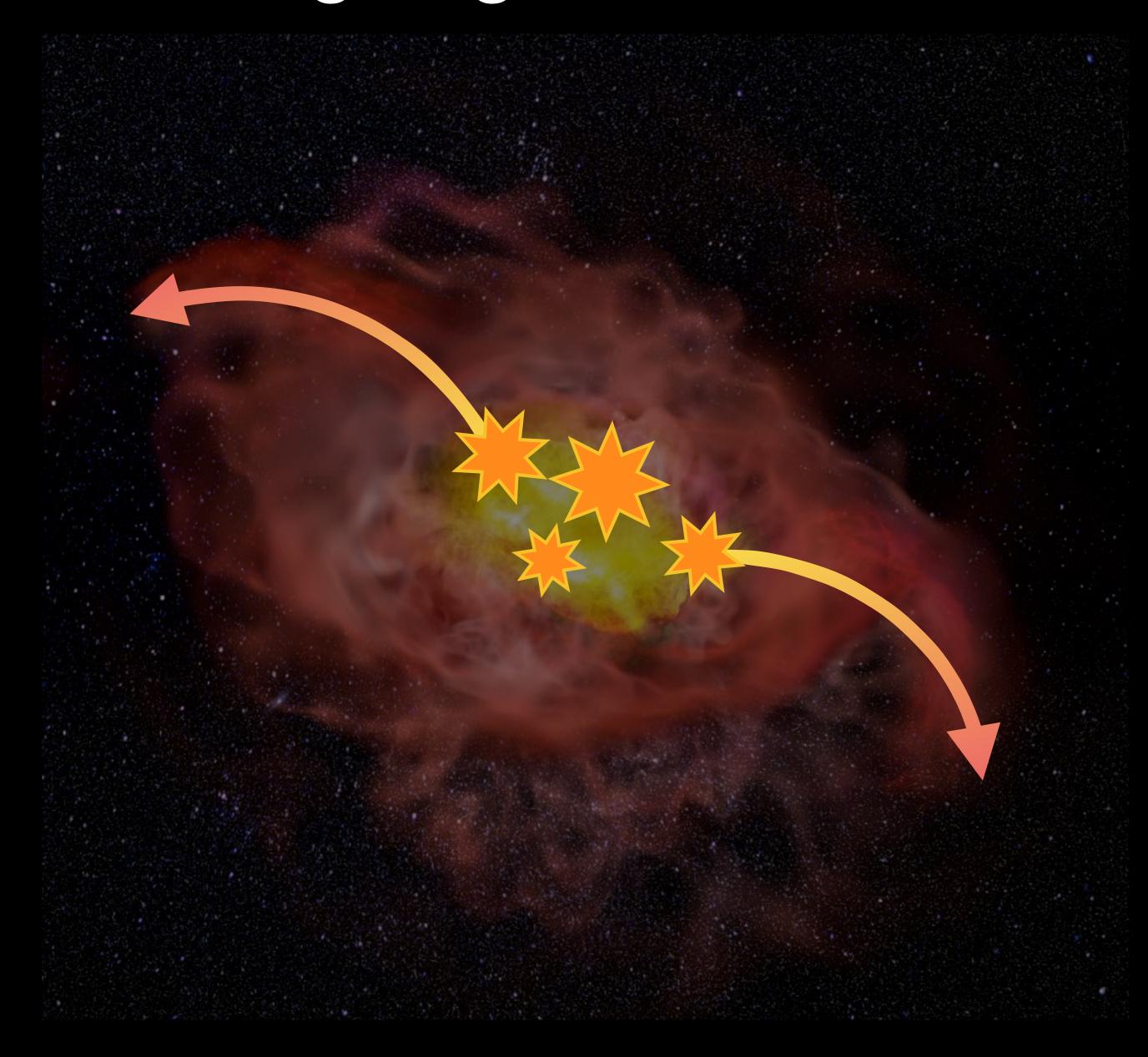
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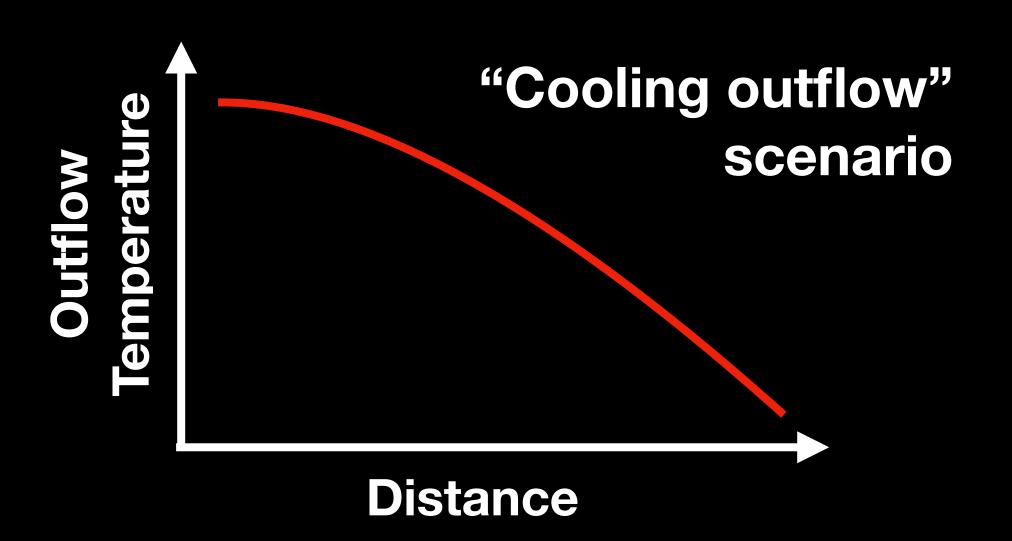


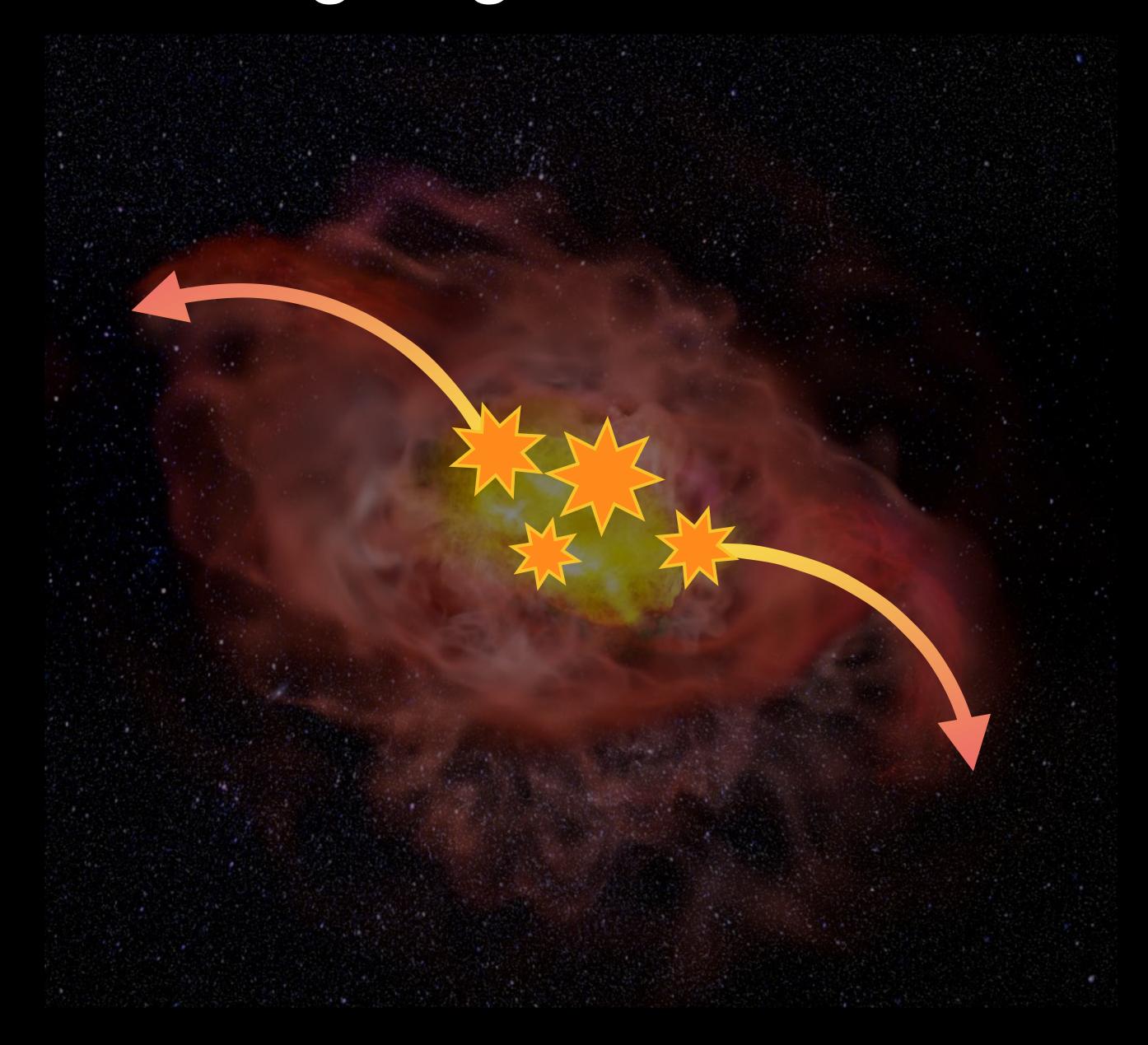


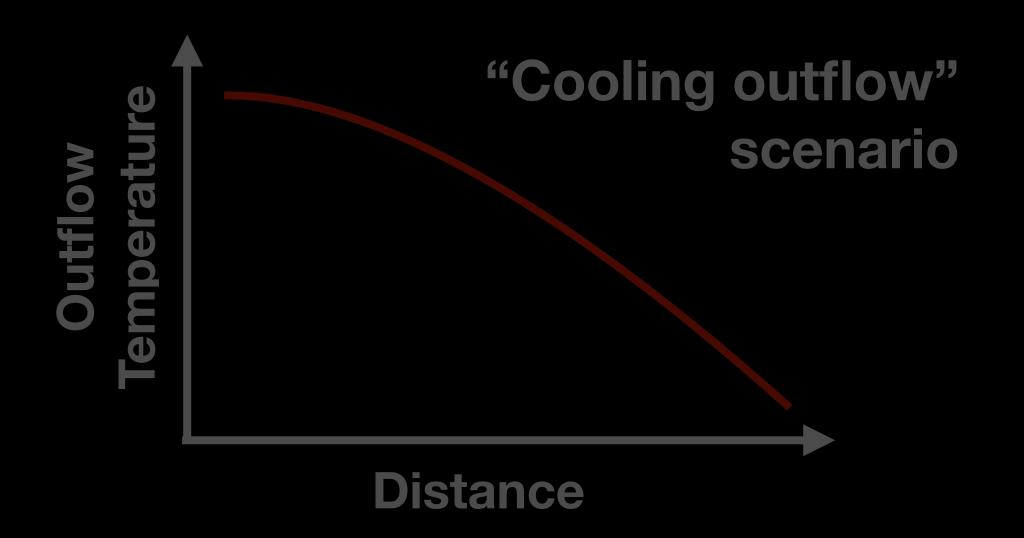




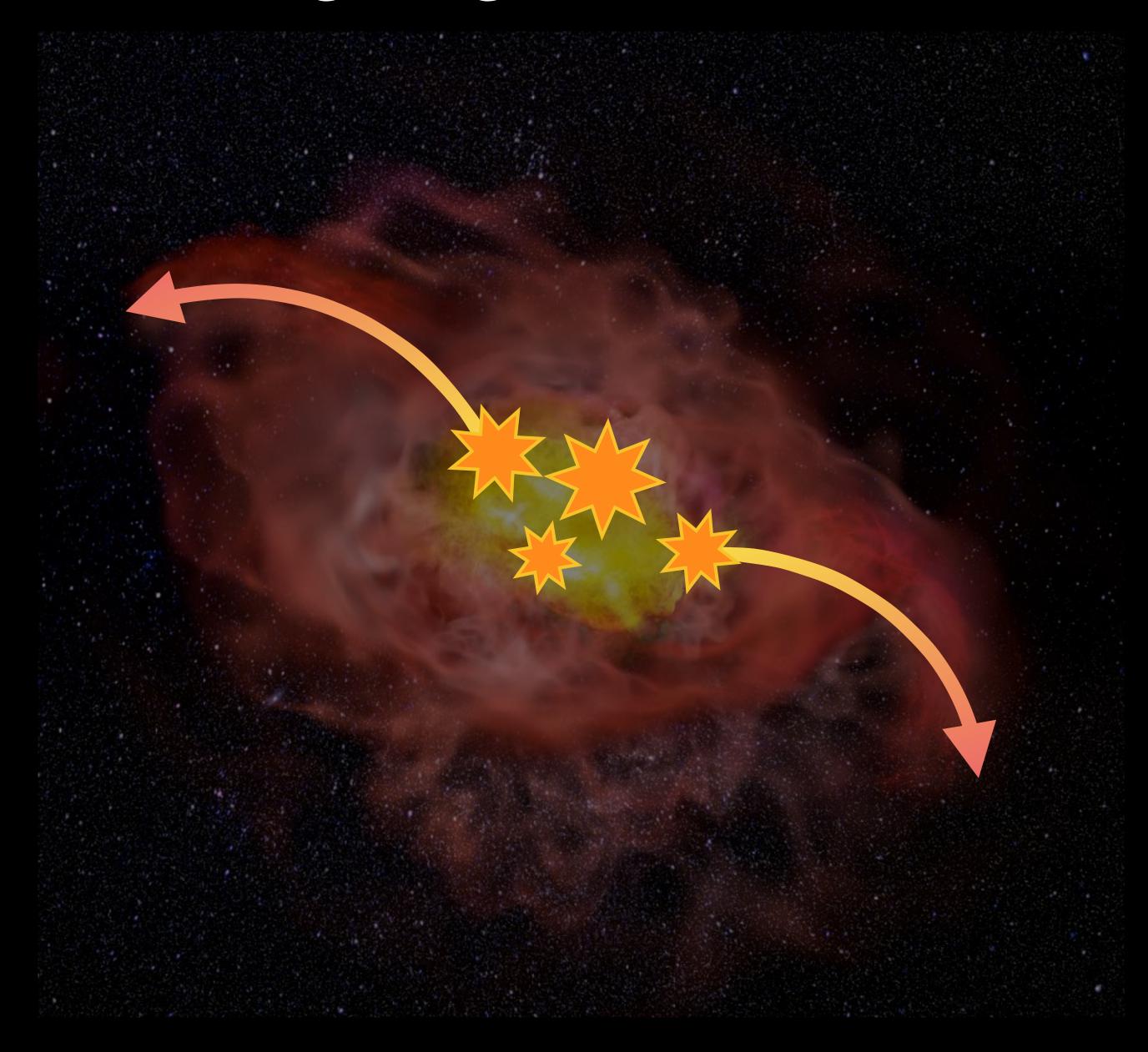


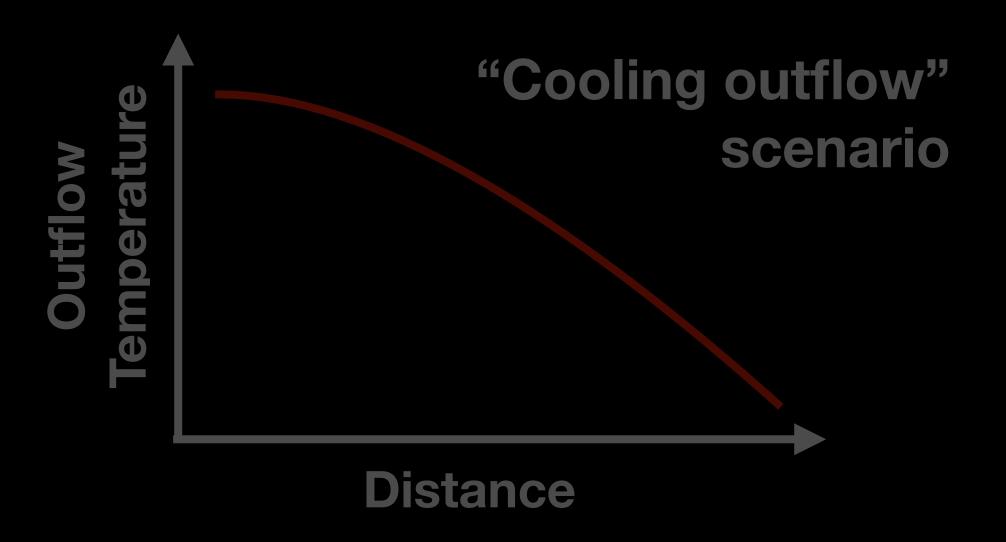






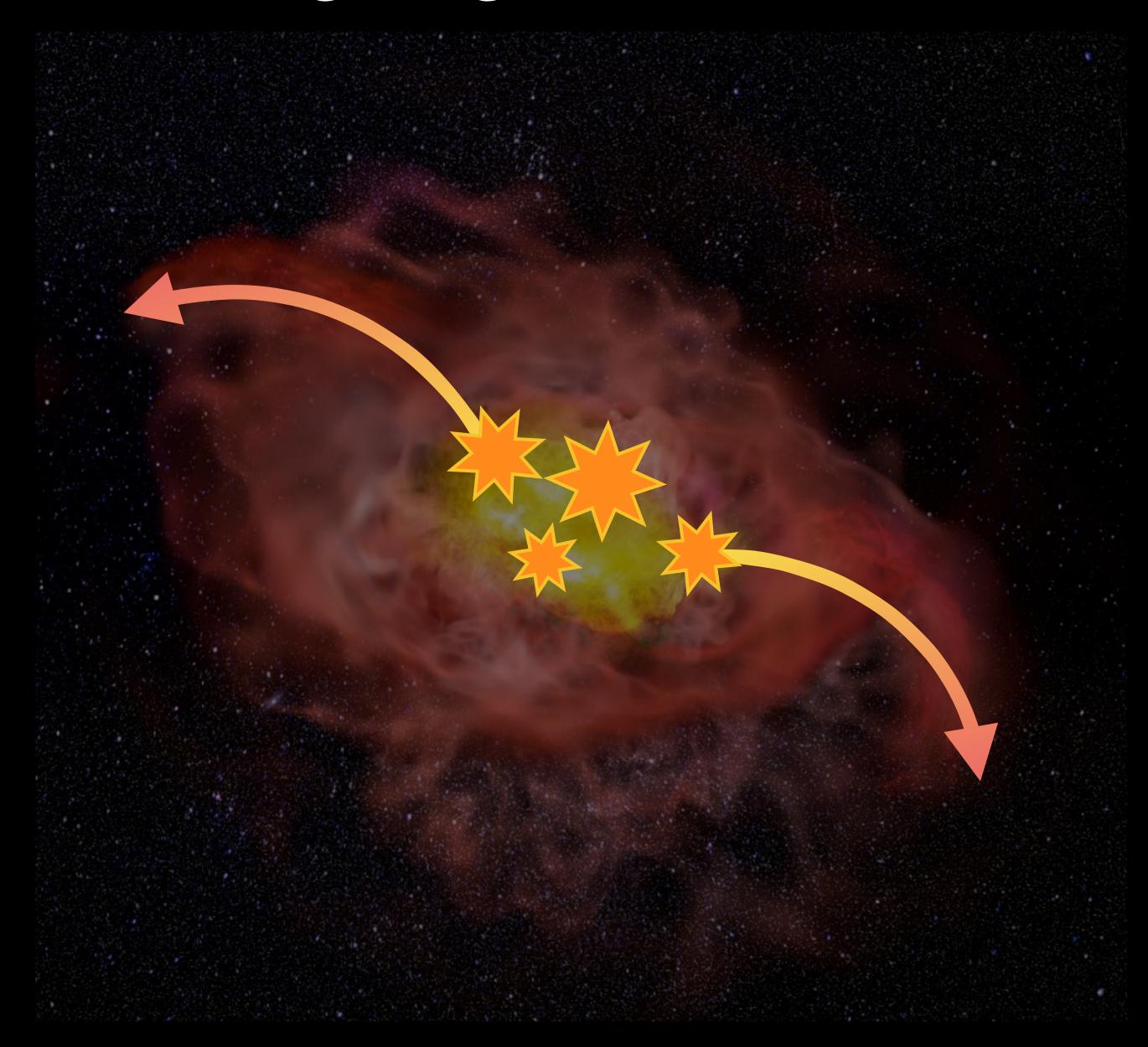
Key Takeaways

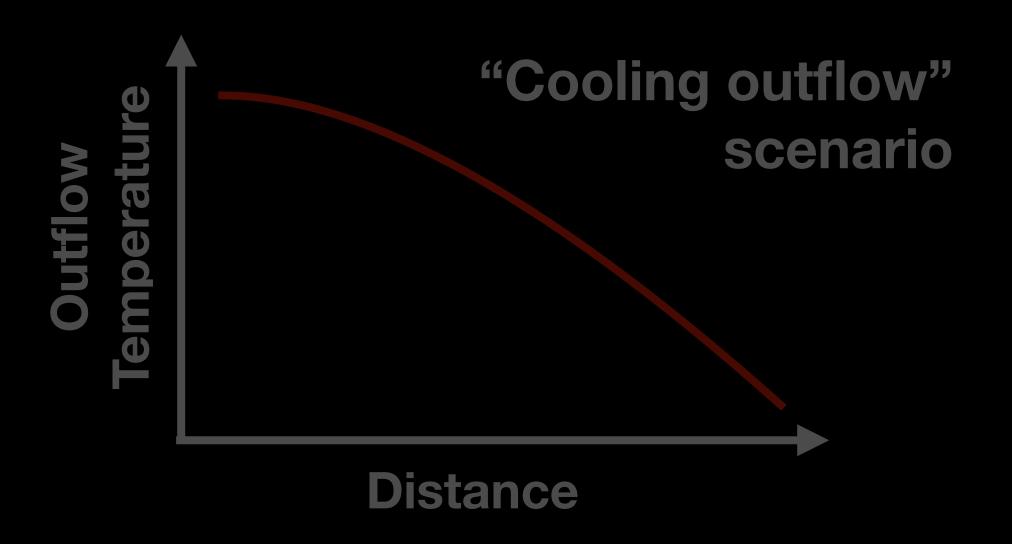




Key Takeaways

 Cooling outflows may be a critical part of the buildup of the CGM





Key Takeaways

- Cooling outflows may be a critical part of the buildup of the CGM
- Outflow activity can shape galaxy evolution even in the early universe and in relatively "normal" galaxies

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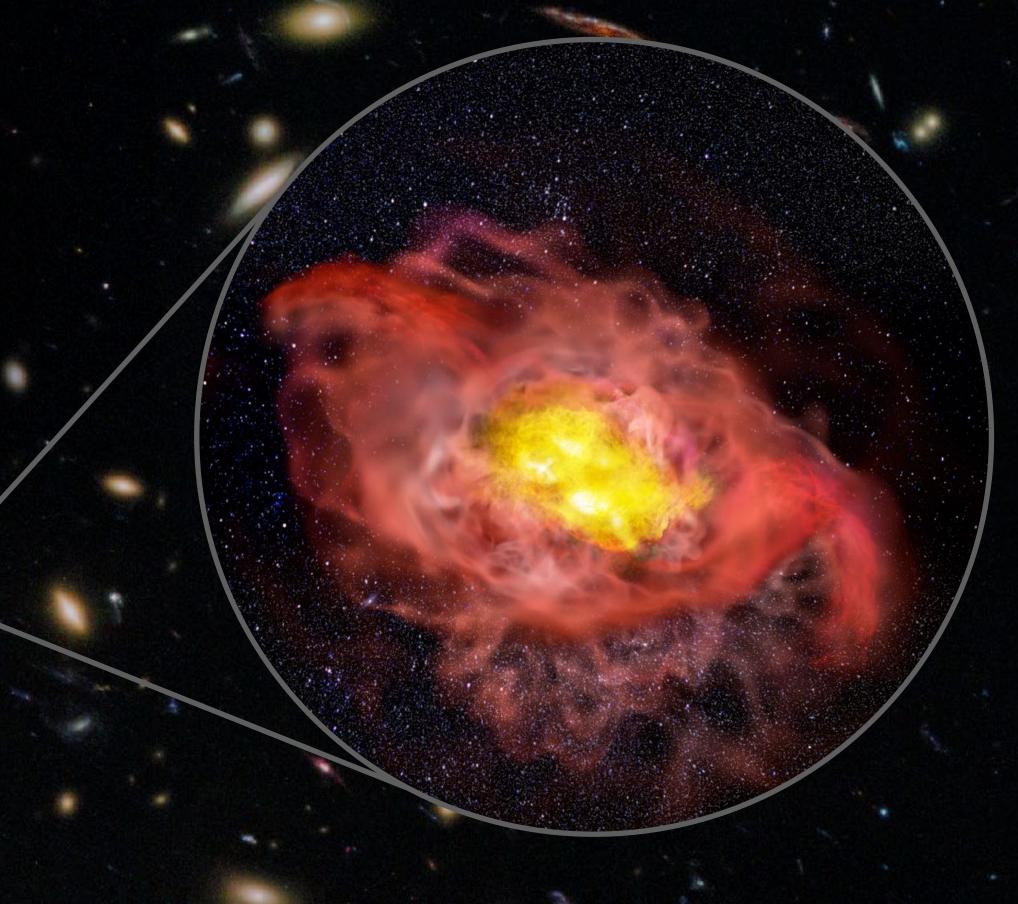
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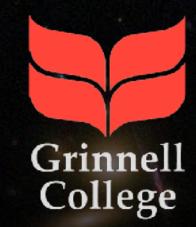
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