



A MULTIWAVELENGTH VIEW OF IC 860

WHAT IS IN ACTION INSIDE QUENCHING GALAXIES

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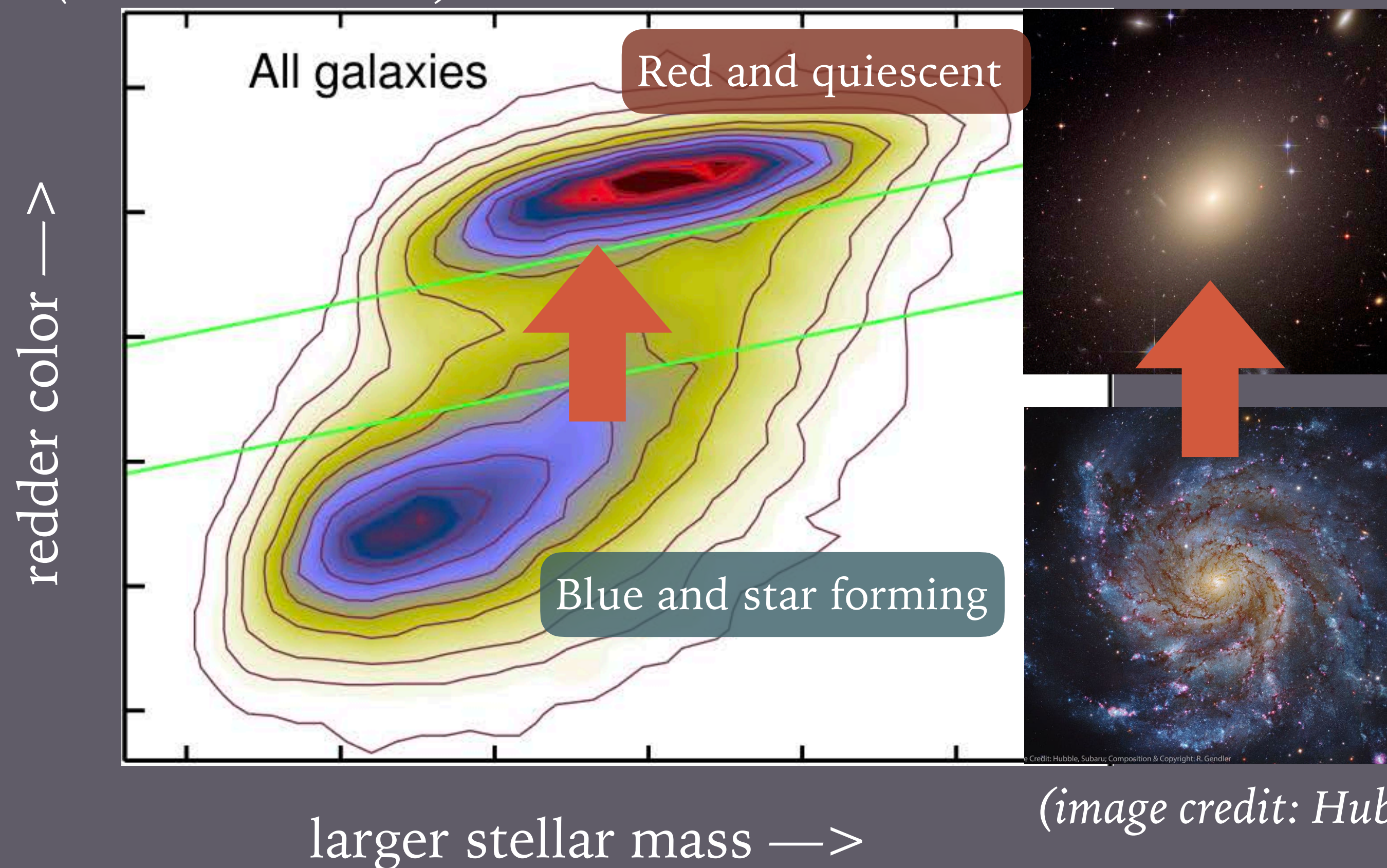
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Quenching (transitioning) Galaxies

(Schawinski+2014)



Quenching:

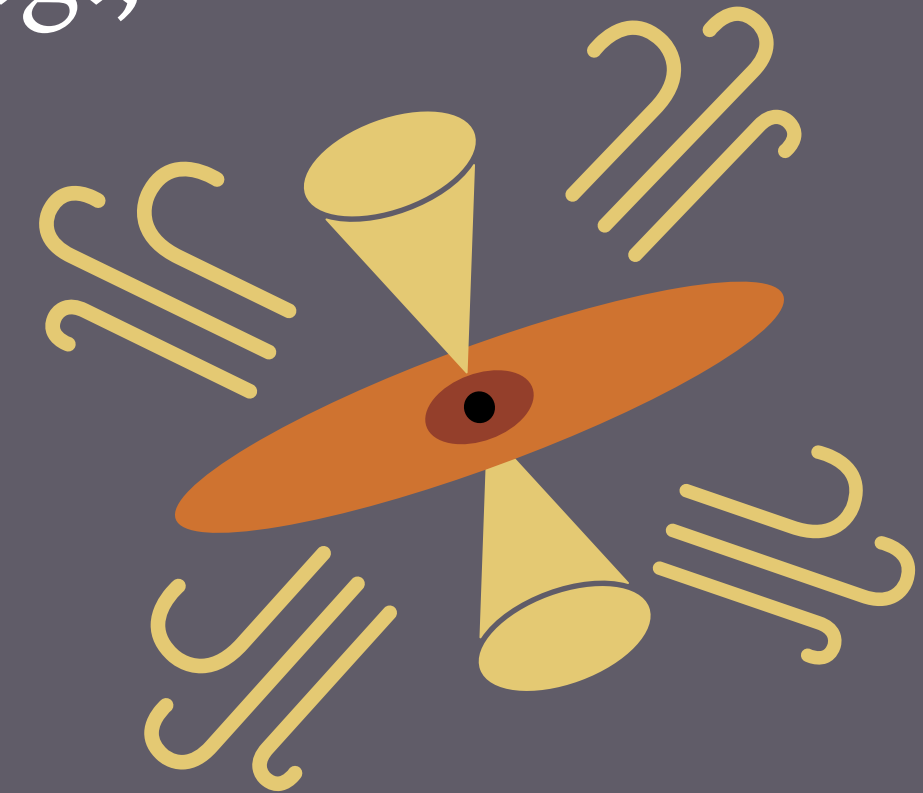
Decline in rate of star formation

“What’s stopping the star formation?”

Galaxies need to lose gas...?

No gas, no fuel for star formation

e.g.,



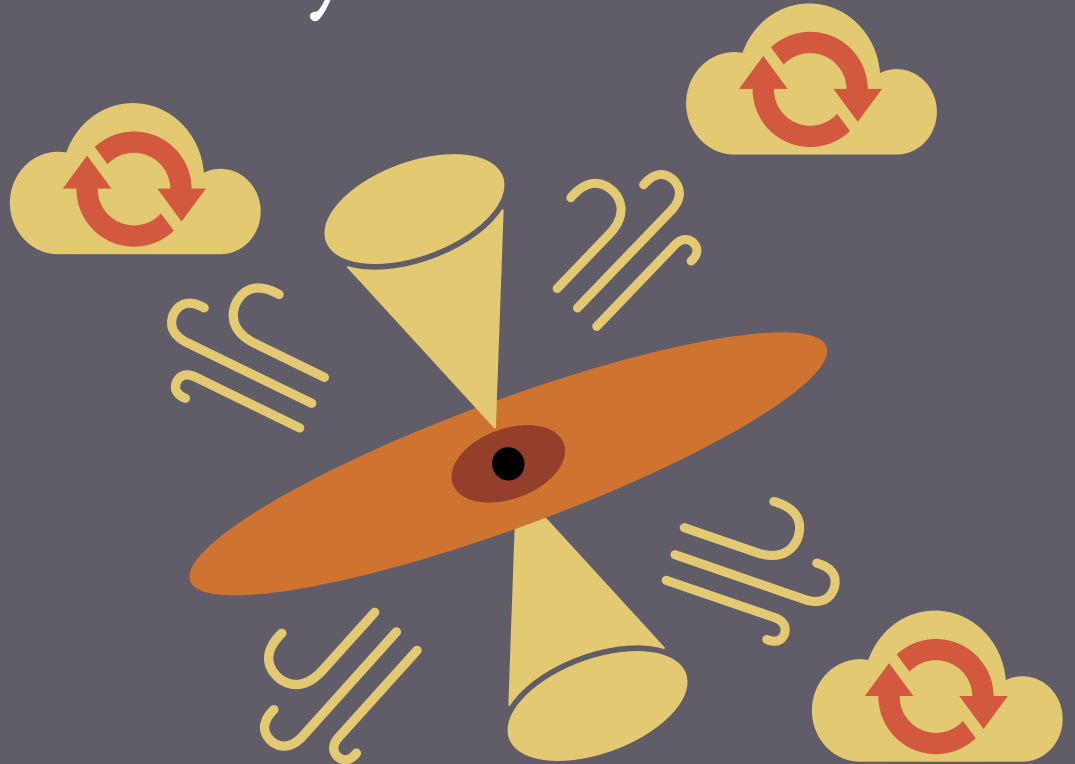
Black hole expels gas

BUT

Found large amounts of cold molecular gas in some quenching galaxies

(Rowlands+2015, French+2015)

Possibly:



Black hole disturbs gas

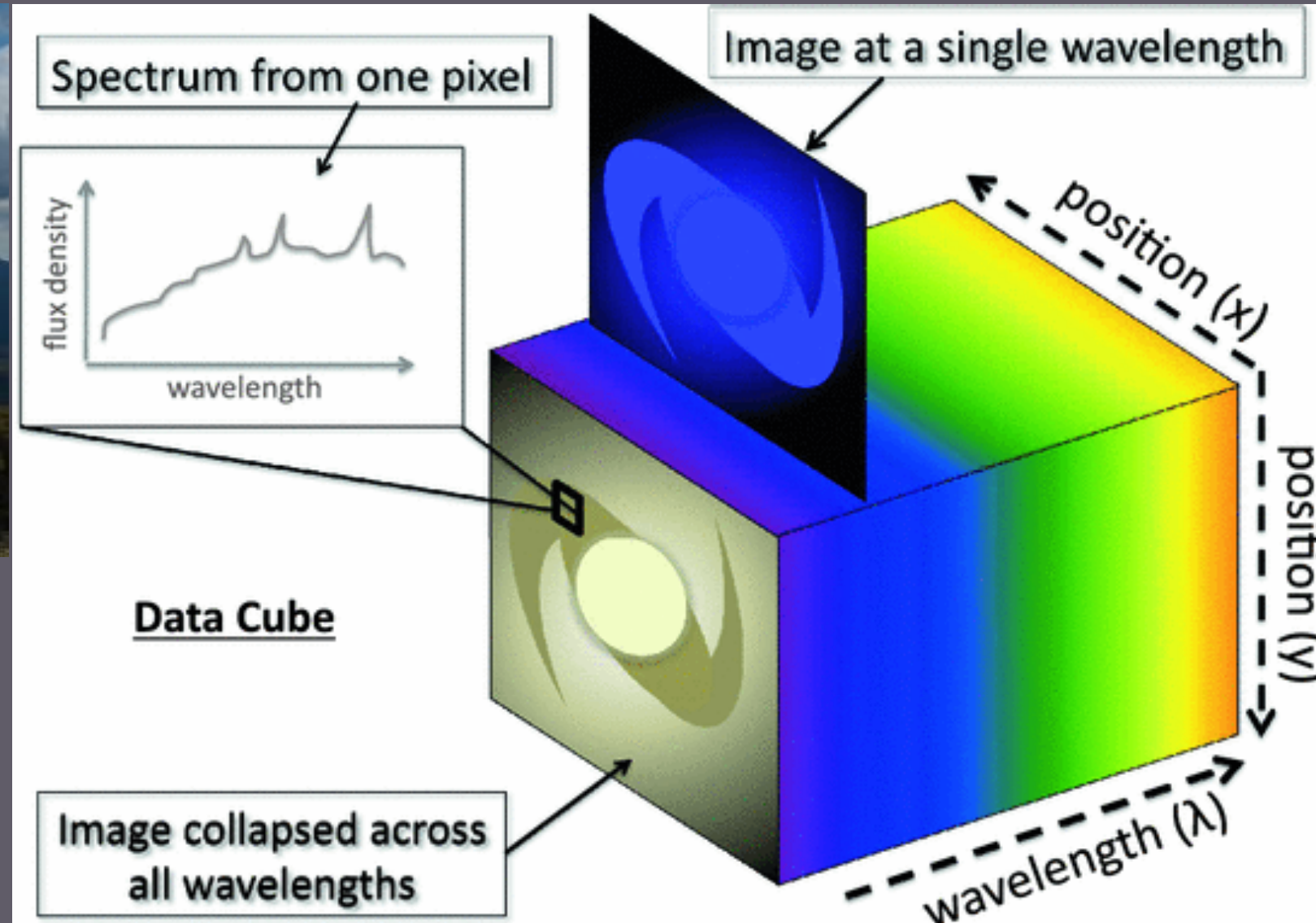
Let's look into one of them: IC 860



Multiwavelength data



Wide Field Spectrograph



CARMA at mm: spatially resolved kinematics of CO gas



Optical integral field spectroscopy: spatially resolved kinematics of stars and ionized gas



HST from NUV to NIR: exquisite multiband imaging

Dusty outflow + AGN + Recent merger

A 3rd spiral arm? Possible tidal feature after merger

r: F814W
g: F606W
b: F438W



0

Centrally concentrated molecular gas: external origin and possible recent merger

CO contours

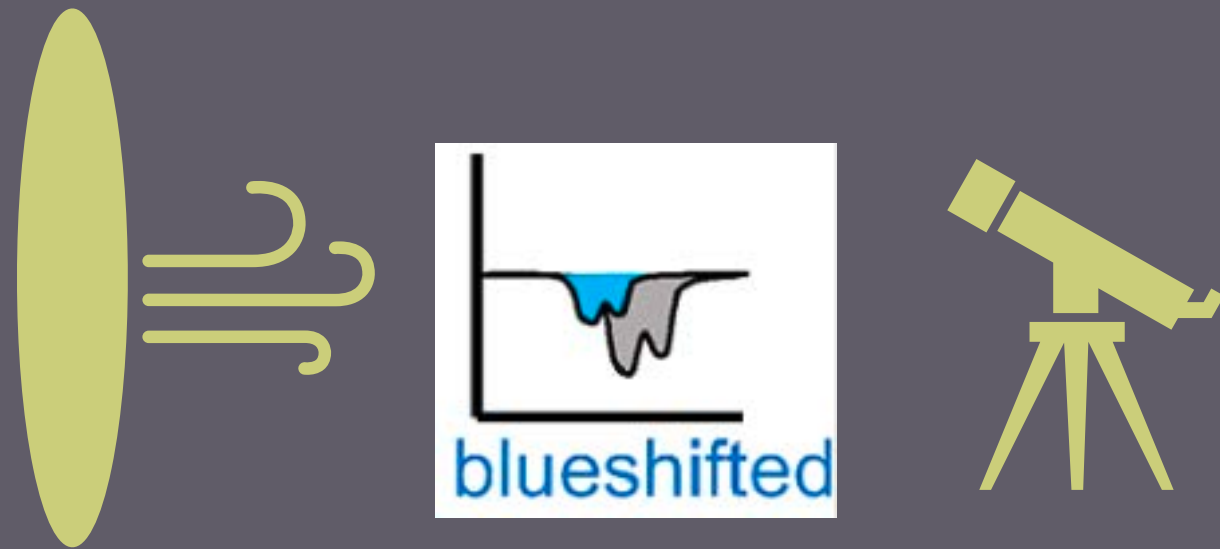
Blue optical color: early stage of transitioning

Compact radio emission at center >> evidence of AGN

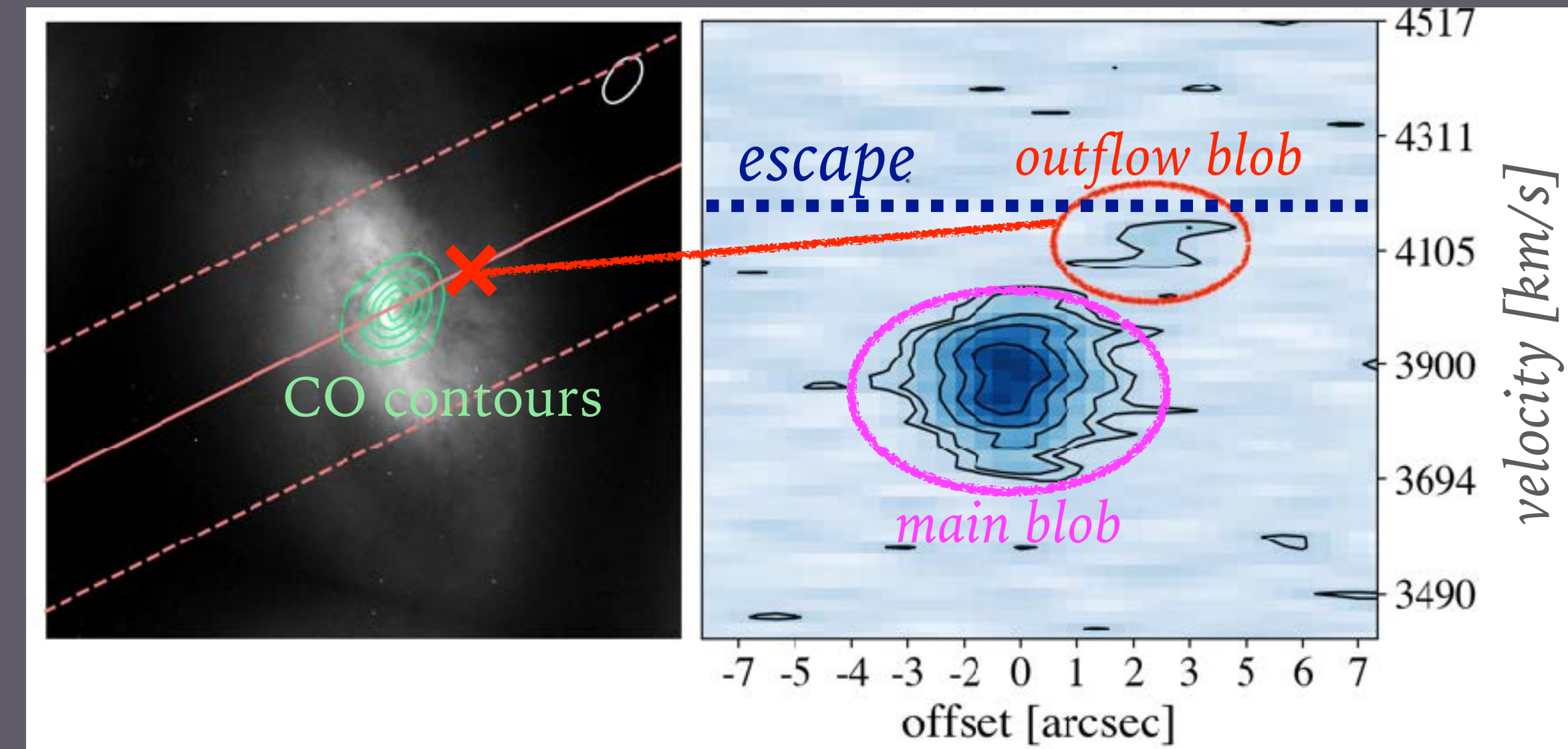
Dust tracing the cone of the outflow

Multiphase outflows

NaD doublet absorption \gg
neutral gas outflow

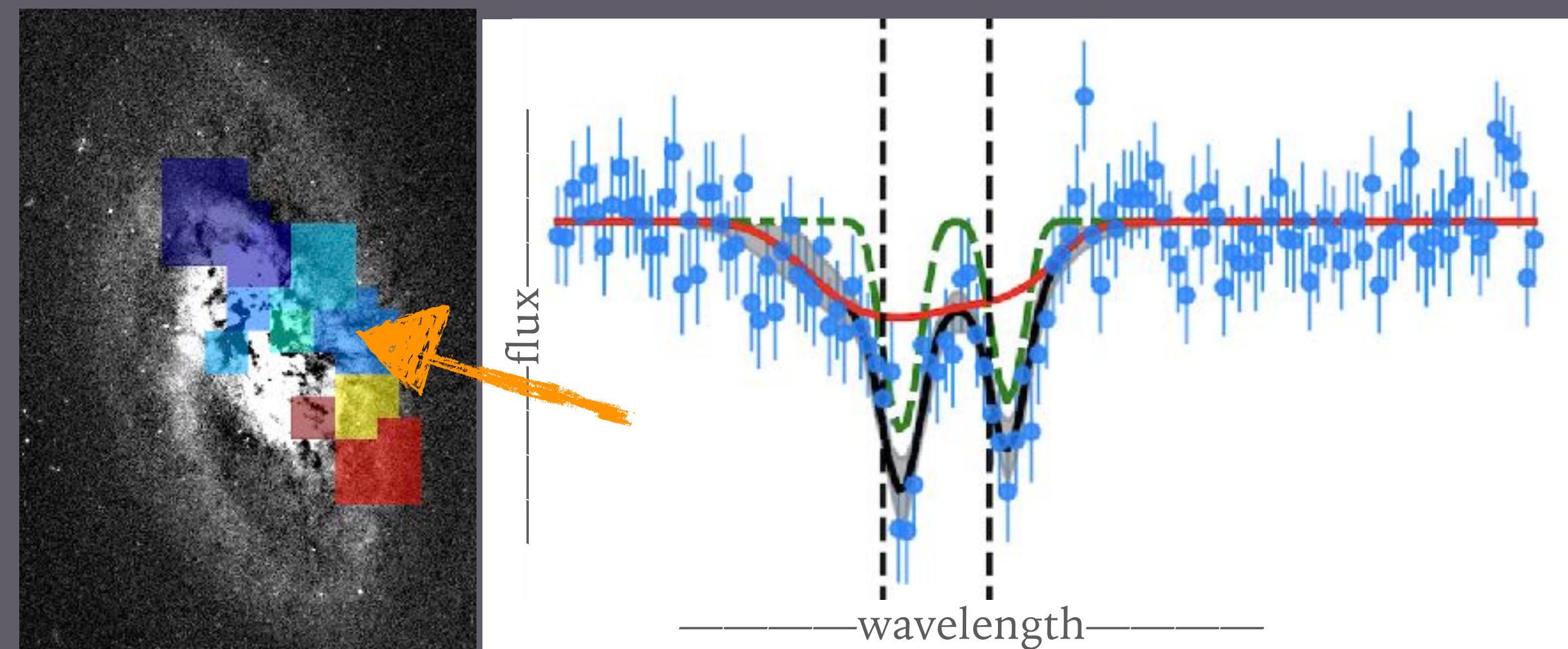


Relatively small mass outflow rate; cannot escape



molecular gas outflow \gg CO gas cloud deviating from bulk motion

Only 4% of total molecular gas mass; cannot escape



Star formation ends not with a bang but a whimper

(inspired by Eliot's poem)

RESULTS

- ▶ Evidence of merger history and AGN activity
- ▶ Presence of multiphase outflows (molecular + neutral)
- ▶ The outflows are not powerful enough to expel gas out of the galaxy

IMPLICATIONS

- ▶ Outflows disturb the gas to prevent star formation, not expel the gas
- ▶ AGN plays a role in galaxy quenching, but more investigation is desirable

(Luo + in prep)